

Matching NLO QCD with Parton Showers

Bryan Webber

TeV4LHC, 16-18 Sept 2004

- Introduction
 - ❖ Motivation & objectives
- Toy Model
 - ❖ Toy Monte Carlo
 - ❖ Modified subtraction
- Real QCD
 - ❖ Gauge boson pair production
 - ❖ Heavy quark production
 - ❖ Higgs boson production
- S Frixione & BRW, JHEP 0206(2002)029 [hep-ph/0204244]; hep-ph/0309186
- S Frixione, P Nason & BRW, JHEP 0308(2003)007 [hep-ph/0305252]
- V Del Duca, S Frixione, C Oleari & BRW, in preparation

<http://www.hep.phy.cam.ac.uk/theory/webber/MCatNLO/>

Motivation

- Reliable prediction of cross sections and final-state distributions for QCD processes is important not only as a test of QCD but also for the design of collider experiments and new particle searches.
- All systematic approaches to this problem are based on perturbation theory, usually truncated at next-to-leading order (NLO).
- For the description of exclusive hadronic final states, perturbative calculations have to be combined with a model for the conversion of partonic final states into hadrons (hadronization). Existing hadronization models are in remarkably good agreement with a wide range of data, after tuning of model parameters.
- However, these models operate on partonic states with high multiplicity and low relative transverse momenta, which are obtained from a parton shower Monte Carlo (MC) approximation to QCD dynamics and not from fixed-order calculations.

Objectives

- Our aim is to develop a **practical** method for combining **existing** parton shower MC programs with NLO perturbative calculations (**MC@NLO**).
- We require MC@NLO to have the following characteristics:
 - ❖ The output is a set of events, which are fully exclusive.
 - ❖ Total rates are accurate to NLO.
 - ❖ NLO results for all observables are recovered upon expansion of MC@NLO results in α_s .
 - ❖ Hard emissions are treated as in NLO computations.
 - ❖ Soft/collinear emissions are treated as in MC.
 - ❖ The matching between hard- and soft-emission regions is smooth.
 - ❖ MC hadronization models are adopted.

Toy Model

- Consider first a toy model that allows simple discussion of key features of NLO, of MC, and of matching between the two.
 - ❖ Assume a system can radiate massless “photons”, energy x , with $0 \leq x \leq 1$, x_s being energy of system before radiation.
 - ❖ After radiation, energy of system is $x'_s = x_s - x$.
 - ❖ System can undergo further emissions, but photons themselves cannot radiate.
- Task of predicting an infrared-safe observable O to NLO amounts to computing the quantity

$$\langle O \rangle = \lim_{\epsilon \rightarrow 0} \int_0^1 dx x^{-2\epsilon} O(x) \left[\left(\frac{d\sigma}{dx} \right)_B + \left(\frac{d\sigma}{dx} \right)_V + \left(\frac{d\sigma}{dx} \right)_R \right]$$

where **Born**, **virtual** and **real** contributions are respectively

$$\left(\frac{d\sigma}{dx} \right)_{B,V,R} = B\delta(x), \quad a \left(\frac{B}{2\epsilon} + V \right) \delta(x), \quad a \frac{R(x)}{x},$$

a is coupling constant, and $\lim_{x \rightarrow 0} R(x) = B$.

- In **subtraction method**, real contribution is written as:

$$\langle O \rangle_{\text{R}} = aBO(0) \int_0^1 dx \frac{x^{-2\epsilon}}{x} + a \int_0^1 dx \frac{O(x)R(x) - BO(0)}{x^{1+2\epsilon}}.$$

Second integral is non-singular, so we can set $\epsilon = 0$:

$$\langle O \rangle_{\text{R}} = -a \frac{B}{2\epsilon} O(0) + a \int_0^1 dx \frac{O(x)R(x) - BO(0)}{x}$$

- Therefore NLO prediction is:

$$\langle O \rangle_{\text{sub}} = BO(0) + a \left[VO(0) + \int_0^1 dx \frac{O(x)R(x) - BO(0)}{x} \right]$$

- We rewrite this in a slightly different form:

$$\langle O \rangle_{\text{sub}} = \int_0^1 dx \left[O(x) \frac{aR(x)}{x} + O(0) \left(B + aV - \frac{aB}{x} \right) \right]$$

Toy Monte Carlo

- In a treatment based on Monte Carlo methods, the system can undergo an arbitrary number of emissions (branchings), with probability controlled by the **Sudakov form factor**, defined for our toy model as follows:

$$\Delta(x_1, x_2) = \exp \left[-a \int_{x_1}^{x_2} dz \frac{Q(x)}{x} \right]$$

where $Q(x)$ is a monotonic function with the following properties:

$$0 \leq Q(x) \leq 1, \quad \lim_{x \rightarrow 0} Q(x) = 1, \quad \lim_{x \rightarrow 1} Q(x) = 0$$

$\Delta(x_1, x_2)$ is the probability that no photon be emitted with energy x such that $x_1 \leq x \leq x_2$.

Modified Subtraction

- We want to interface NLO to MC. Naive first try:

$$O(0) \Rightarrow \text{start MC with 0 real emissions: } \mathcal{F}_{\text{MC}}^{(0)}$$

$$O(x) \Rightarrow \text{start MC with 1 emission at } x: \mathcal{F}_{\text{MC}}^{(1)}(x)$$

so that overall **generating functional** is

$$\int_0^1 dx \left[\mathcal{F}_{\text{MC}}^{(0)} \left(B + aV - \frac{aB}{x} \right) + \mathcal{F}_{\text{MC}}^{(1)}(x) \frac{aR(x)}{x} \right]$$

- This is **wrong**: MC starting with no emissions will generate emission, with NLO distribution

$$\left(\frac{d\sigma}{dx} \right)_{\text{MC}} = aB \frac{Q(x)}{x}$$

We must subtract this from second term, and add to first:

$$\begin{aligned} \mathcal{F}_{\text{MC@NLO}} &= \int_0^1 dx \left[\mathcal{F}_{\text{MC}}^{(0)} \left(B + aV + \frac{aB[Q(x) - 1]}{x} \right) \right. \\ &\quad \left. + \mathcal{F}_{\text{MC}}^{(1)}(x) \frac{a[R(x) - BQ(x)]}{x} \right] \end{aligned}$$

$$\mathcal{F}_{\text{MC@NLO}} = \int_0^1 dx \left[\mathcal{F}_{\text{MC}}^{(0)} \left(B + aV + \frac{aB[Q(x) - 1]}{x} \right) + \mathcal{F}_{\text{MC}}^{(1)}(x) \frac{a[R(x) - BQ(x)]}{x} \right]$$

This prescription has several good features:

- $\mathcal{F}_{\text{MC}}^{(0)} = \mathcal{F}_{\text{MC}}^{(1)}$ to $\mathcal{O}(1)$, so added and subtracted terms are equal to $\mathcal{O}(a)$;
- Coefficients of $\mathcal{F}_{\text{MC}}^{(0)}$ and $\mathcal{F}_{\text{MC}}^{(1)}$ are now separately finite;
- Same resummation of large logs in $\mathcal{F}_{\text{MC}}^{(0)}$ and $\mathcal{F}_{\text{MC}}^{(1)} \Rightarrow \mathcal{F}_{\text{MC@NLO}}$ gives same resummation as $\mathcal{F}_{\text{MC}}^{(0)}$, renormalised to correct NLO cross section.

Note, however, that some events may have **negative weight**.

Toy Model Observables

- As an example of an “exclusive” observable, we consider the energy y of the hardest photon in each event. The NLO and MC predictions are

$$\left(\frac{d\sigma}{dy}\right)_{\text{NLO}} = a \frac{R(y)}{y}$$

$$\left(\frac{d\sigma}{dy}\right)_{\text{MC}} = aB \frac{Q(y)}{y} \Delta(y, 1)$$

- As an “inclusive” observable, consider the fully inclusive distribution of photon energies, z :

$$\left(\frac{d\sigma}{dz}\right)_{\text{NLO}} = a \frac{R(z)}{z}$$

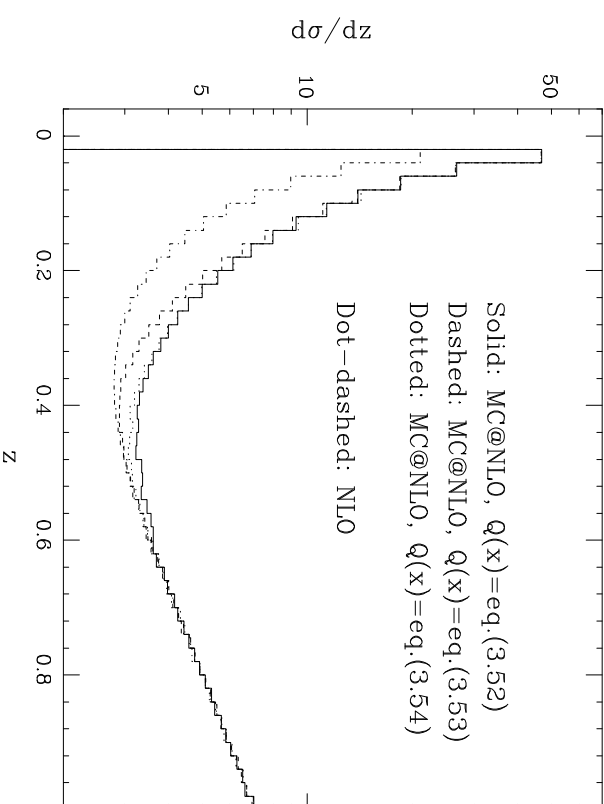
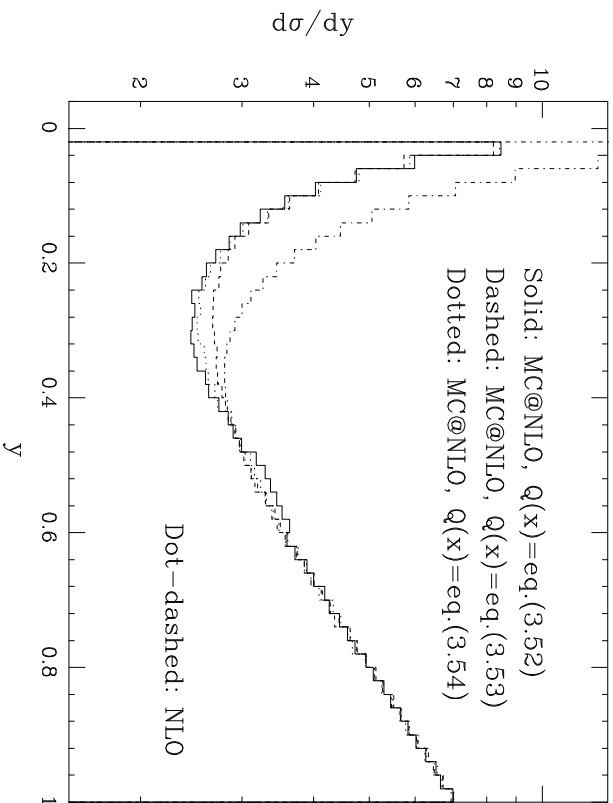
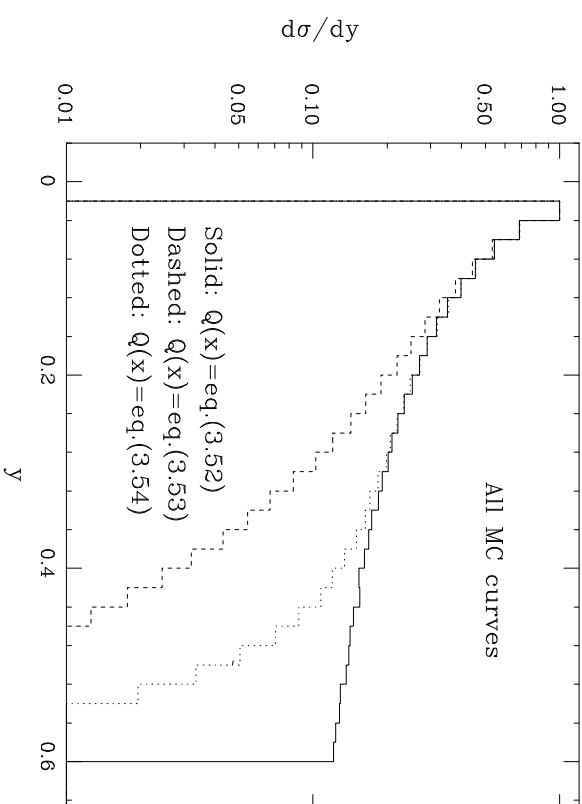
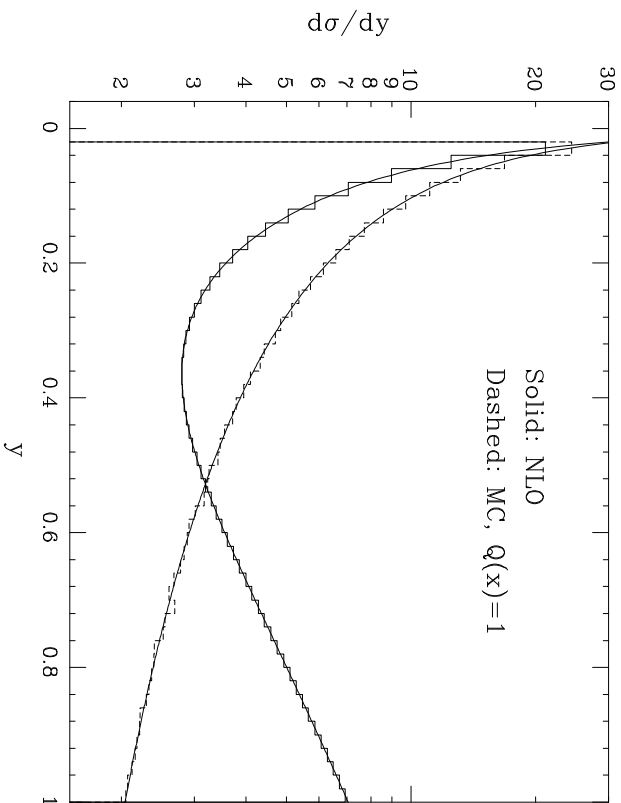
$$\left(\frac{d\sigma}{dz}\right)_{\text{MC}} = aB \frac{Q(z)}{z}$$

- Toy model results below are for

$$a = 0.3, \quad B = 2, \quad V = 1,$$

$$R(x) = B + x(1 + x/2 + 20x^2)$$

- For MC we have assumed a “dead zone” $Q(x) = 0$ for $x > 0.6$ with variable smoothing at boundary (see figure).



Modified Subtraction for Real QCD

- Consider a hadron collider process which is $2 \rightarrow 2$ at LO, e.g. W^+W^- or $Q\bar{Q}$ pair production. Schematic expression for any observable O , evaluated by subtraction method, is

$$\begin{aligned} \langle O \rangle_{\text{sub}} &= \sum_{ab} \int_0^1 dx_1 dx_2 d\phi_3 f_a(x_1) f_b(x_2) \left[O^{(2 \rightarrow 3)} \mathcal{M}_{ab}^{(h)}(x_1, x_2, \phi_3) \right. \\ &\quad \left. + O^{(2 \rightarrow 2)} \left(\mathcal{M}_{ab}^{(b,v,c)}(x_1, x_2, \phi_2) - \mathcal{M}_{ab}^{(c.t.)}(x_1, x_2, \phi_3) \right) \right] \end{aligned}$$

- ❖ $\mathcal{M}_{ab}^{(h)}$ is NLO real-emission contribution;
- ❖ $\mathcal{M}_{ab}^{(b,v,c)}$ are LO Born, NLO virtual and collinear (finite parts);
- ❖ $\mathcal{M}_{ab}^{(c.t.)}$ are counter-terms which cancel divergences of $\mathcal{M}_{ab}^{(h)}$.
- Naively, for MC@NLO we would replace $O^{(2 \rightarrow 2,3)}$ by $\mathcal{F}_{\text{MC}}^{(2 \rightarrow 2,3)}$ (MC generating functionals starting from $2 \rightarrow 2, 3$ hard subprocesses), to obtain $\mathcal{F}_{\text{MC@NLO}}$.
- This would be **wrong** because $\mathcal{F}_{\text{MC}}^{(2 \rightarrow 2)}$ also generates $2 \rightarrow 3$ configurations, which must be subtracted from weight of $\mathcal{F}_{\text{MC}}^{(2 \rightarrow 3)}$ (and added to that of

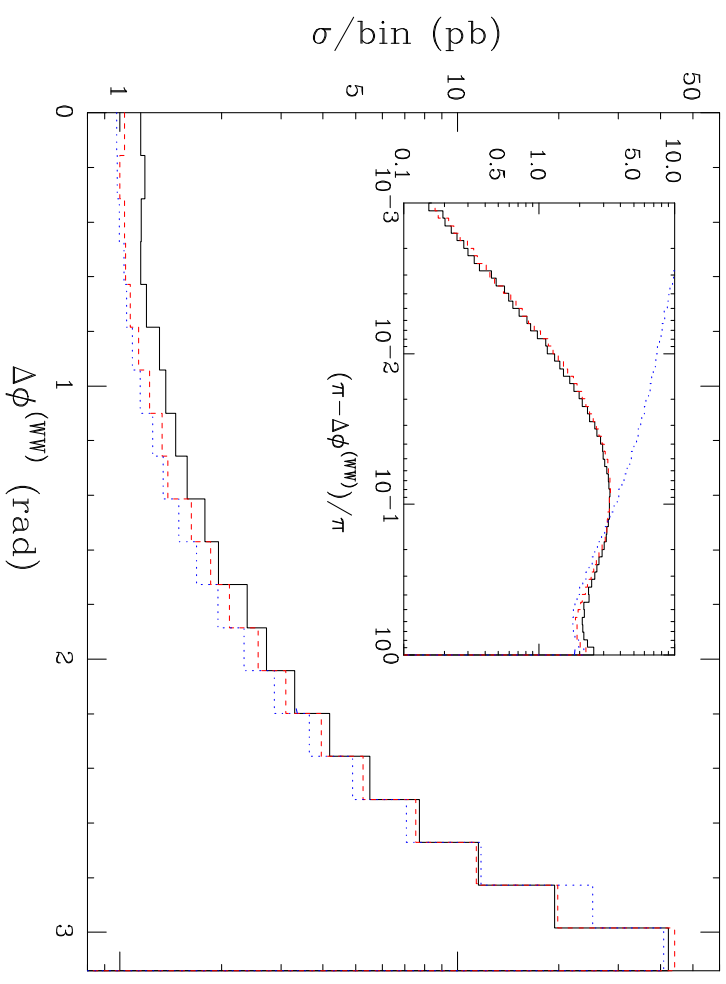
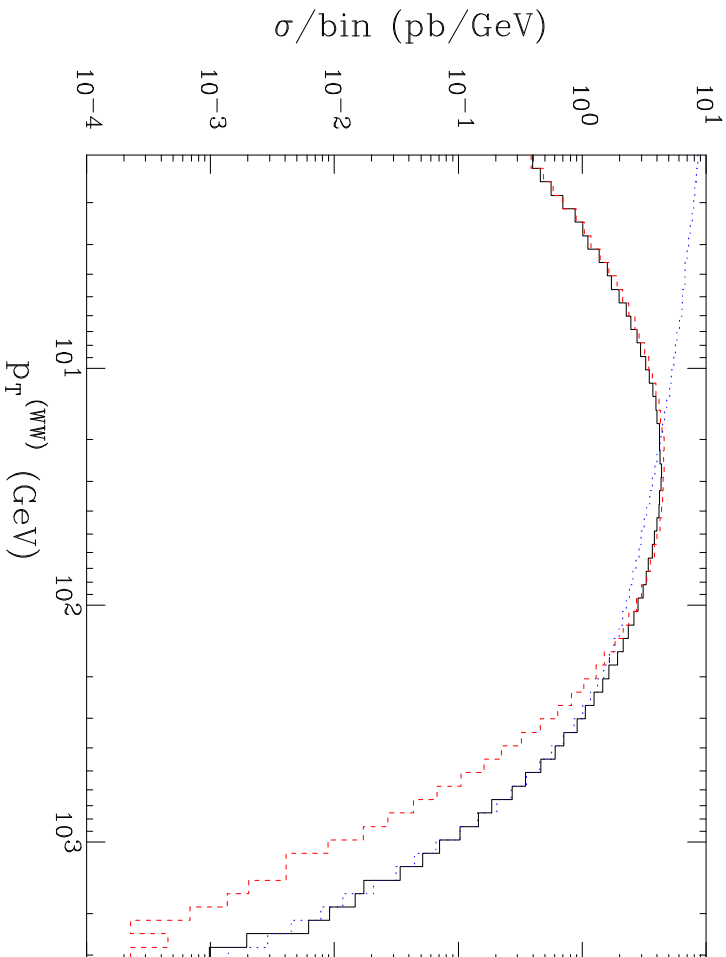
$$\mathcal{F}_{\text{MC}}^{(2 \rightarrow 2)}).$$

- Therefore for MC@NLO we define

$$\mathcal{F}_{\text{MC@NLO}} = \sum_{ab} \int_0^1 dx_1 dx_2 d\phi_3 f_a(x_1) f_b(x_2) \left[\mathcal{F}_{\text{MC}}^{(2 \rightarrow 3)} \left(\mathcal{M}_{ab}^{(h)}(x_1, x_2, \phi_3) - \mathcal{M}_{ab}^{(\text{MC})}(x_1, x_2, \phi_3) \right) + \mathcal{F}_{\text{MC}}^{(2 \rightarrow 2)} \left(\mathcal{M}_{ab}^{(b,v,c)}(x_1, x_2, \phi_2) - \mathcal{M}_{ab}^{(c.t.)}(x_1, x_2, \phi_3) + \mathcal{M}_{ab}^{(\text{MC})}(x_1, x_2, \phi_3) \right) \right]$$

- Provided MC does a good job in all soft and collinear limits, coefficients of $\mathcal{F}_{\text{MC}}^{(2 \rightarrow 2)}$ and $\mathcal{F}_{\text{MC}}^{(2 \rightarrow 3)}$ are now **separately finite**.
- But coefficients may be negative \Rightarrow some events have **negative weight**.
- Number of negative weights can be reduced by tuning counterterms. Typically we find 10 – 20%.

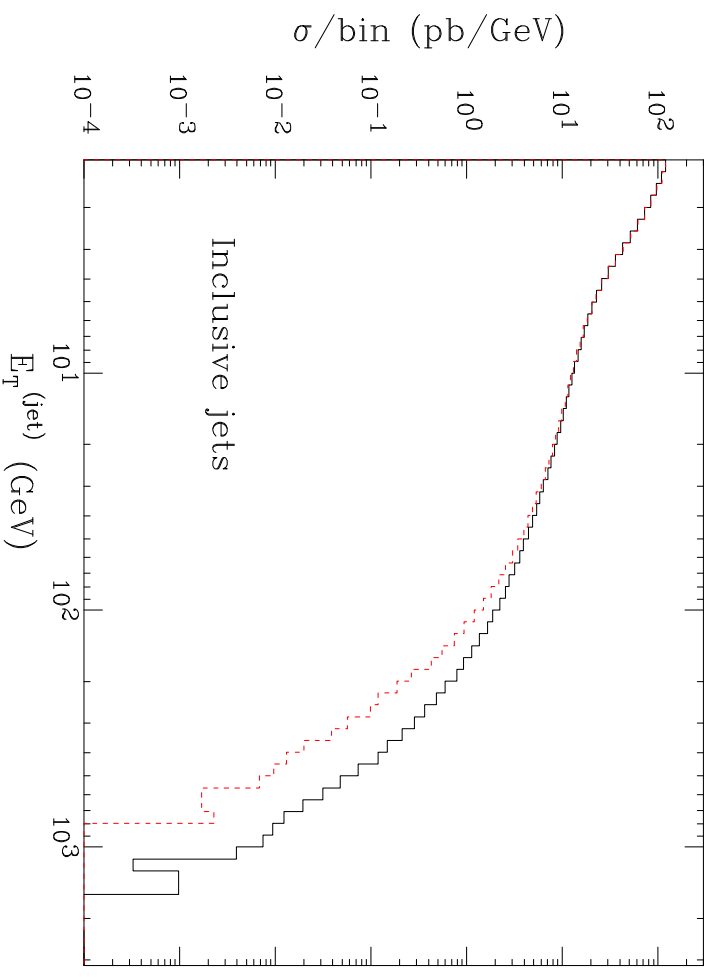
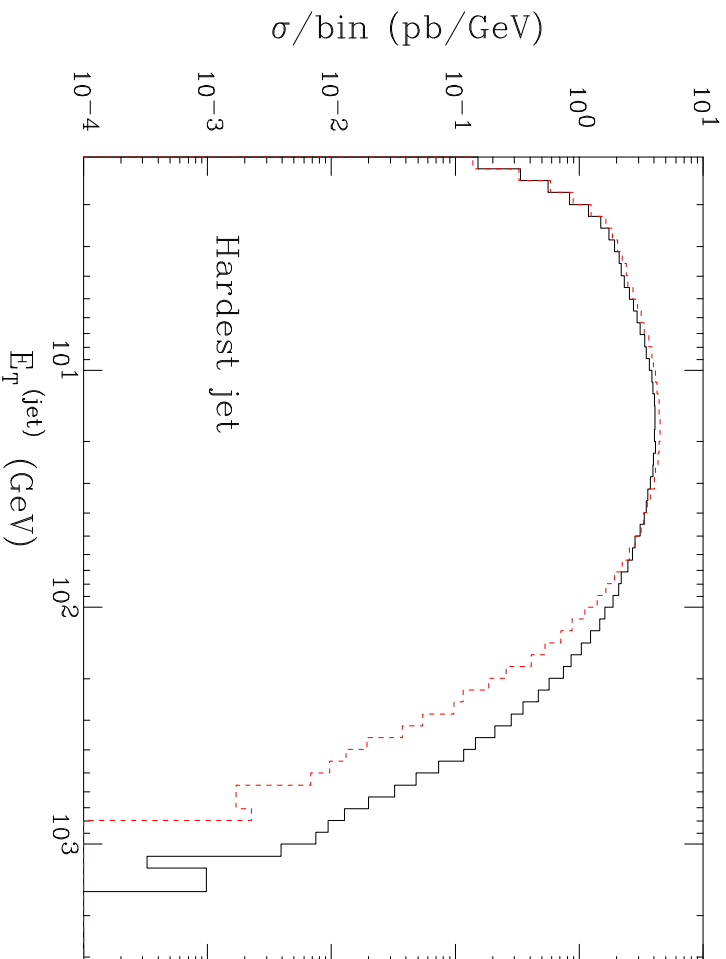
W^+W^- Observables



These correlations are problematic: the soft and hard emissions are both relevant. MC@NLO does well, resumming large logarithms, and yet handling the large-scale physics correctly

Solid: MC@NLO
Dashed: HERWIG $\times \frac{\sigma_{NLO}}{\sigma_{LO}}$
Dotted: NLO

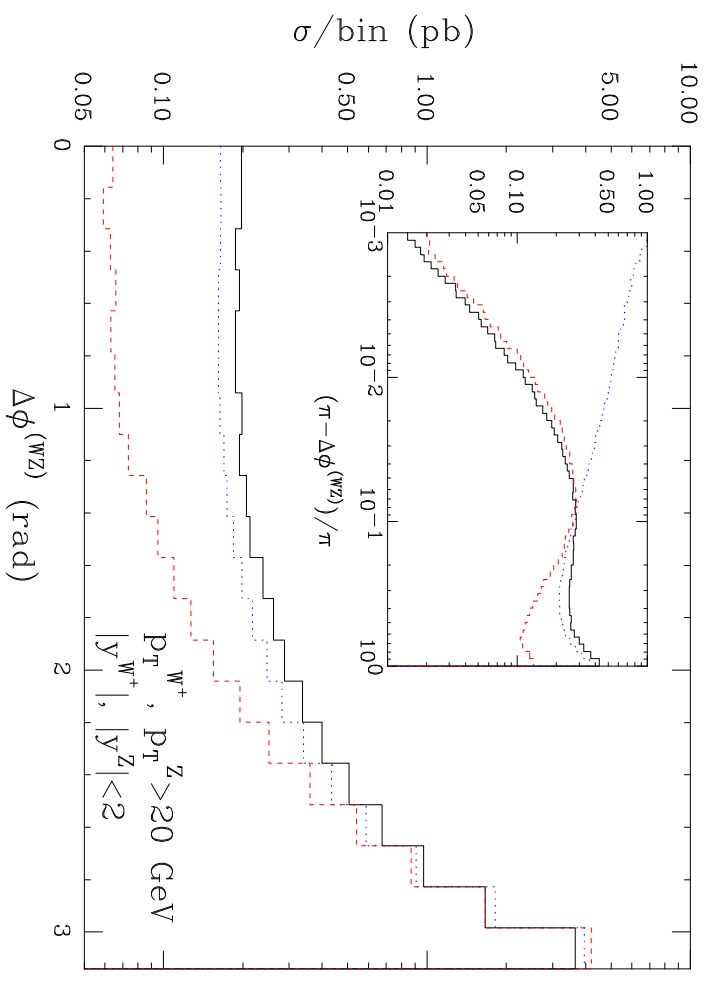
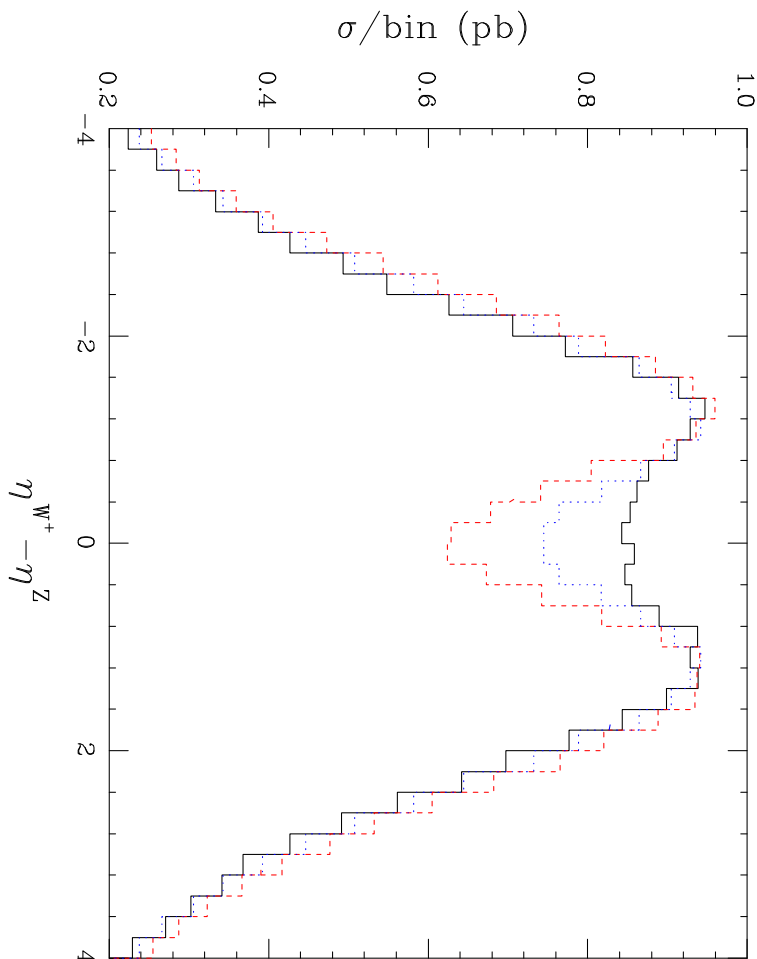
Jet Observables in W^+W^- Production



Jets have been reconstructed with a k_T algorithm. It is striking that inclusive jet distribution displays the same behaviour as in the toy model: MC@NLO/MC=K factor for $p_{\tau} \rightarrow 0$

Solid: MC@NLO
Dashed: HERWIG $\times \frac{\sigma_{NLO}}{\sigma_{LO}}$

W^+Z Observables



Solid: MC@NLO

Dashed: HERWIG $\times \frac{\sigma_{NLO}}{\sigma_{LO}}$

Dotted: NLO

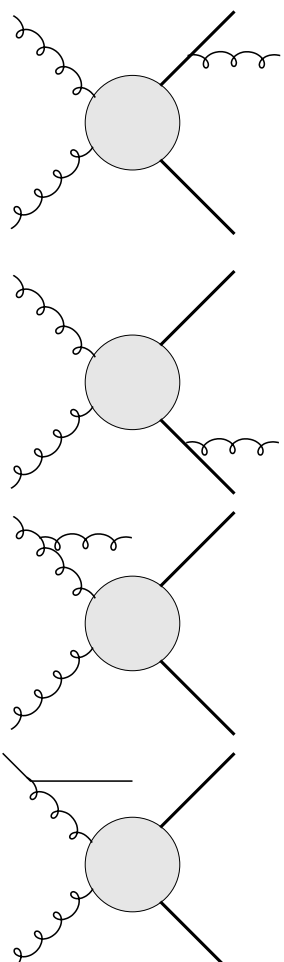
It is interesting that the MC@NLO fills further the kinematic dip at $\eta_{W^+} - \eta_Z = 0$. The difference between MC@NLO and MC is enhanced by the cuts in the $\Delta\phi$ tail

Heavy Quark Production

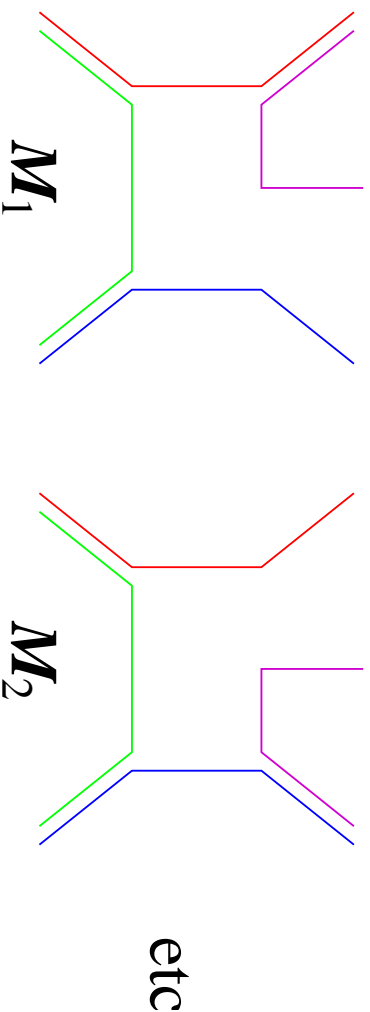
- Modified subtraction formula above can be used for any process.
 - ❖ Take standard subtraction formula;
 - ❖ Calculate analytically **exactly** what MC does at NLO;
 - ❖ Insert $\mathcal{M}_{ab}^{(\text{MC})}(x_1, x_2, \phi_3)$ terms;
 - ❖ Generate $2 \rightarrow 2$ and $2 \rightarrow 3$ parton configurations and weights;
 - ❖ Feed into MC (using Les Houches interface, hep-ph/0109068).
- Most difficult part is calculating what MC does!
 - ❖ Details in FNW, JHEP 0308(2003)007 [hep-ph/0305252]

MC Heavy Quark Production

- MC starts from $2 \rightarrow 2$ subprocess \Rightarrow momentum reshuffling is done after real emission.

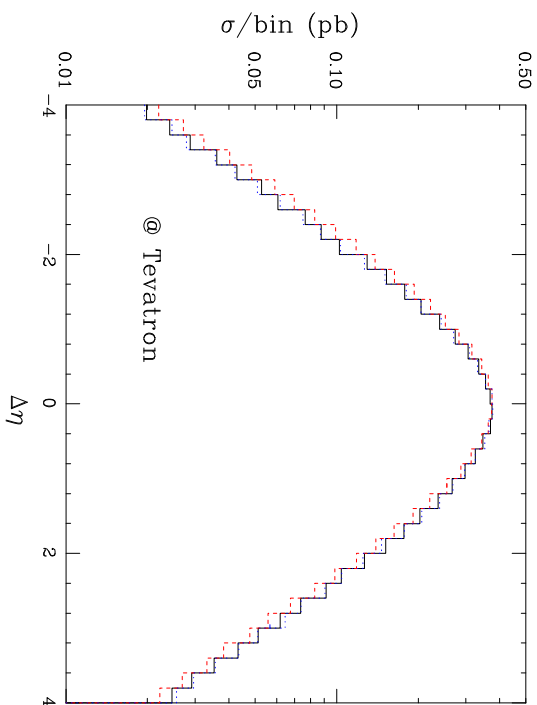
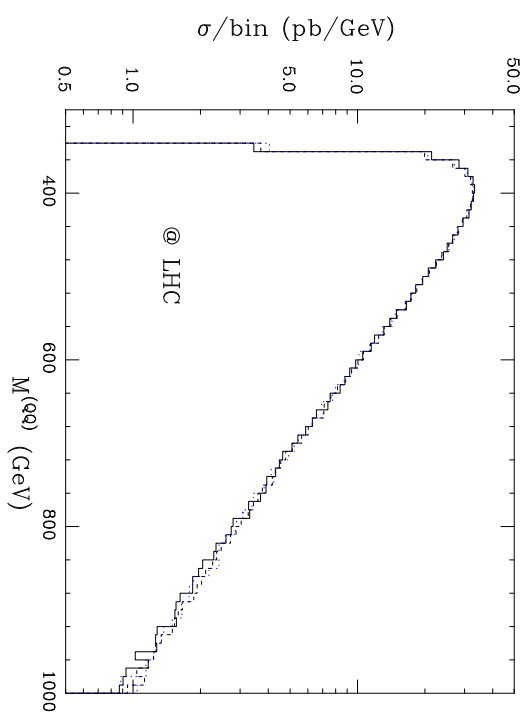
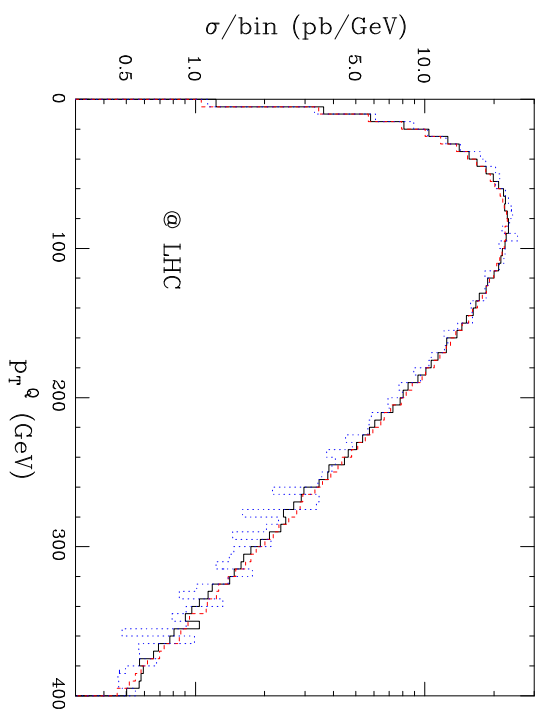


- Relation between invariants and shower variables depends on which leg emits!
- Colour structure assigned (for shower/hadronization) according to $N \rightarrow \infty$ limit.



$$\text{Prob}_i = \frac{|\sum_j M_j^{(3)}|^2 |M_i^{(\infty)}|^2}{\sum_j |M_j^{(\infty)}|^2}$$

t, \bar{t} Observables at Colliders



Solid: MC@NLO

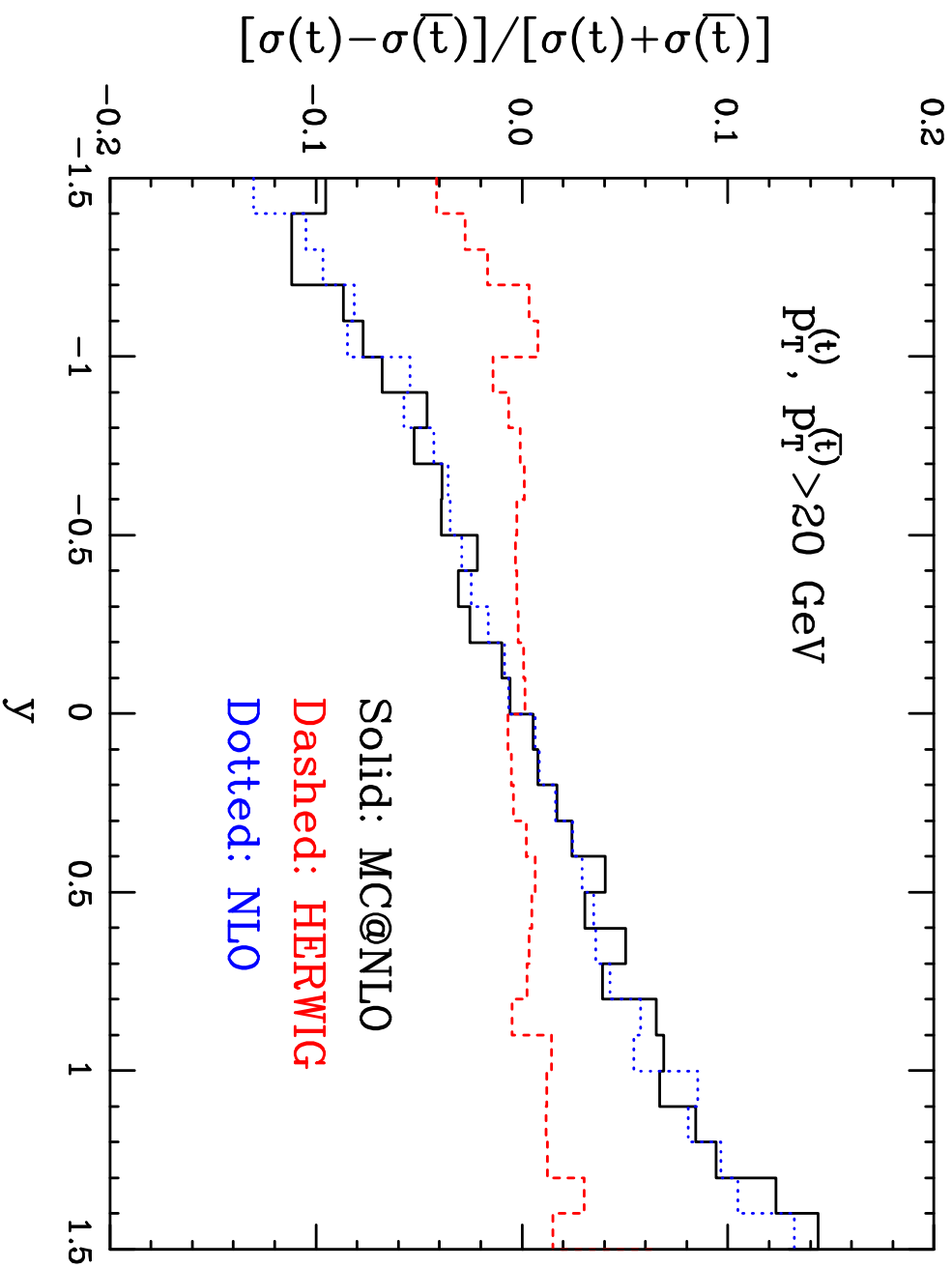
Dashed: HERWIG $\times \frac{\sigma_{NLO}}{\sigma_{LO}}$

Dotted: NLO

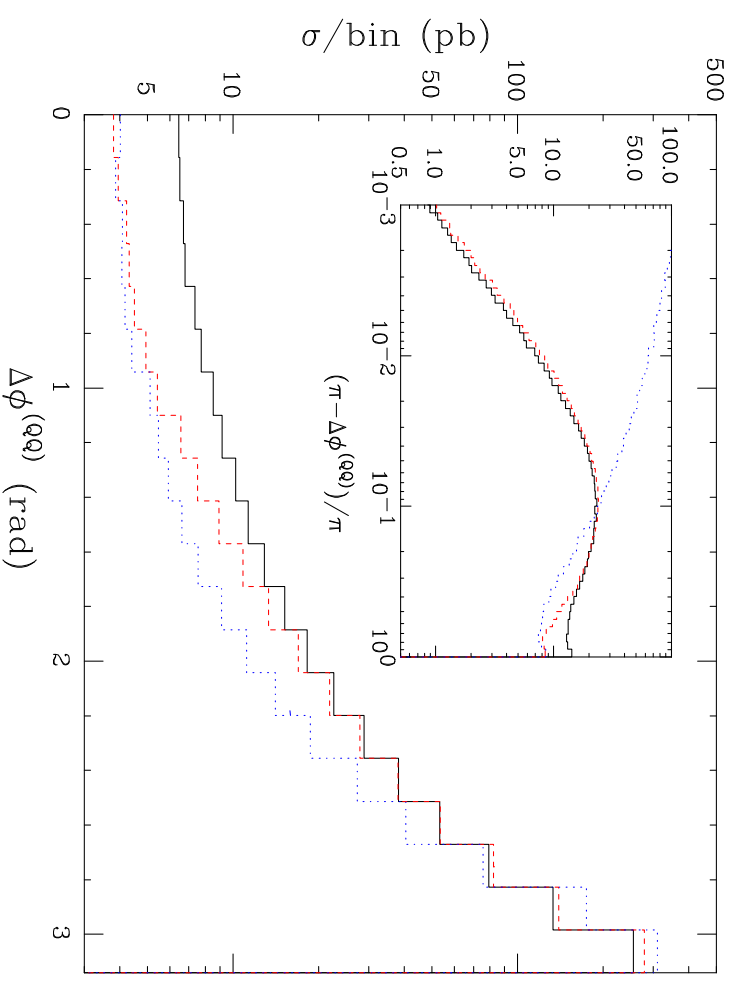
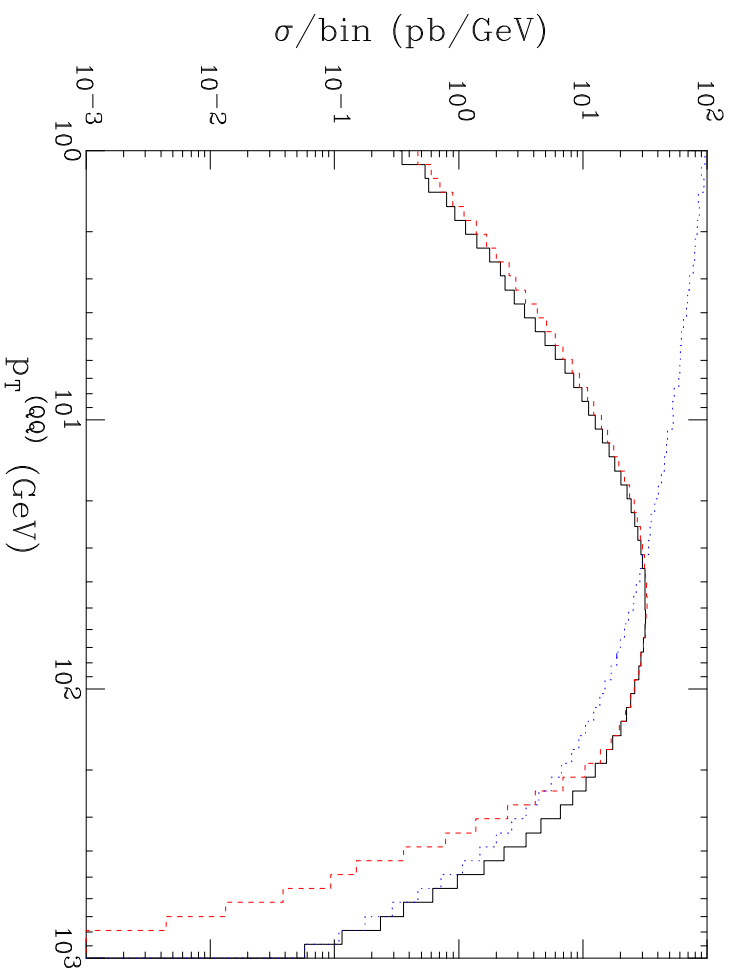
MC@NLO \approx NLO here.

New feature in MC: $Q\bar{Q}$
asymmetry at Tevatron.

Top Rapidity Asymmetry at Tevatron



$t\bar{t}$ Correlations at LHC



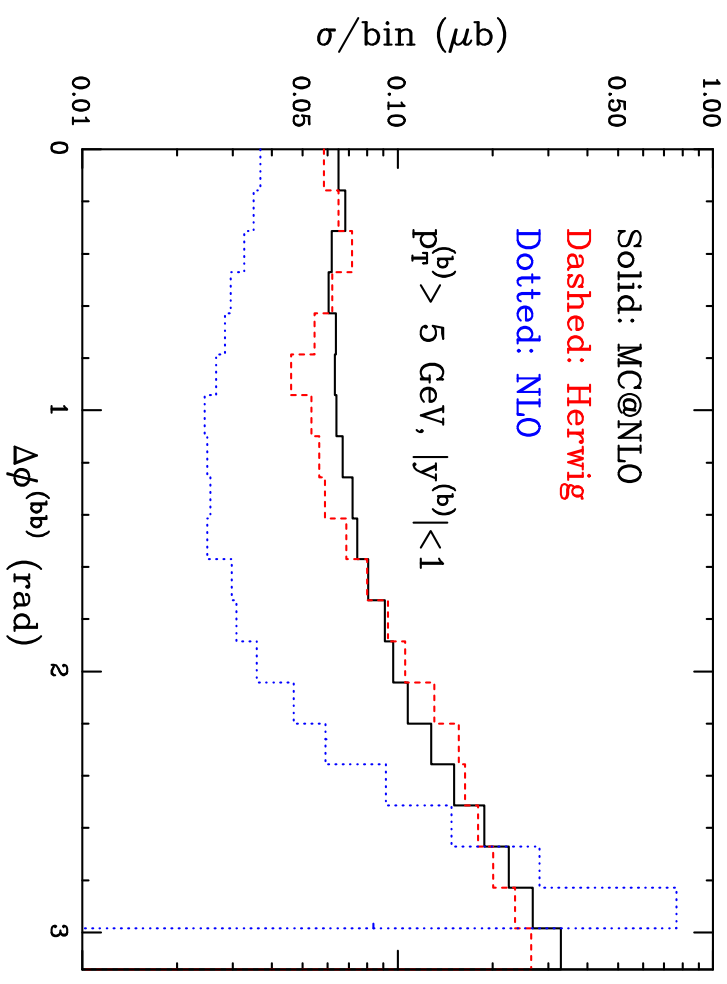
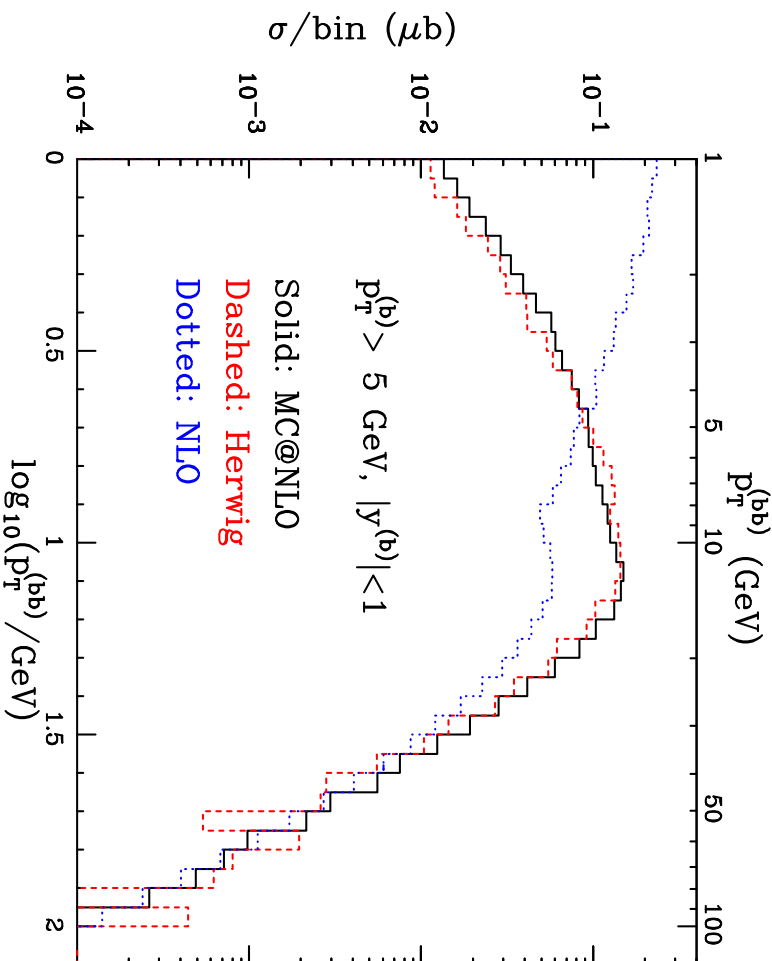
These correlations display the same patterns as those for vector boson pair production. Hard- and soft-scale physics are both treated correctly.

Solid: MC@NLO

Dashed: HERWIG $\times \frac{\sigma_{NLO}}{\sigma_{LO}}$

Dotted: NLO

$b\bar{b}$ Correlations at Tevatron

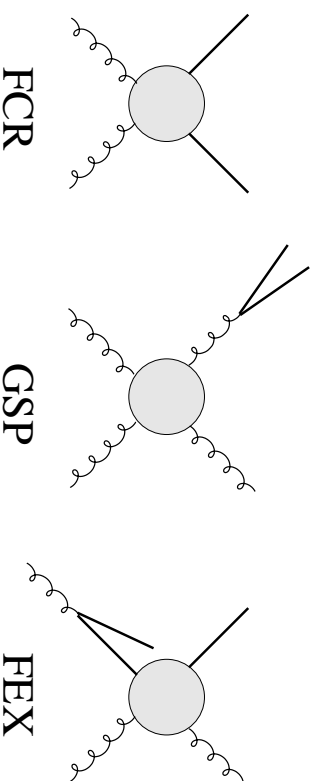


HERWIG does well (after cuts) but needs much more CPU: 14 million events vs 1 million for MC@NLO

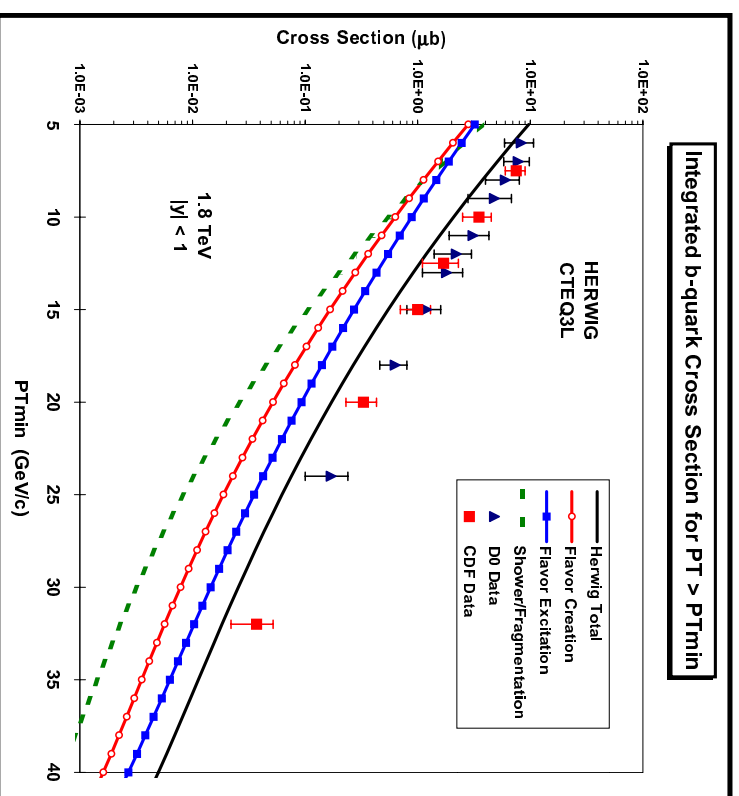
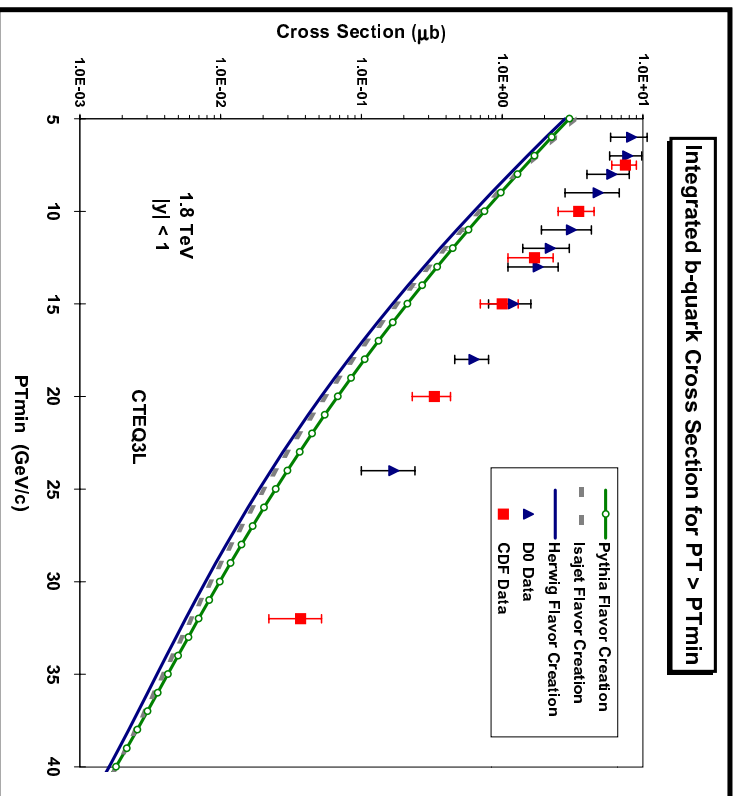
Solid: MC@NLO
Dashed: HERWIG (no K-factor)
Dotted: NLO

b Production with HERWIG

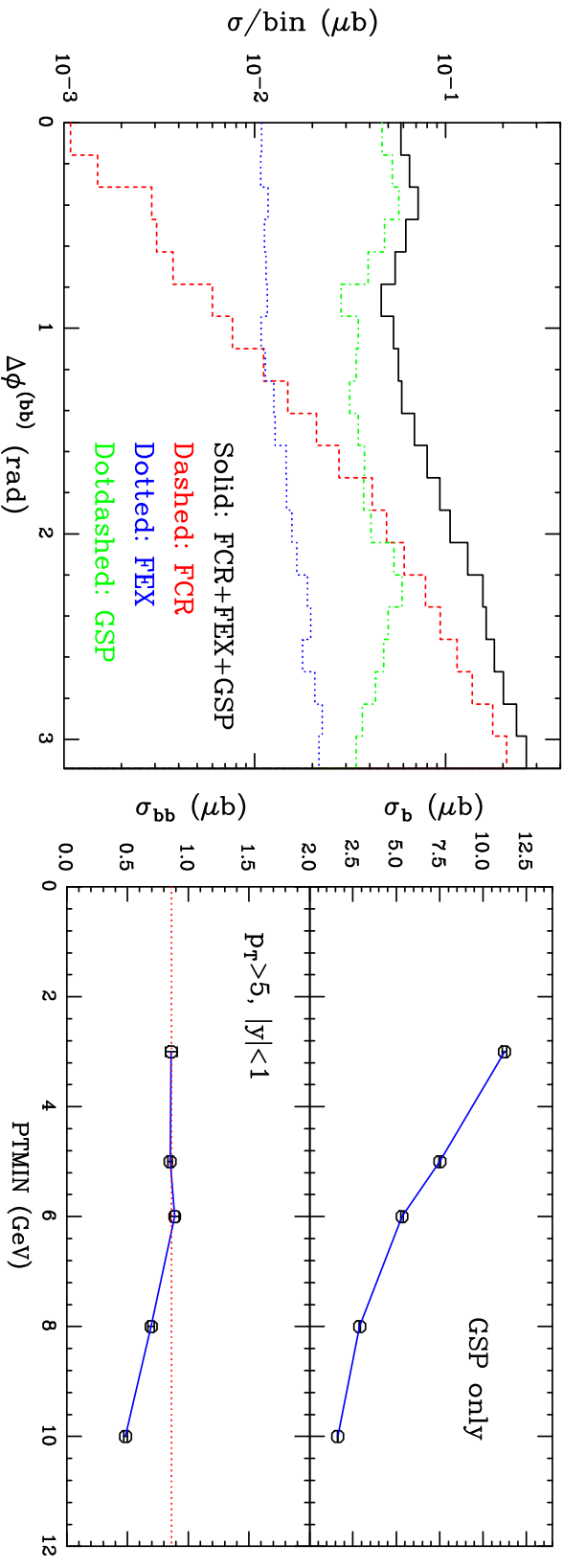
- In parton shower MC's, 3 classes of processes can contribute:



- All are needed to get close to data (RD Field, hep-ph/0201112):



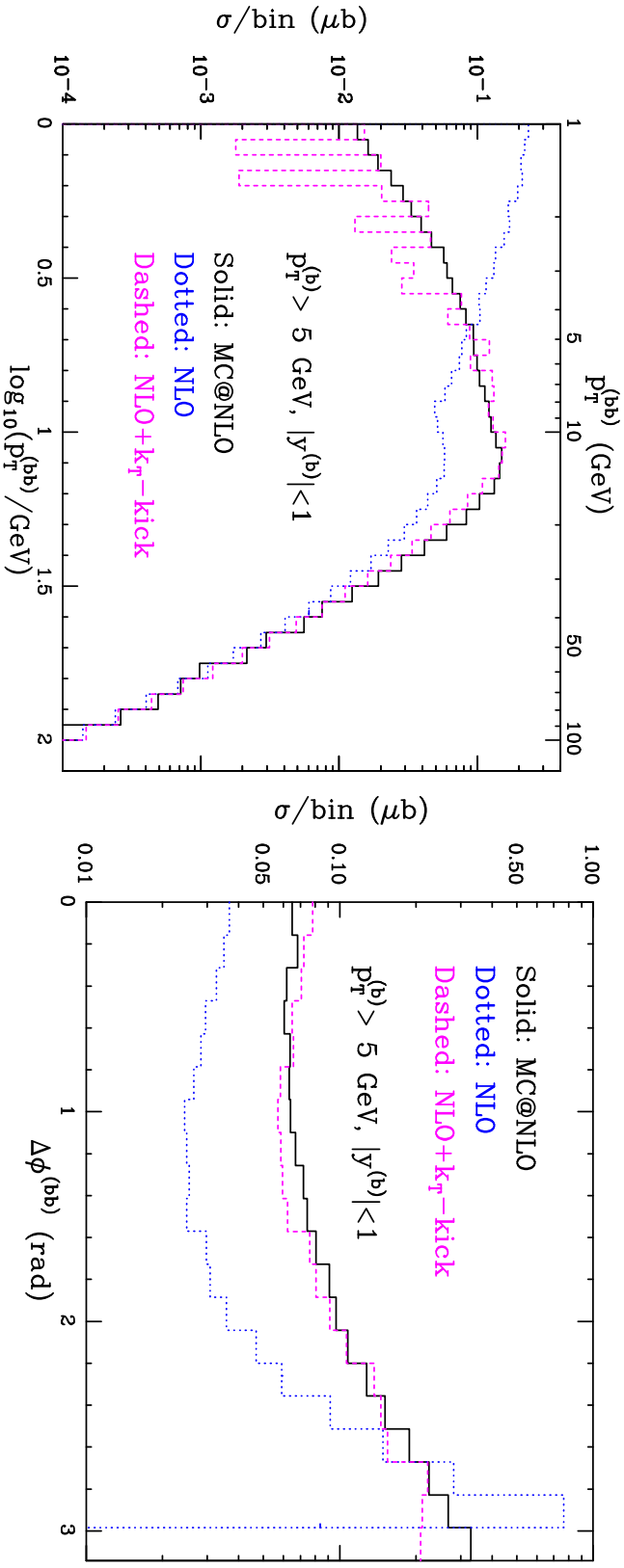
GSP and FEX contributions in HERWIG



- GSP, FEX and FCR are complementary and all must be generated
 - ❖ GSP cutoff (PTMIN) sensitivity depends on cuts and observable
 - ❖ FEX sensitive to bottom PDF
 - ❖ GSP efficiency very poor, $\sim 10^{-4}$
- All these problems are avoided with MCC@NLO!

NLO + k_T -kick vs MC@NLO

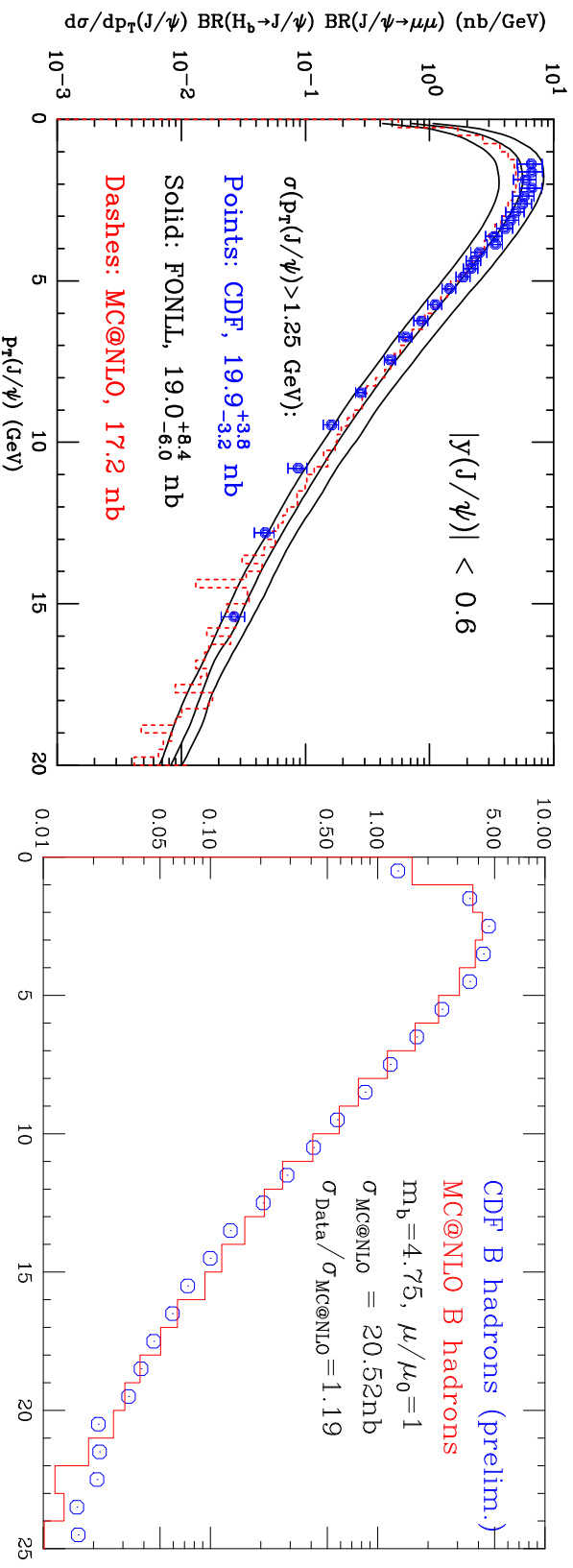
- (NLO + k_T -kick) with $\langle k_T \rangle = 4 \text{ GeV} \simeq \text{MC@NLO}$ (at Tevatron)



- This does **NOT** mean that there is $\langle k_T \rangle = 4 \text{ GeV}$ inside proton: it simply emulates the effect of initial-state parton showers.

Hadron-level Results on B production

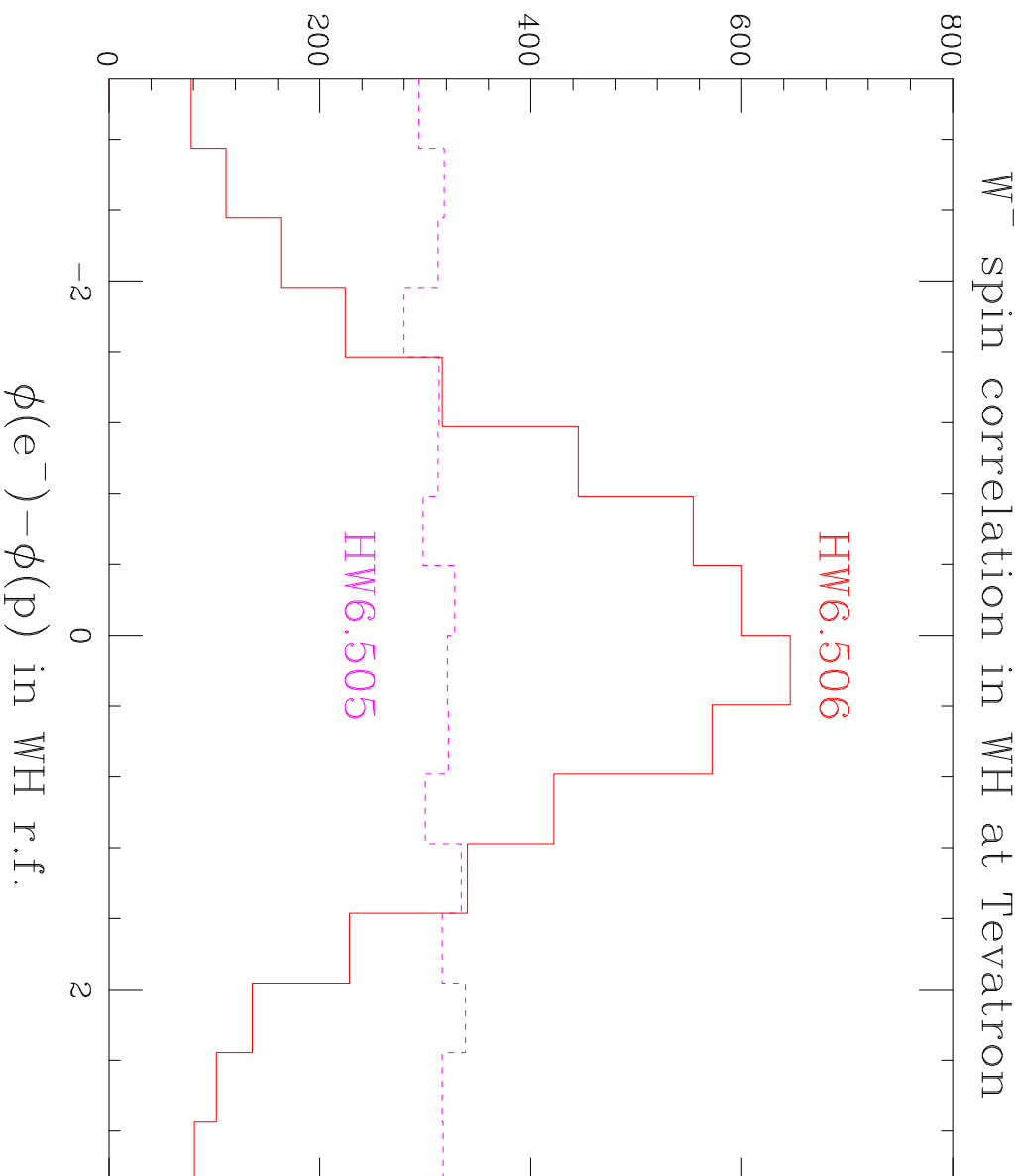
- $B \rightarrow J/\psi$ results from Tevatron Run II \Rightarrow B hadrons (includes BR's)



- No significant discrepancy!

Associated Higgs + Vector Boson Production

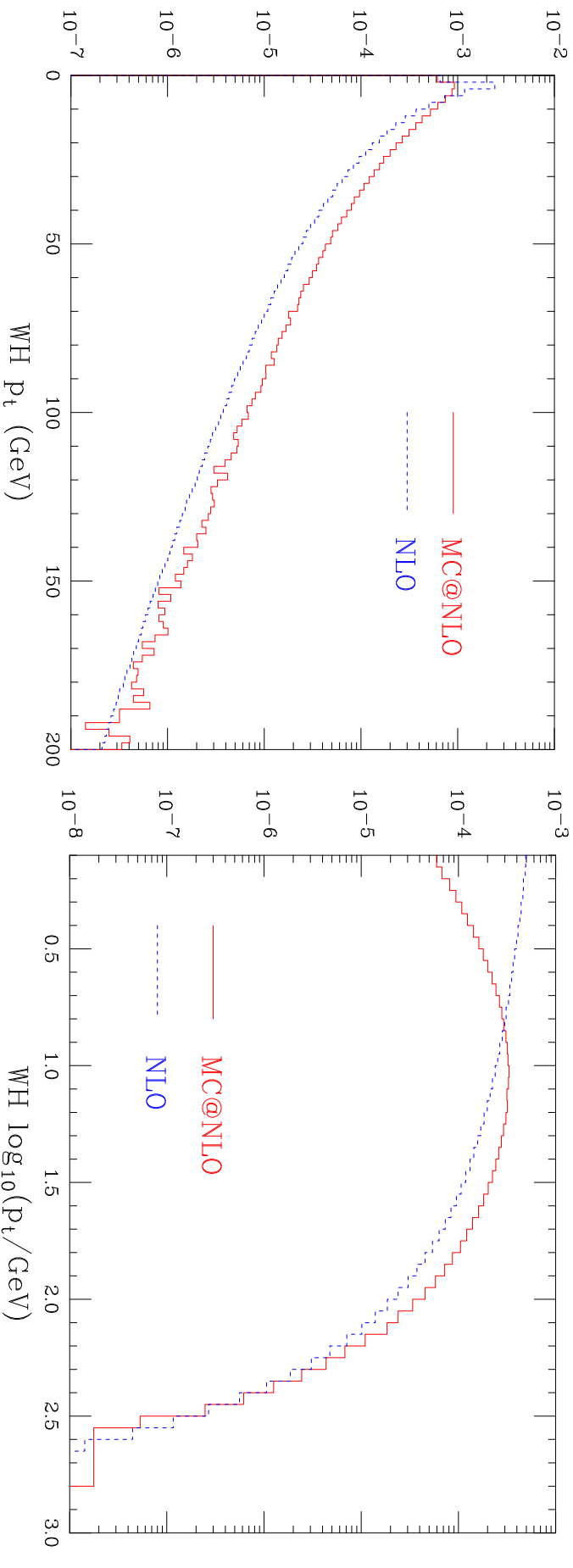
- Associated Higgs production implemented with full decay correlations



- LO in HERWIG 6.506, NLO in MC@NLO 3.1 (in preparation)

Associated Higgs + Vector Boson Production

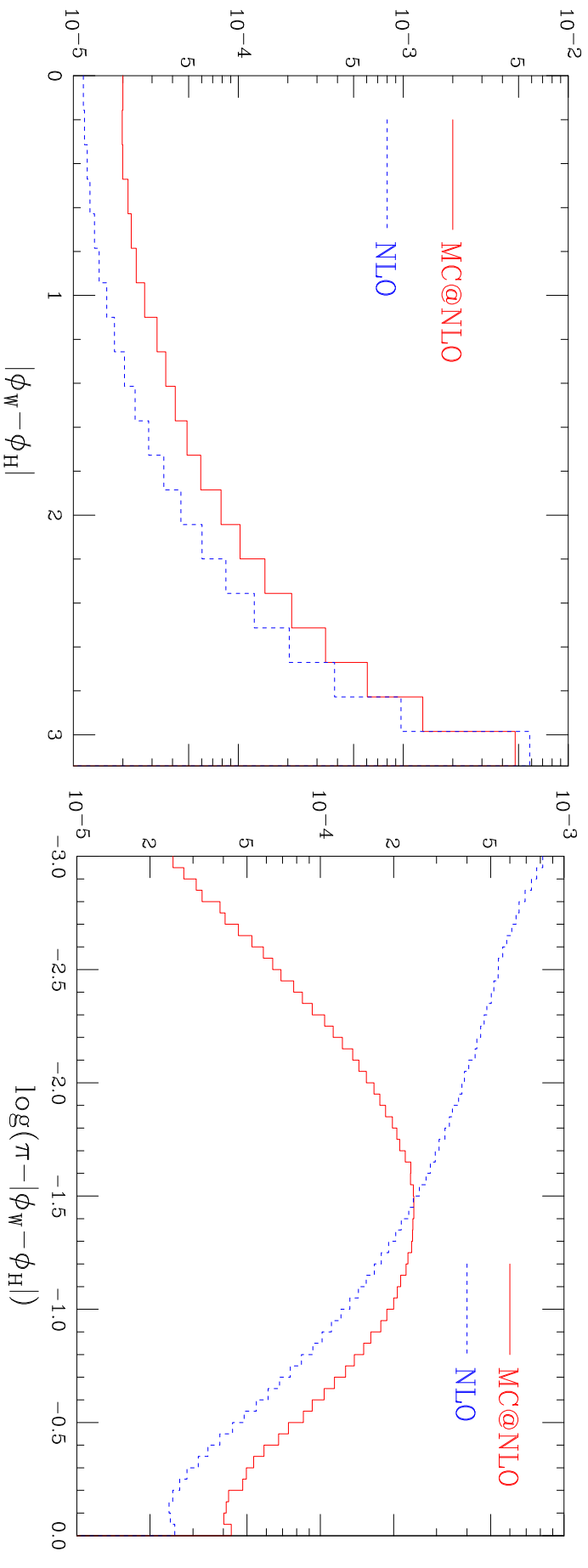
- p_t of WH pair in $p\bar{p} \rightarrow W^+H^0X$ at Tev II



- Qualitatively similar to WW

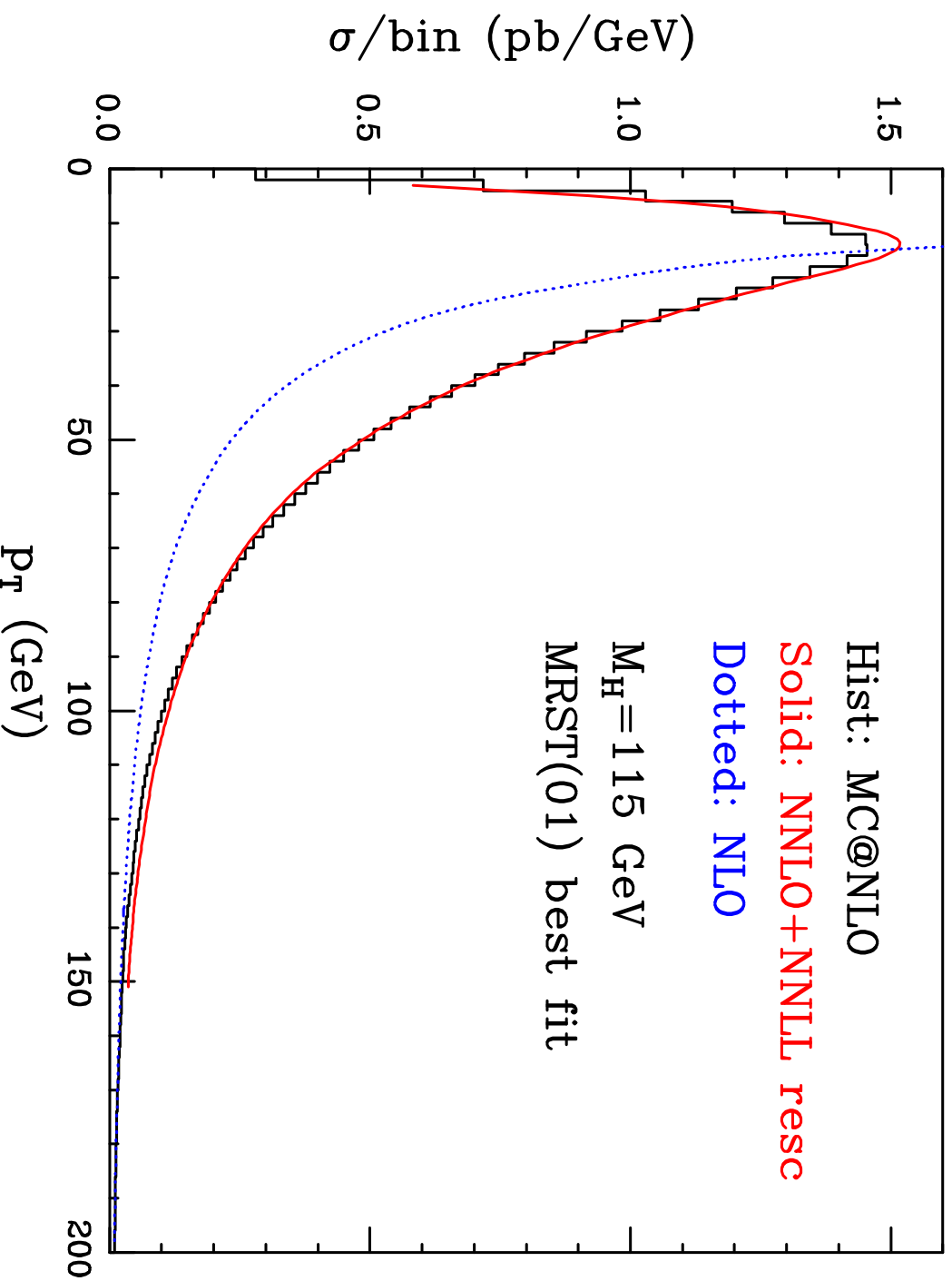
Associated Higgs + Vector Boson Production (cont'd)

- WH azimuthal separation in $\bar{p}p \rightarrow W^+H^0X$ at Tev II



Unassociated Higgs Production at LHC

- Good agreement with (N)NLO+NNLL



Conclusions and Future Prospects

- MC@NLO exists and works well for W, Z, H, WW, WZ, ZZ, WH, ZH, $t\bar{t}$ and $b\bar{b}$ production. Negative weights $\sim 10\%$ ($t\bar{t}$) to 20% ($b\bar{b}$) not a problem.
- Decay correlations implemented for W, Z, WH, ZH, not yet for others.
- Jet production needs more work.
- Shower modification to avoid negative weights looks possible (P Nason).
- General interface to NLO (subtraction method) programs feasible.