

PHYSICS OF THE LINEAR COLLIDER

F. RICHARD

*Laboratoire de l'Accélérateur Linéaire, IN2P3-CNRS et Université de Paris-Sud,
F-91405 Orsay Cedex, France
Email: richard@lal.in2p3.fr*

This presentation intends to illustrate the specific capabilities of an e^+e^- sub-TeV collider to provide answers on the basic issues in physics: the origin of mass, hierarchy of masses, and cosmological problems. Some foreseeable scenarios are discussed with a possible synergy with the LHC.

1. Introduction

In this presentation I intend to summarize the main physics prospects of the Linear Collider (LC) presently under consideration in North America, Asia and Europe. These prospects have been studied by 3 communities and there exist various documents describing them in detail.¹

After recalling briefly the baseline for the future LC, I will mention some important features of the detectors under consideration and the requirements needed for the machine parameters.

Concerning the physics, I will focus on 3 aspects.

1. The mechanism of electro-weak symmetry breaking (EWSB), in other words what is the origin of mass in particle physics. This aspect will be my main emphasis.
2. The problem of mass hierarchy, in particular in the Higgs sector, and the need for new mechanisms beyond the Standard Model (SM) like Supersymmetry (SUSY).
3. The input on cosmology of this model which can explain the origin of Dark Matter in the universe.

In a short presentation one can only give a very partial view of ongoing studies performed in the 3 regions. Examples of uncovered or very partially covered topics are: e^+e^- physics at MultiTeV (CLIC scheme²), the various SUSY breaking scenarios, SUSY and CP-violation, SUSY and the neutrino sector, extra dimensions with different schemes either alternate or combined with SUSY, e^-e^- , γe and $\gamma\gamma$ physics, precise tests of QCD with a LC.

2. The TeV LC

The present goal is to construct an e^+e^- LC covering an energy between the Z boson mass and 500 GeV, with polarized electrons (at least 80%) and collecting 500 fb^{-1} in the first four years of running. This LC should be able to reach, in a second stage, an energy $\sim 1 \text{ TeV}$, collecting about $500 \text{ fb}^{-1}/\text{year}$.

Various options are considered, which would require additional equipment and whose priorities will depend on the physics priorities emerging after LHC and LC operation:

1. positron polarization, at the 60% level, is not easy to implement but is needed to fully exploit the precision of a Z factory (the GigaZ scheme). This polarization is also needed for transverse polarization measurements which have recently been emphasized;³
2. an e^-e^- collider, is possible with reduced hardware changes; and
3. a $\gamma\gamma$ (or a γe collider), which could operate with a maximum energy $\sim 80\%$ of the nominal energy. This scheme would require major changes in the interaction region. As for the e^-e^- case, there will be a reduction in luminosity at the same energy but a high degree of polarization is feasible. Contrary to e^+e^- , a zero crossing angle is not possible even with the superconductive technology.

3. The Detector

Given the LC luminosity, two orders of magnitude above LEP, in several analyses the precision will be limited by systematic uncertainties. Part of these uncertainties come from detector limitations but this will improve using better technologies.

1. Improved vertexing which will allow a clean and efficient separation of charm quarks and also tagging of tau leptons. The improvement with respect to LEP is due to a small radius beam pipe and to the resolving power of thin pixel detectors.
2. Improved energy flow: the aim is to improve the jet energy resolution by about a factor of 2 with respect to LEP/SLD with fine segmentation of the calorimeters which are inside the magnetic coil. This improvement should allow one to cope efficiently with 6/8 jet topologies from ZHH and ttH channels. This resolution will also allow a clean separation between ZZ and WW hadronic final states to isolate the $WW\nu\bar{\nu}$ channel.
3. Momentum resolution will be improved by a factor 10 with respect to LEP/SLD, with polar angle coverage down to 100 mrad.
4. Hermeticity on energetic γ 's and electrons should go down to about 5 mrad with instrumented masks and good segmentation of the very forward calorimeter.

A detailed discussion of the new technologies implied can be found in the presentation given by T. Behnke at this conference.

A limiting factor will also come from our knowledge of the differential luminosity, of the polarization and of the energy calibration. Physics groups and machine experts are actively investigating these issues.

4. Which Physics Scenario for EWSB ?

This is clearly the central issue for the LHC and the LC. While there is no doubt that the LHC should discover a SM/MSSM type Higgs boson, one should be prepared for unconventional answers from Nature and I will illustrate this schematically in the following sections.

From LEP/SLD/Tevatron precision measurements (PM) one can derive 2 important consequences.

1. The Standard Model (SM) or its supersymmetric (SUSY) minimal extension (MSSM) are compatible with PM.

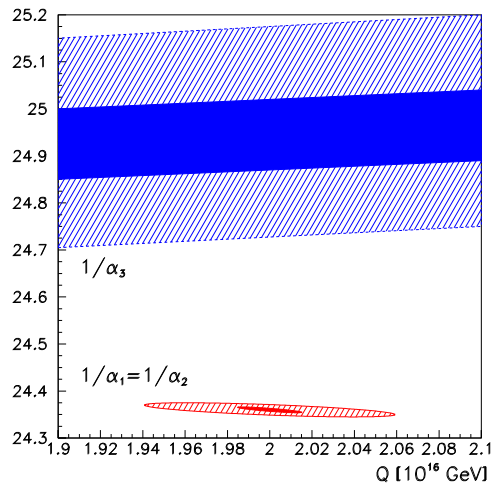


Figure 1. This figure shows that while LEP/SLD data (hatched) are compatible with SUSY-GUT, there is an indication of a discrepancy which, if real, should become clear with the precision of GigaZ (full).

2. Unification of the 3 interactions occurs at $\sim 2 \times 10^{16}$ GeV provided that the MSSM mass spectrum is at ~ 1 TeV (see Fig. 1⁴).

A light Higgs is therefore expected, below 130 GeV as a consequence of MSSM, below 250 GeV as a consequence of PM assuming that there is no contribution other than SM.

With a critical view⁵ of PM, one can notice that $\sin^2\theta_W$ from the charge asymmetry on b quarks measured at LEP1 is hardly consistent with the value obtained at SLD from polarization asymmetry.⁶ So far no compelling explanation has emerged from theory nor experiment. New physics interpretations, although not impossible,⁷ are severely limited by the absence of a significant deviation on the b quark cross section.

Another discrepancy, found in NuTeV, is still under investigation, but it should be noticed that this determination of $\sin^2\theta_W$ has no sensitivity to the value of the Higgs boson mass.⁶

When removing from the fit the b quark asymmetry result one finds that the Higgs mass value indicated by the data is too low given the direct search limit. At this stage one can remark that the effect is at the 2 sigma level and that the recent update on the top mass found by D0 with Run 1 data and the shift in the W mass after a reanaly-

sis of ALEPH data should even reduce it further. In our opinion one should therefore wait for the new top mass measurements at FNAL before drawing any definite conclusion on this discrepancy.

Another possibility could be that there are non SM/MSSM contributions coming from an alternate scheme (see elsewhere⁸ for a general discussion). I will therefore consider, for simplicity, that a LC could be dealing with three EWSB scenarios.

1. There is a light Higgs boson consistent with MSSM. There will then be emphasis on the precise measurement of Higgs couplings based on the collection of $\sim 10^5$ HZ events.
2. There is a Higgs boson but with a mass incompatible with SM/MSSM. A LC would then focus its effort in understanding the underlying mechanism and search for direct or indirect signals of new physics.
3. There is no Higgs boson, implying that the longitudinally polarized gauge bosons will strongly interact. A LC would then run at its maximal energy and focus on WW final states.

I will illustrate with some examples how a sub-TeV LC can provide the appropriate answers in these three scenarios.

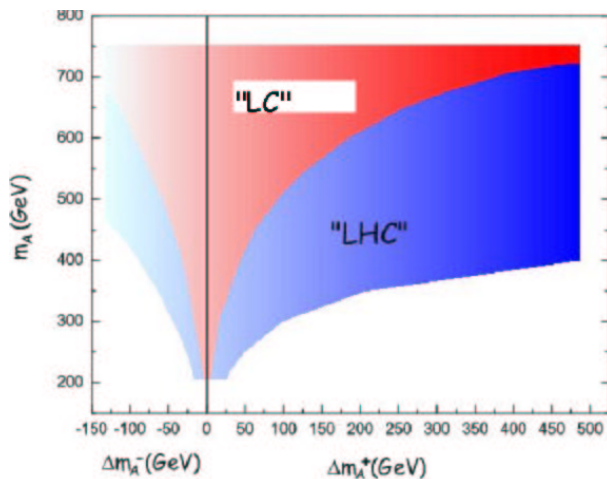


Figure 2. Sensitivity to the pseudoscalar Higgs boson mass at LC and LHC. This figure assumes an integrated luminosity of 300 fb^{-1} and the combination of 2 experiments for LHC.

4.1. MSSM Scenario

The first questions are obviously: are we dealing with a CP-even scalar, are the coupling to fermions and to bosons consistent with MSSM?

1. A LC will identify unambiguously the spin-parity of a Higgs boson by measuring the shape of the HZ cross section near threshold and the angular distributions provided by the ZH channel.
2. All fermionic (except for the top) and bosonic couplings will be measured at the % level for tree level couplings, at 5% for the gluonic width Γ_{gg} and 20% for $\Gamma_{\gamma\gamma}$. For the latter one can reach the % level using the photon collider scheme.
3. The Higgs coupling to the top quark can be measured with a 7–15% accuracy in the 120–200 GeV mass range. For the Higgs self-coupling λ_{HHH} the accuracy would be 10% (20%) at 800 GeV (500 GeV) center-of-mass energy.

With such accuracies one can detect with high sensitivity the presence of a non-standard component.

1. In the MSSM itself, it is possible to detect the influence of the heavier Higgs bosons. As shown in Fig. 2, one can estimate⁹ the mass of these bosons up to about 700 GeV. This goes beyond the mass reach of an e^+e^- LC but, given this information, could motivate a photon collider scheme in which these Higgs bosons can be singly produced.

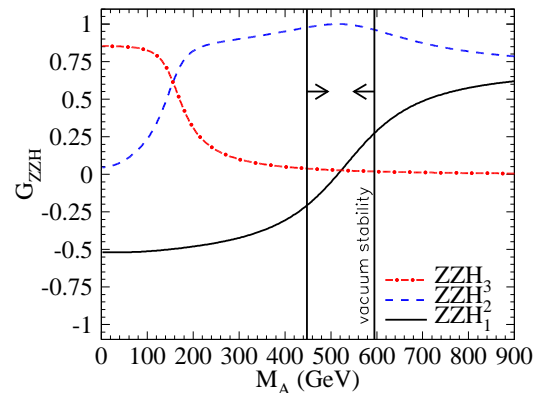


Figure 3. Variation of the ZZH coupling with the mass of the pseudoscalar A , in an NMSSM model, for the 3 scalar states.

2. Beyond MSSM one could have CP-violation in the Higgs sector or NMSSM with an additional Higgs isosinglet which can mix with the isodoublets. This could result in a significant drop in the ZZH coupling as shown¹⁰ in Fig. 3. In this respect a LC is very robust and can stand a reduction of a factor 100 in the ZH cross section.
3. Similarly it is worth recalling that the Higgs detection does not depend on the final-state branching ratios, in particular the LC can very well detect an “invisible” Higgs, which may occur in various schemes.¹¹ Moreover if Γ_{inv} is at the 2% level, the LC could still give a 5 standard deviation evidence on the presence of an invisible channel. An example¹² is given within SUSY where one assumes an unusual hierarchy between gaugino masses $M1$ and $M2$ ($M1 \sim M2/10$), such that the limit from LEP2 on the chargino does not eliminate the possibility of a very light LSP. Figure 4 shows the large effects possible in this scheme.

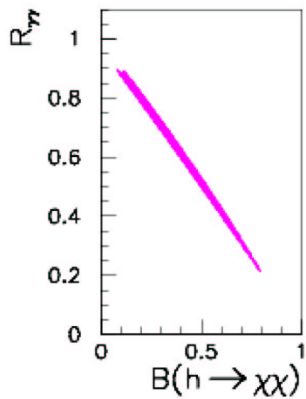


Figure 4. Variation of $R_{\gamma\gamma}$, the rate at LHC of Higgs into a pair of photons normalized to the SM rate, versus the branching ratio of the Higgs into a pair of LSP in a model with non universal gaugino masses.

4. Theories with large extra dimensions also provide valid schemes for EWSB. Within these theories the radion is a scalar field introduced to stabilize the small dimensions. It can therefore mix with the Higgs boson and, accordingly, modify the couplings of the Higgs to bosons and fermions as seen in Fig. 5. This figure¹³ shows that the modification is similar for W and fermions and therefore cancels in the ra-

tio. A LC however has access to the couplings and would measure this variation. The gluon pair BR should also provide a valid input in the interpretation of this effect.

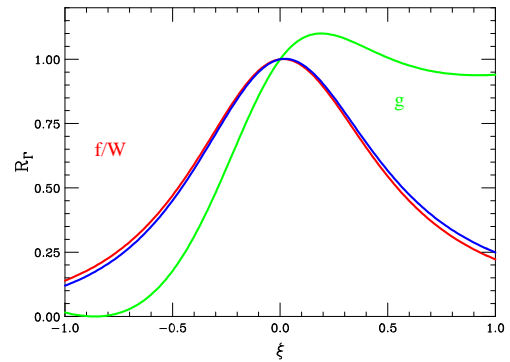


Figure 5. Ratio of the Higgs widths to their SM value for fermions, W bosons and gluons as a function of the mixing parameter in a radion model, assuming a Higgs mass of 125 GeV.

4.2. Quantum Level Consistency

The accuracy on the indirect determination of the Higgs mass can be improved by an order of magnitude with respect to LEP1/SLD. GigaZ should measure $\sin^2\theta_W$ with an error $\sim 10^{-5}$ provided one has polarized positrons. Figure 6 shows the other limiting accuracies and how they should evolve.

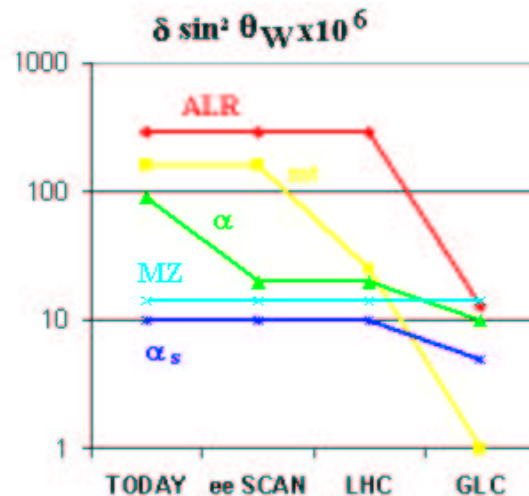


Figure 6. Predicted evolution of the experimental and theoretical accuracy on $\sin^2\theta_W$ with the expected progress on the polarization asymmetry, the top mass measurement, the decrease in uncertainty on $\alpha(M_Z)$ provided by e^+e^- scans and the improved precision on α_s at GigaZ.

The B - and K -factors should provide the input needed to settle the issue on $\alpha(M_Z)$ while the top mass accuracy with a LC should eliminate this source of error.

In parallel, there is a continuous and fruitful effort in reducing the theoretical uncertainty which is now coordinated in the “LoopVerein” working group.¹⁴ As an example, Fig. 7 indicates the continuous progress¹⁵ achieved on theoretical errors for the luminosity. For measurements which concern $\sin^2\theta_W$, the present theoretical uncertainty¹⁶ of $\sim 6\times 10^{-5}$ has to be reduced to match the expectations given in Fig. 6.

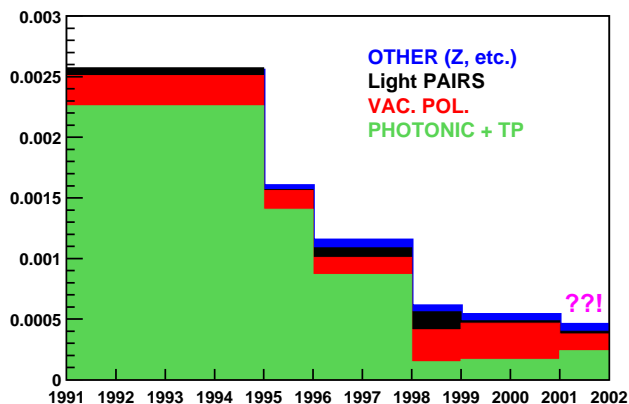


Figure 7. Components of the luminosity theoretical error at LEP1.

With GigaZ one could therefore test at the 5% level the equality:

$$M_H^{Direct} = M_H^{Indirect}$$

and draw valid conclusions on the global consistency of SM or MSSM.

An alternative to GigaZ, if positron polarization is not available, would be an improved measurement of the W mass using a threshold scan. The critical issue, there, is energy calibration: one has to extrapolate the beam measurement from the Z mass to the W threshold with a precision of $\sim 5\times 10^{-5}$. One expects $\delta M_W \sim 6$ MeV, which gives a 10% accuracy on the indirect Higgs boson mass.

4.3. Non-MSSM Higgs Scenario

Let us assume that a Higgs boson has been found with a mass inconsistent with PM, meaning above 200 GeV. This mass is therefore also inconsistent with MSSM. While PM are still constraining the gauge sector since, for instance, one could measure the BR of the Higgs into WW at 5% with $M_h=250$ GeV, the main mission of a LC would then be to find the “guilty part”.

If there is direct evidence for new physics at the LHC, e.g. if one observes a candidate Z' , the LC can decipher the message with precision measurements in the channel $e^+e^- \rightarrow f\bar{f}$.

With measurements in the TeV range, the LC provides high discrimination against the various Z' predicted by several symmetry groups¹⁷ as shown in Fig. 8. Knowing the Z' mass from LHC, LC PM at high energy will extract the vector and axial couplings of this Z' .

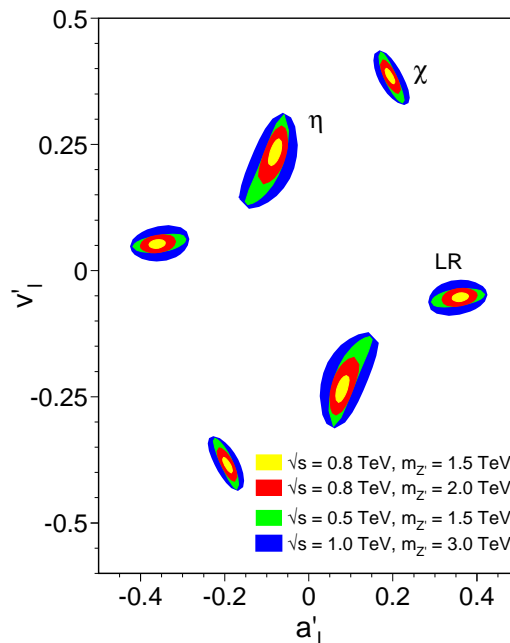


Figure 8. Accuracies expected on Z' vectorial and axial couplings from a LC operating up to 800 GeV for various extensions of the SM.

Similarly one can use the GigaZ to measure Z - Z' mixing, providing extra information.¹⁸ These measurements should also allow one to restore the consistency of PM with the observed Higgs mass.

As shown in Fig. 9 the mass domain of a LC covers¹⁸ and, in some instances, surpasses, the LHC

domain. It may therefore turn out that only the LC information will be left to solve the puzzle of a Higgs mass inconsistent with PM.

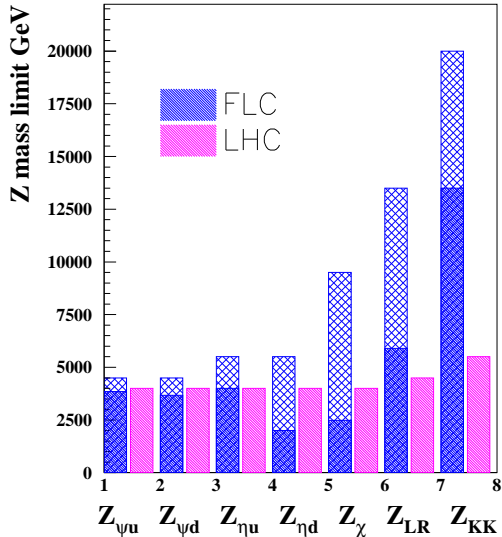


Figure 9. Mass domain covered by LHC (shaded/pink) and LC (hatched/blue) for various scenarios predicting a Z' . In the case of the LC the dark hatched part corresponds to the limit of sensitivity through mixing, while the light hatched part corresponds to the sensitivity through interference for a LC operating at 800 GeV.

In some cases the Z' can decouple from ordinary fermions as in the UED model (Universal Extra Dimensions¹⁹) where the new particles are mass degenerate Kaluza-Klein excitations of ordinary particles which carry a conserved quantum number and therefore need to be pair-produced. Remarkably, given the absence of direct coupling to fermions, this model has much weaker mass bounds than usual extensions with new Z' or extra dimensions and therefore it is not yet excluded that there could be pair-production at a TeV LC. If not, there still remains the effect on weak isospin violation in the top sector which contributes to the ρ parameters and therefore is measurable at the GigaZ.

4.4. The Little Higgs Scenario

Among the possible non-MSSM extensions of the SM, the “Little Higgs” model²⁰ has received special attention since it offers a viable solution to the hierarchy problem in the Higgs sector. In this scheme, the Higgs boson is a pseudo-goldstone boson originating from a symmetry broken at 10–100 TeV.

There is a perturbative theory below this scale and quadratic divergences of the Higgs mass are cancelled at first order by the contribution of new particles originating from the new symmetries. In particular there could be a light U(1) gauge boson, the B' , which contributes to the ρ parameter in the minimal model called “Littlest Higgs”.

If the LHC finds this B' then, as stated previously, the LC will allow one to identify its origin. If not, the LC can predict¹⁸ the mass of this object from PM and indicate which improvements in luminosity/energy are needed at the LHC (or at future colliders) to discover it.

To conclude on this non-MSSM Higgs scenario, one can say that it shows in a quantitative form how there could be a strong synergy between the LC and the LHC and that it suggests that we may need the full information of a LC from the Z -pole to the maximum energy.

4.5. No Higgs Scenario

In the case of no Higgs or, equivalently of a very heavy Higgs boson, with a mass above 1 TeV, one needs, even more than for the previous scenario, some kind of conspiracy cancelling the contribution of the Higgs term in PM.

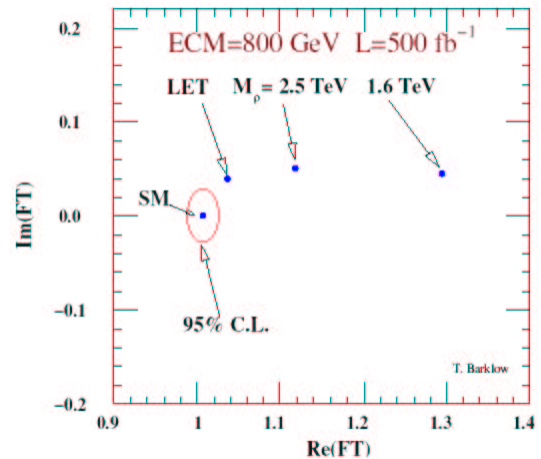


Figure 10. Sensitivity on the W_L form factor FT with a LC operating at 800 GeV and an integrated luminosity of 500 fb^{-1} when there is a ρ resonance at 1.6 TeV or 2.5 TeV or, with no resonance, the effect as predicted by the Low Energy Theorem.

The most direct manifestation of the heavy Higgs should occur in the gauge sector, where longitudinally polarized W bosons should strongly interact as a result of the absence of a Higgs exchange term cancelling the divergence in the process $W_L W_L \rightarrow W_L W_L$.

A likely scenario would be that the strong interaction, as in QCD, results into a ρ -type resonance with a mass below $\Lambda_{EWSB} = 4\pi v \sim 3$ TeV. Using the channel $e^+e^- \rightarrow W^+W^-$ and extracting the longitudinal part, one can observe a clear effect²¹ with a LC operating at 800 GeV as shown in Fig. 10.

Figure 11 displays the energy dependence of the signal if there is a resonance. Even if there is no resonance (Fig. 10), one should still observe a minimal effect, the so-called LET term predicted on a model-independent basis (Low Energy Theorem).

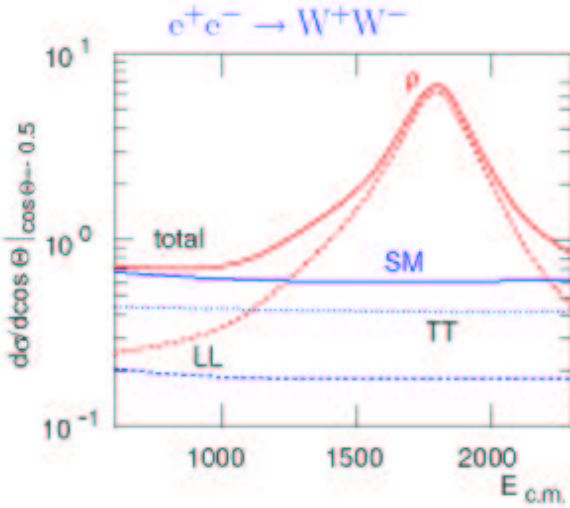


Figure 11. Variation with the LC energy of the backward differential cross section for W^+W^- with a ρ resonance. The upper dashed (red) curve corresponds to the LL component part to be compared to the lower dashed curve (blue) curve corresponding to the SM.

Table 1 summarizes the evidence foreseen at the LC and the LHC. One can clearly see an advantage for the LC if there is a resonance, while the 2 machines are roughly equivalent otherwise.

In case there is no resonance in the $J=1$ and $I=1$ channel, one could alternatively use the channels $e^+e^- \rightarrow \nu\bar{\nu}W^+W^-$ or $\gamma\gamma \rightarrow W^+W^-$ to access other quantum numbers.

One can draw similar conclusions in the language of triple gauge couplings, TGC, which should also

Table 1. Sensitivity of the LC and LHC to the presence of a strong interaction component in W^+W^- for different luminosities, energies and assuming the presence of a ρ resonance or without it (LET).

	\sqrt{s} GeV	\mathcal{L} fb $^{-1}$	M_ρ 1.6 TeV	LET
LC	0.5	300	16 σ	3 σ
LC	0.8	500	38 σ	6 σ
LC	1.5	200	204 σ	5 σ
LHC	14	100	6 σ	5 σ

manifest deviations in this scenario. One has 5 TGC preserving parity, custodial symmetry (to avoid effects on PM). Three are measured at GigaZ and in $e^+e^- \rightarrow W^+W^-$, 2 with $e^+e^- \rightarrow \nu\bar{\nu}W^+W^-$. These couplings can be expressed as:

$$\alpha = \frac{\Lambda_{EWSB}^2}{\Lambda^2}.$$

To be significant, a limit on Λ should be above Λ_{EWSB} . This condition is satisfied for the first 3 couplings (also true for the other 2) with the LC as shown in Fig. 12. Note that there is a significant improvement on the determination of these parameters with the input of the GigaZ.

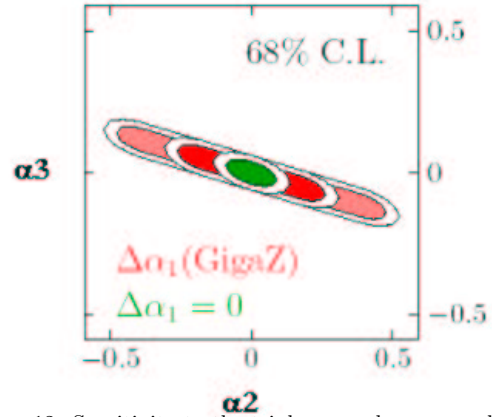


Figure 12. Sensitivity to the triple gauge boson couplings α_3 and α_2 with LC. The central ellipse assumes no error on α_1 , while the middle (dark red) one assumes a precision given by GigaZ.

5. The SUSY Scenario

Although there is very little experimental support for it, SUSY is considered as the leading candidate theory beyond the SM. On the theoretical side it provides a consistent way to understand EWSB through the Higgs mechanism (no hierarchy problem and

EWSB naturally driven by the large Yukawa coupling of the top quark). On the experimental side it allows unification of the 3 forces and it provides natural links to cosmology, in particular a mechanism for generating Dark Matter (DM) in the universe. The $g-2$ deviation with respect to the SM, which would constitute a precious indication of a SUSY effect, is still uncertain given the 2 contradictory results²² obtained for the hadronic correction. If confirmed it would favor a SUSY solution with light sleptons and light gauginos.

The basic issues for this theory, after it is revealed at the LHC or Tevatron, will be:

1. to fully confirm the SUSY hypothesis by measuring the spin and couplings (precisely predicted in this theory) of these particles. Recall that, e.g. in the UED framework,²³ one is able to fake the presence of SUSY but unable to pass the above criteria; and
2. to understand the SUSY breaking (SSB) mechanism, for which there is plethora of proposed schemes, none of them clearly emerging as preferable to the others.²⁰

For pedagogical reasons only, I will use the simplest of SSB schemes, that is mSUGRA in which one can represent the parameter space in two dimensions, in terms of the common scalar and gaugino masses at GUT m_0 and $M_{1/2}$. $\tan\beta$ is a free parameter and μ is derived in absolute value by imposing EWSB. Figure 13 shows²⁴ that various experimental conditions, in particular the need to provide the adequate amount of neutral DM and the need to generate EWSB, put severe restrictions on these parameters. One can distinguish, in Fig. 13, four allowed domains.

1. The *blob* domain, studied so far in LC and LHC studies. This region in which SUSY parameters are moderate is favored by the advocates of *Fine-Tuning criteria* (FT), meaning that one expects that the SUSY-EWSB generation of the W and Z mass should not result from fine-tuned cancellations between large SUSY masses. In this region the DM candidate is a Bino and the presence of light sleptons allows one to get sufficient t-channel annihilation to keep DM under control.

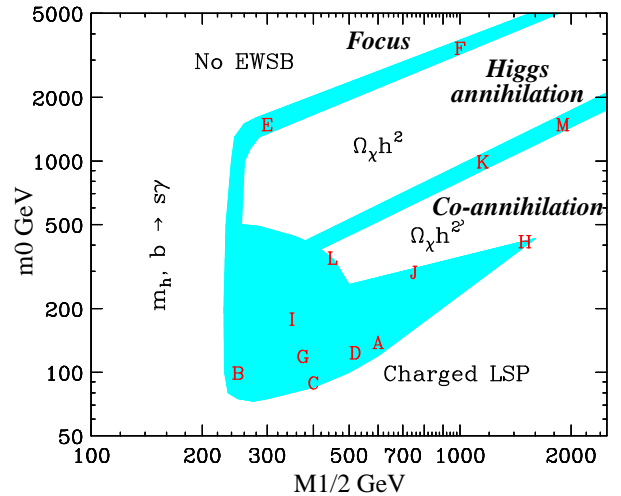


Figure 13. Allowed regions for the m_0 (scalar mass) and $M_{1/2}$ (gaugino mass) parameters within mSUGRA taking into account EWSB, radiative decay of the b quark and the DM constraints. The letters indicate the points which are proposed as benchmarks for the LHC and the LC. There are 3 regions called “focus” (large m_0 solutions), “co-annihilation” (sleptons almost degenerate with LSP) and “Higgs annihilation” (LSP about 1/2 the heavy Higgs mass) which allow large values of these parameters.

2. In the *co-annihilation* domain, the Bino LSP gets heavier and, for DM, one needs to assume that the LSP is almost degenerate in mass with the sleptons, in particular with the stau particle which is usually the lightest slepton. The recent results from WMAP, have further restricted this domain as will be discussed later (see Fig. 15).
3. The *Higgs annihilation* domain corresponds to a region of parameters for which the heavier Higgs particle have a mass close to twice the LSP mass, in such a way that efficient s-channel annihilation can occur to reduce the amount of DM. This solution allows one to reach high SUSY masses, not only beyond LC reach but even beyond the LHC.
4. The *focus* domain is remarkable in the sense that it can still claim absence of FT with large values of m_0 . This is so since the coefficient of m_0 in the EWSB equation is minute.²⁵ In this scenario, DM can be controlled by noting that the LSP can be a Higgsino if μ is small, favoring s-channel annihilation. Such a scenario could be quite peculiar since one might only see the 2 first lightest neutralinos and the first chargino, all with masses $\sim \mu$.

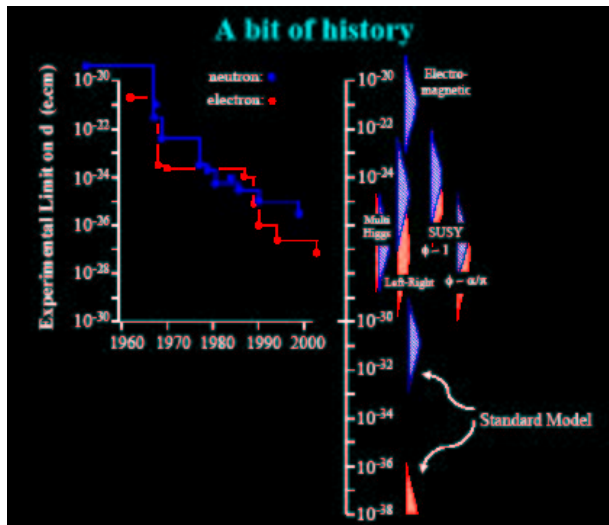


Figure 14. Experimental limit on EDM for neutrons and electrons versus time. On the right-hand-side are indicated regions predicted by various theories. The SUSY prediction assumes moderate scalar masses.

Before leaving this part of the discussion, let me point out an important experimental aspect of these arguments on the LSP sector. In scenario 2 and 4 there could be a degeneracy in mass between the sleptons and the LSP and between the gaugino themselves. This conclusion, simply expressed within mSUGRA is certainly quite general²⁶ and deserves attention. As shown at LEP2, techniques can be used to detect the charginos in such cases given the large production cross section involved. This might not be the case for the slepton with a much reduced cross section.

5.1. The Flavor Sector

Previous discussion could lead to the conclusion that nature will choose the less FT solution, that is the “blob” sector. This conclusion does not take into account the so-called *flavor problem* of SUSY.

This problem extends to the various observables where experimental information puts very severe limitations on SUSY parameters: FCNC constraints, CP-violation in the K sector, EDM limits for electrons and neutrons (in rapid progress²⁷ as shown in Fig. 14) and proton lifetime limits. Several authors propose to avoid these constraints by setting the scalar masses, at least for the first two families, to very high value. This could happen in the framework of the “focus” scenario without drastic FT.

There are other ways to avoid some of the problems and restore the “blob” scenario. One can for example postulate a hidden symmetry (of the Left-Right type²⁸) which forces the SUSY phases to 0, or assume a cancellation mechanism²⁹ which reduces the importance of CP-violation in the EDM process. The later seems to also imply some kind of fine-tuning given that one already needs to assume a very small phase for the μ term.

5.2. The Three Scenarios

From previous discussion and for the sake of simplicity, one may foresee 3 types of scenarios representing different challenges for a LC.

1. All SUSY particles are very heavy except the lightest Higgs boson h and, possibly, the two lightest neutralinos and the lightest chargino accessible at the LC and the gluino at the LHC. This would correspond to a “focus” type solution. The LSP could be a Higgsino with $|\mu| < M_1$ or a Wino if the usual inequality $M_1 < M_2$ is violated. In both cases there would be a substantial s-channel annihilation. The three lightest gauginos are almost degenerate in mass but, as shown at LEP2, observable at the LC.
2. Same as 1 but, in addition, the third generation of scalar particles is light. Then co-annihilation can take place between the LSP and the stau particle. As has been pointed out,³⁰ the new results from WMAP would require a very close degeneracy and therefore the stau would be extremely hard to observe. The good news however is that the WMAP results imply that the LSP cannot be heavier than ~ 500 GeV and therefore would fall within the range of observation of a LC.
3. The “blob” scenario which allows a wide range of possibilities subject however to the Higgs mass limit of LEP2 and, possibly, to the g-2 measurement from BNL. Figure 15 indicates a recent update³⁰ illustrating these features.

5.3. DM at the LC

If within reach, the LC will accurately measure the mass of the LSP and its couplings, in other words its Higgsino/Wino/Bino content.

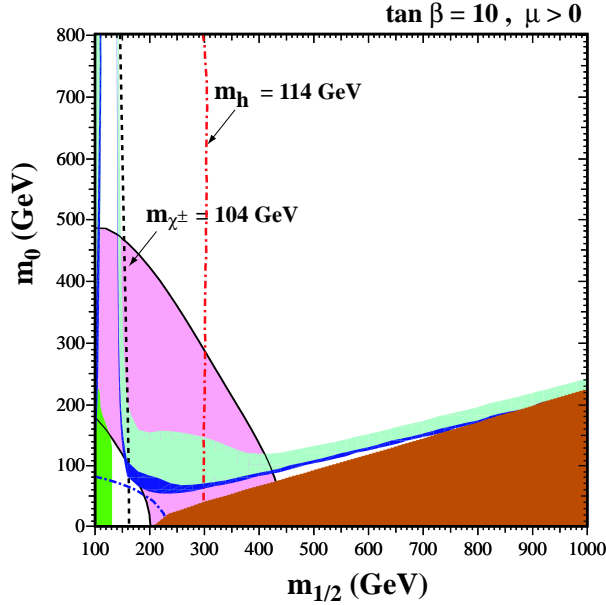


Figure 15. Allowed regions within mSUGRA assuming $\mu > 0$ and $\tan\beta=10$. Indicated also is the mass limit from LEP2 on charginos and on the Higgs boson mass. The small bands correspond to DM solutions with (darkest shaded region/dark blue) and without (lightest shaded region/pale green) WMAP derived results. The ~ 300 GeV wide band (decreasing from the mid-left side to the bottom/pink) is favoured by an interpretation of the g-2 results, still under investigation.

There are various tools to perform such an analysis, threshold scans, polarization asymmetries which I will not describe here but can be found elsewhere.¹⁷

These measurements will constitute an input for cosmology as will be illustrated shortly. They will also allow one to precisely interpret the non-accelerator searches on primordial DM performed with various techniques.³¹

Taking point B within mSUGRA (see Fig. 13), a case treated with LHC inputs,³² I have estimated the accuracy on $\Omega_{DM}h^2$ from SUSY using the code MicroMEGAs,³³ given that a LC can measure the LSP and τ slepton masses with accuracies of 0.1 GeV and 0.6 GeV respectively.¹⁷ Any significant discrepancy with cosmology may reveal extra sources of DM (e.g. axions or very heavy objects produced in the early phases of the universe).

It is also worth recalling¹⁷ that the chargino and neutralino measurements, if there is a Bino component, provide an indirect sensitivity on the masses of selectrons and sneutrinos masses which are precise up to masses ~ 1 TeV, that is well above the reach of

Table 2. Relative precision on DM obtained by cosmology in the present phase (WMAP) and foreseen (Planck), to be compared with the accuracy foreseen on its determination at the LHC and at the LC assuming point B as defined in Fig. 13.

WMAP	7%
LHC	$\sim 15\%$
Planck	$\sim 2\%$
LC	$\sim 3\%$

a TeV linear collider.

6. LC and the GUT Scale

As already stated, understanding the origin of SSB will be the major goal of our field once SUSY is discovered at the Tevatron or LHC. To reach this goal, it will be essential to measure SUSY masses and couplings precisely and in a model independent framework. As an example, in the gaugino sector, the LC should allow one to measure precisely M_1 and M_2 while M_3 is provided by the LHC from the gluino mass. These three quantities would be extrapolated to the GUT scale and provide a crucial test of SSB. Figure 16⁴ gives an example of such a confrontation in the case of a string inspired model.

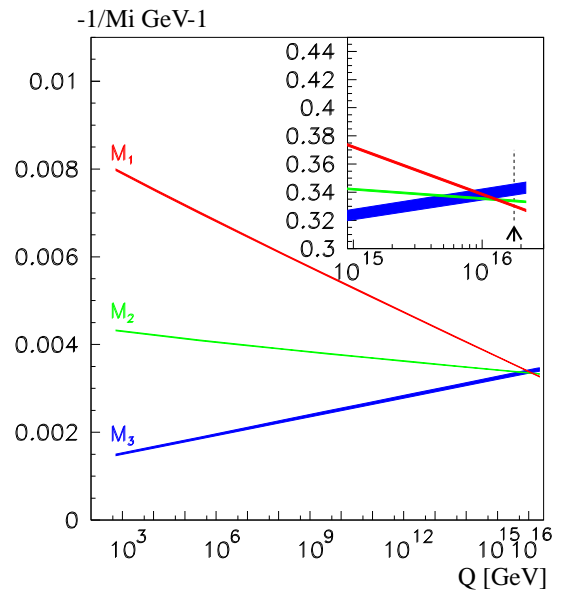


Figure 16. Evolution of the gaugino masses up to the GUT scale with a window showing the breaking of universality due to loop corrections in a string inspired model.

One can observe that while the convergence is satisfied, it occurs at a mass below the GUT scale given by the couplings. This shift is due to loop corrections present in such a scheme and the LC accuracy is needed to observe this subtle effect present in string theories.³⁴ This example shows how a LC may turn out to be crucial to understand the underlying fundamental theory.

Moreover one can observe in Fig. 16 the unmatched accuracy of M3. Recently³⁵ it has been shown that this accuracy can be greatly improved if the LHC can use LC information, in particular the mass determination of the LSP. One could therefore reduce the error given in Fig. 16 by a factor 3.

This example illustrates the potential of a LC+LHC combination to reach a very ambitious goal, pioneered by LEP/SLD: explore the GUT/string scale to reconstruct the fundamental parameters of an underlying theory which no accelerator will ever be able to test directly. Studies of this type are systematically being done in the framework of LHC/LC collaborative studies.³⁵

7. Summary: Why Do We Need a LC ?

From this very incomplete description of the prominent topics which will be addressed by a LC, one can derive the following list of goals.

1. To provide the full picture on an SM/MSSM Higgs.
2. To provide an answer on the issue of EWSB even in more difficult or unexpected situations, e.g. a reduced cross section, a heavy Higgs or no Higgs scenario.
3. To access to the SSB mechanism by combining LC and LHC measurements and, even more ambitiously, to access the GUT scale and to the parameters of the underlying string theory.
4. To predict very precisely, within SUSY, the amount of DM and confront it to the observation by cosmology.
5. To interpret unambiguously an unexpected discovery at LHC, e.g. a Z' or a Kaluza Klein excitation.
6. To estimate mass scales beyond LC/LHC reach:

- observing deviations on PM translated, e.g. into a heavy Higgs, a heavy sfermion or a Z' mass; and

- testing the theory at the quantum level and eventually predicting new mass scales as has already occurred with LEP/SLD/Tevatron for the Higgs mass.

These various examples indicate how PM at a LC may reveal new frontiers in energy which would call for an upgraded LHC or for a new generation of colliders like CLIC or VLHC.

Acknowledgments

It is a pleasure to thank and congratulate the team of this Symposium for the excellent organization of the meeting.

References

1. See the following web sites and documents referred therein:
 - For America <http://blueox.uoregon.edu/~lc/alcpag/>
 - For Asia <http://acfahep.kek.jp>
 - For Europe <http://www.desy.de/conferences/ecfa-desy-lcext.html>
2. CLIC Physics Study Group: <http://deroeck.home.cern.ch/deroeck/clic/spin2.html>
3. T.G. Rizzo, JHEP 0308:051,2003, hep-ph/0306283.
4. G.A. Blair *et al.*, *Eur. Phys. J. C* 27, 263-281 (2003).
5. M.S. Chanowitz, *Phys. Rev. D* 66 073002,2002, hep-ph/0207123.
6. Talk of P. Gambino at this conference.
7. D. Choudhury *et al.*, *Phys. Rev. D* 65 053002 (2002), hep-ph/0109097.
8. M.E. Peskin, J.D. Wells, *Phys. Rev. D* 64 093003 (2001), hep-ph/0101342.
9. E. Gross *et al.*, to be published.
10. D.J. Miller *et al.*, hep-ph/0304049.
11. Examples of scenarios with a Higgs decaying invisibly can be found in:
 - A. Djouadi, M. Drees hep-ph/9703452 (gravitino)
 - A.S. Joshipura, S.D. Rindani hep-ph/9205220 (Majorons)
 - G.F. Giudice *et al.*, hep-ph/0002178 (graviscalar)
 - N. Arkani-Hamed *et al.*, hep-ph/9811448 (KK of ν_R)
 - A. Widom, Y.N. Srivastava hep-ph/0004231 (gravitons).
12. F. Boudjema *et al.*, hep-ph/0206311.
13. J.L. Hewett and T.G. Rizzo, SNOWMASS-2001-P338, hep-ph/0112343.

14. <http://wwwth.mppmu.mpg.de/members/loopverein/TheLoopVerein.html> and references therein.
15. S. Jadach, hep-ph/0306083.
16. See for instance W. Hollik, talk given at the Mini-Workshop on Electroweak Precision Data and the Higgs Mass, Zeuthen, Germany, 28 Feb - 1 Mar 2003, DESY-PROC-2003-01.
17. TESLA TDR part 3. Physics at an $e^+ e^-$ Linear Collider, hep-ph/0106315.
18. F. Richard, hep-ph/0303107.
19. T. Appelquist and Ho-Ung Yee, *Phys. Rev. D* **67** 055002 (2003), hep-ph/0211023.
20. Talk of G.F. Giudice at this conference.
21. Proceedings of the 5th International Linear Collider Workshop (LCWS2000) and Reference.¹⁷
22. M. Davier *et al.*, LAL-03-50, hep-ph/0308213.
23. Hsin-Chia Cheng *et al.*, *Phys. Rev. D* **66** 056006 (2002), hep-ph/0205314.
24. J.R. Ellis, hep-ex/0210052.
25. J.L. Feng, K. T. Matchev, and T. Moroi, *Phys. Rev. D* **61** 075005 (2000).
26. D. Hooper and T. Plehn, *Phys. Lett. B* **562** 18-27 (2003), hep-ph/0212226.
27. See B. E. Sauer, status of EDM measurements in <http://frontierscience.lnf.infn.it/2002/talks/sauer.pdf>
28. K.S. Babu *et al.*, *Phys. Rev. D* **65** 016005 (2002), hep-ph/0107100.
29. T. Ibrahim and P. Nath, *Phys. Rev. D* **58** 111301 (1998) [Erratum-ibid. D 60, 099902 (1999)], hep-ph/9807501.
30. J.R. Ellis, *Phys. Lett. B* **565** 176-182.
31. Talk given by M. de Jesus at this Conference.
32. M. Battaglia *et al.*, hep-ph/0306219.
33. G. Belanger, *et al.*, hep-ph/0112278.
34. P. Binetruy, M.K. Gaillard and B.D. Nelson, *Nucl. Phys. B* **604** 32 (2001).
35. LHC/LC Study Group <http://www.ipp.dur.ac.uk/~georg/lhclc/> and references therein.

DISCUSSION

Chang Kee Jung (Stony Brook): What is the most current MGUT apparent value with an error? What is the MGUT value newly derived by this particular string model shown in your talk?

Francois Richard: In the plot shown, the gaugino masses unify at about 10^{16} GeV in contrast with the coupling constants which unify at twice this value. The claim is that such an effect can be seen at a LC. In this particular case the effect originates from loop corrections within a string inspired theory.

Thomas Germann (Zurich University): You mention the option of using transverse polarization. What are the physics observables to be probed with this?

Francois Richard: I had in mind the possibility of measuring polarization asymmetries which are insensitive to vector contributions but can reveal

tensor terms originating from low scale gravity contributions. This topic has been recently advocated by T. Rizzo *et al.*, and they find a sensitivity of this method up to very high mass scales.

Andreas Kronfeld (FNAL): If I understood your remarks on dark matter correctly, cosmology could determine the total component of dark matter at the 2% level. Meanwhile, LC could determine the contribution of a single identified component (e.g. the LSP) with similar precision. Then you could imagine $DM_{Total} = (30.0 \pm 0.5)\%$ and $DM_{LSP} = (20.0 \pm 0.5)\%$. That would be fascinating. Did I understand this correctly?

Francois Richard: Yes, there is no reason that the two numbers coincide but it matters that the precisions are similar for an optimal comparison.