

# MiniBooNE:

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### Mixing in Leptons and Quarks

$$egin{aligned} |d_lpha
angle &=\sum_i V^*_{lpha i} |d_i
angle \ |
u_lpha
angle &=\sum_i U^*_{lpha i} |
u_i
angle \end{aligned}$$

Lepton Sector: Neutrino oscillations

- Mass eigenstates ≠
   Flavor eigenstates
- Allows flavor-changing interactions
- No theoretical guidance

#### Quark Sector:

Flavor-changing decays Mixing/oscillations CP violation

$$\begin{split} P(\mathbf{v}_{\alpha} \rightarrow \mathbf{v}_{\beta}) = & \delta_{\alpha\beta} \\ & -4\sum_{i>j} \Re(\mathbf{U}) \sin^2[1.27\Delta m_{ij}^2(L/E)] \quad \text{sum over mass} \\ & +2\sum_{i>j} \Im(\mathbf{U}) \sin[2.54\Delta m_{ij}^2(L/E)] \quad \text{eigenstates} \end{split}$$

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Lepton Sector: Neutrino oscillations

 $P(\mathbf{v}_{lpha} 
ightarrow \mathbf{v}_{eta}) = \delta_{lpha eta}$ 

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ctor:<br/>ationsNeutrino of type α, energy E<br/>Traverses distance L<br/>Interacts as neutrino of type β<br/>Observe as: deficit of  $v_{\alpha}$ <br/>appearance of  $v_{\beta}$  $\delta_{\alpha\beta}$ appearance of  $v_{\beta}$  $-4\sum_{i>j}\Re(U)\sin^2[1.27\Delta m_{ij}^2(L/E)]$ <br/> $+2\sum_{i>j}\Im(U)\sin[2.54\Delta m_{ij}^2(L/E)]$ 

### Mixing in Leptons and Quarks

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#### Lepton Sector: Neutrino oscillations

 $\sin^2 2\theta_{12} \sin^2 [1.27 \Delta m_{12}^2 (L/E)]$ Two flavor oscillations

 $P(\mathbf{v}_{\alpha} \rightarrow \mathbf{v}_{\beta}) = \delta_{\alpha\beta} -4\sum_{i>j} \Re(\mathbf{U}) \sin^{2}[1.27\Delta m_{ij}^{2}(L/E)] +2\sum_{i>j} \Im(\mathbf{U}) \sin[2.54\Delta m_{ij}^{2}(L/E)] - CP \text{ violation}$ 

### **Solar Neutrino Oscillations**



$$\mathbf{v}_e \rightarrow \mathbf{v}_x$$

#### $\Delta m^2 \sim 8 \times 10^{-5} \, eV^2$ , $\sin^2 2 \vartheta \sim 0.3$

Solar neutrinos, confirmed by reactor antineutrinos

### **Atmospheric Neutrino Oscillations**



 $\mathbf{v}_e \rightarrow \mathbf{v}_{\chi} \qquad \mathbf{v}_{\mu} \rightarrow \mathbf{v}_{\chi}$ 

 Δm<sup>2</sup>~8x10<sup>-5</sup> eV<sup>2</sup>, sin<sup>2</sup> 2θ~0.3
 Δm<sup>2</sup>~2.5x10<sup>-3</sup> eV<sup>2</sup>, sin<sup>2</sup> 2θ~1.0
 Atmospheric ν, confirmed by accelerator ν Aspen Conference on Particle Physics February 2005

### **LSND** Oscillations:

 $v_e \rightarrow v_x$ 



 $\mathbf{v}_{\mu} 
ightarrow \mathbf{v}_{\chi}$ 

 $\mathcal{V}_{e}$ 

 $\bar{\nu}_{\mu}$  $\Delta m^2 \sim 8 \times 10^{-5} \, eV^2$ ,  $\sin^2 2 \vartheta \sim 0.3$ ×  $\Delta m^2 \sim 2.5 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2 \vartheta \sim 1.0$  $\Delta m^2 \sim 10^{-1} - 10^1 \text{ eV}^2$ ,  $\sin^2 2\theta \sim 10^{-4} - 10^{-2}$ LSND Aspen Conference on Particle Physics February 2005

### The LSND Result



Search for excess  $\bar{v}_e$  in  $\bar{v}_\mu$  beam • Stopped pion beam produces pure  $\bar{v}_\mu$   $\pi^+ \rightarrow \mu^+ v_\mu$   $\mu^+ \rightarrow e^+ v_e \bar{v}_\mu$ • Excess of 87.9 ±22.4 ±6.0 events • Oscillation probability:  $(0.264\pm 0.067\pm 0.047)\%$ 

#### A challenge to the Standard Model:

Δm<sup>2</sup>~10<sup>-5</sup>, 10<sup>-3</sup>, 10<sup>-1</sup>eV<sup>2</sup> cannot result from three neutrinos
 Cannot be explained by additional light active neutrinos
 Fundamentally new physics (additional particles, broken symmetries) needed to explain all three modes.

# MiniBooNE

### Confirm/refute LSND evidence for $\bar{v}_{\mu} \rightarrow \bar{v}_{e}$ oscillations

### LSND vs. MiniBooNE LSND MiniBooNE

Neutrino Energy	~40 MeV	~800 MeV
Baseline	30 meters	540 meters
Signal Process	inverse β decay	v <sub>e</sub> CC quasi-elastic
Signal Identification	Double coincidence (e <sup>+</sup> , n capture)	Č ring, Sci. profile ν <sub>e</sub> from μ/Κ NC π <sup>0</sup>
Backgrounds	π⁻ wrong sign decay	
S/B Yield	~88/30	~300/800

• Sensitive to same parameters with different method Aspen Conference on Particle Physics February 2005

## Overview of MiniBooNE

- Produce a pure beam of  $v_{\mu}$ 
  - proton interactions on Be produce  $\pi^+$
  - $\pi^+ \rightarrow \mu^+ \nu_{\mu}$  in decay region



- 8 GeV Protons delivered by FNAL Booster
- Look for  $v_e$  interactions in the detector



### Focussing the Beam

- Electromagnetic focussing horn
  - 170 kA pulse New horn installed an
  - focuses positive secondaries

New horn installed and working well

- x5 enhancement in neutrino flux
  - Polarity can be reversed to focus negative secondaries



## "Intrinsic" Background (II)

- Sources of  $v_e$  not due to neutrino oscillations
  - $\mu^+$  produced from  $\pi^+$  decay can also decay
  - Kaons are produced in p-Be interactions and decay via K<sub>e3</sub>
- Source of irreducible v<sub>e</sub> background



### The Neutrino Beam



### • Primary p-Be interaction:

- $\pi^{\pm}$  from global fit to available data
- K<sup>+</sup> from global fit
- K<sup>0</sup> scaled according to GFLUKA
- Use existing data including E910
- High purity  $v_{\mu}$  beam
  - ~0.5%  $v_e$  contamination from:
  - Kaons produced at target (K<sub>e3</sub>)
  - μ decays from pion decay
- 540 m baseline to detector

### HARP (Hadron Production)



#### **Dedicated Measurement:**

- 8 GeV protons on Be
  - Replica targets
     0.1, 0.5 and 1 interaction length

• Tracking (TPC, Drift Chambers) Particle ID (TOF and Cherenkov)

Precision Pion and Kaon production measurement
 Spectrum and rate of incident neutrino flux
 Backgrounds from intrinsic v<sub>e</sub> (Kaon decay)
 First measurements of pion production released
 Aspen Conference on Particle Physics February 2005

## Little Muon Counter (LMC)



**Decay Region Monitor:** 

- Wide angle (7°), high p (2 GeV/c) muons
- Kaon decays in the decay pipe.

#### Detector:

- Collimator to select angle range
- Fiber tracker/magnet
- Range stack

### Detector installed: Analysis in progress



### The MiniBooNE Detector

- 800 ton mineral oil target
- 610 cm radius
- Optical barrier at 575 cm
  - Inner "tank" volume
     1280 photomultipliers
  - Outer "veto" region
     240 photomultipliers



Detect neutrino interactions via Č and scintillation light



## Particle Identification:

Cherenkov radiation:

- Charged particles with produce cone of radiation
- Minimum ionizing particles (muons) sharp-edged rings
- Electrons (photons) scatter, shower, convert, etc.
   →more diffuse rings
- Multiple particles: reconstruct by identifying rings



### Reducible Backgrounds

### • Signal Process: $v_e$ CCQE

- $v_e + n \rightarrow p + e$
- proton typically under threshold
- single electron-like ring
- Backgrounds from high energy photons
  - NC  $\pi^0$  production:  $\nu + (n/p) \rightarrow \nu + \pi^0 + (n/p) = \pi^0 \rightarrow \gamma\gamma$
  - NC radiative  $\Delta$  decays:  $\mathbf{v} + (n/p) \rightarrow \mathbf{v} + \Delta$  $\Delta \rightarrow (n/p) + \gamma$
- Background rejection by topology of PMT hits

Highly sensitive to photon propagation in mineral oil

## Primary light production



- Č light production
  - Occurs when  $n\beta > 1$  (n~1.47)
  - Emitted in cone
  - $1/\lambda^2$  wavelength distribution
- Scintillation light
  - Emission from molecular excitations from ionization
  - Emits isotropically
  - Several lifetime, emission modes

•  $\lambda = 270-340$  nm

Optical properties of light change dramatically over wavelength range Aspen Conference on Particle Physics February 2005

### Absorption



# Photon disappears: Thermally dissipated

### Rayleigh/Raman Scattering



### Rayleigh Scattering:

- Density pertubations
- Prompt, no  $\lambda$  shift
- Raman scattering
  - Excitation of vibrational states
  - Prompt with  $\lambda$  shift

#### Dominant process at $\lambda$ >350 nm

### Fluorescence



### Excite molecular states

- Emission at different wavelength
- Decay lifetime
- Multiple components
  - Different lifetimes (0.35-33 ns)
  - Different emission (270-340 nm)
- Stokes Shift:
  - UV photons red-shifted to visual

#### Dominant process in UV region (<300 nm)

## Summary of Processes



Measurements of

- Index of refraction
- Raman/Rayleigh
- Fluorescence
  - time-resolved
  - steady state
- Overall rate (extinction)
- Scintillation

Complement "test beam" measurements with *in-situ* calibrations Recent "push" propagating through analyses Aspen Conference on Particle Physics February 2005

### **Calibration Systems**



### Tracker/Cube System

- Scintillator hodoscope
- Seven scintillator cubes at various depths (15 cm-6 m)
   Muons with well-known pathlength

Laser Flask System: 397 and 438 nm pulsed lasers 4 Ludox flasks scatter light 1 bare fiber (collimated light)



## **Calibration Systems**



Tracker/Cube reconstructed muons

- Energy estimate from pathlength and dE/dx
- Compare with reconstructed energy

#### Michel electrons:

Decay of stopped muons

Well-defined energy spectrum

Reconstructed energy compared with theory and resolution model



### **Space/Time Distributions**



#### Laser data:

- Scattering/absorption from time profile Tracker/Cube Muons:
  - Scintillation/Fluorescence from time and angular distribution



## Neutrino Interactions at O(1GeV)



v<sub>u</sub> CCQE events



Selected based on: **Ring profile** Time profile of hits 80% purity Neutrino energy based on Energy, angle of muon Two body kinematics 28K events selected

## NC $\pi^0$ Events:





• Two ring fit for each event

- Č/Sci light from each
- Direction
- Mean shower point

Kinematic reconstruction of  $\pi^0 \rightarrow \gamma \gamma$  decay Aspen Conference on Particle Physics February 2005

# NC $\pi^0$ Kinematic Distributions



#### $\pi^0$ misidentification driven by

- Collimation of photons, energy asymmetry of photons
- Momentum, CM decay axis

### Expected Signal/Background

#### Process All Events After Selection

$v_{\mu}$ CC quasi-elastic	553,000	8
$\nu_{\mu} NC \pi^{0}$	110,000	290
Radiative $\Delta$ decay	1,080	80
Intrinsic v <sub>e</sub>	2,500	350
Oscillation Signal	1,500	300

Signal/Background 300/780=0.38For  $10^{21}$  protons-on-target NC  $\pi^0$  is dominant reducible background Aspen Conference on Particle Physics February 2005

### **Expected Sensitivity**





Energy distribution fit to extract signal, background yield

### Looking ahead: FY 2006

- MiniBooNE approved for FY06 running
- FY06 running may be in antineutrino mode: Studies of O(1 GeV)  $\bar{v}_{\mu}$  interactions
  - Challenge: wrong-sign  $(v_u)$  contamination (30%)
    - Angular distributions
    - Muon lifetime ( $\mu^+$  vs.  $\mu^-$  with capture)
    - CC  $\pi^+$  events (from  $v_{\mu}$  events only)
- Prepare for  $\bar{\mathbf{v}}_{\mu} \rightarrow \bar{\mathbf{v}}_{e}$  oscillation search

#### Summary and Outlook MiniBooNE: Confirm/refute LSND evidence for neutrino oscillations ! Confirmation has dramatic implications for neutrino physics Accumulated $4 \times 10^{20}$ pot (400K neutrino interactions) Detector/reconstruction functioning well Beamline functioning well (>100 million horn pulses with 1st horn, new horn installed) **Current Activities** Systematic studies: Bring offline measurements and in-situ into agreement For both beam and detector

Accumulating data towards 10<sup>21</sup> pot goal

### Neutrinos in the Standard Model



- Lepton sector has
  - charged leptons
  - neutrinos
- Neutrinos identified by flavor



 $v_l$  produces lepton *l* (e, μ, τ) in the weak charged current interaction Aspen Conference on Particle Physics February 2005

## "Intrinsic" Background (I)

- $\mu^+$  produced from  $\pi^+$  decay can also decay
  - Produce v<sub>e</sub> in detector not due to oscillations
  - Irreducible "intrinsic" v<sub>e</sub> background
- μ<sup>+</sup> intrinsic background ∝ decay region length



### "MisID" Background

- Some  $v_u$  interactions produce high energy  $\gamma$ s
  - $\pi^0$  production
  - Radiative Delta decays  $(\Delta \rightarrow N\gamma)$
- γ conversions produces e<sup>+</sup>e<sup>-</sup> pairs
  - Reducible by analyzing topology of event



### **Cosmic Rays**

Shallow overburden reduces rate to 10 kHz

- Use combination of active veto, beam timing
- e from stopped µ provide valuable calibration



### Fluorescence







#### 1 cm cell:

- Time-resolved (285, 300 nm excitation)
- Steady-state excitation/emission matrix

### **Cosmic Rays**



Rate and angular distribution of Rayleigh scattering

# $v_{\mu}$ CCQE Kinematics

