

Kaluza-Klein Dark Matter: a review

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Based on work with

-Tim Tait: hep-ph/0206071 (NPB)
 hep-ph/0209262 (New J.Phys.)

-Bertone & Sigl: hep-ph/0211342 (PRD)

-Kaustubh Agashe: hep-ph/0403143 (PRL)
 hep-ph/0411254 (JCAP)

-Dan Hooper: to appear

ADD models

only gravity
in bulk

$$R \sim \text{meV}^{-1} \text{ (flat)}$$

- radion dark matter
 $m \sim \text{meV}$; (fine-tuned)
- branon dark matter
(fine-tuned)

$$\text{TeV}^{-1} \text{ X-dims}$$

gauge bosons
in bulk

all SM fields
in bulk

"Universal" X-dims

$$R \sim \text{TeV}^{-1} \text{ (flat)}$$

- radion dark matter
 $m \sim \text{meV}$; (fine-tuned)

- KK dark matter
WIMP!

Warped geometries

(Randall-Sundrum)

(AdS)

if GUT in bulk

- radion unstable

- KK dark matter
WIMP!

$$R \sim M_{\text{Pl}}^{-1}$$

but

$$M_{\text{KK}} \sim \text{TeV}$$

Hierarchy pb solved

WIMP KK dark matter

So far, two working models :

✓ Universal Extra Dimensions (UED)

WIMP = Lightest KK particle (LKP)
stability symmetry = KK parity

✓ Warped GUTs

WIMP = Lightest Z_3 charged particle (LZP)
stability symmetry = Z_3 symmetry

+ a potential link between the LZP and baryogenesis...

Literature on KK dark matter: the complete list

Kolb & Slansky '84

Thought about it, but in 1984 $R^{-1} \sim \text{TeV}$ was inconceivable...

Dienes, Dudas & Gherghetta '99

Mohapatra & Perez-Lorenzana '02

mentionned the idea in passing

Servant & Tait '02

Cheng, Feng & Matchev '02

Servant & Tait '02

Majumdar '02

Hooper & Kribs '02

Bertone, Servant & Sigl '02

Hooper & Kribs '04

Bergstrom, Bringmann, Eriksson & Gustafsson '04

Baltz & Hooper '04

Bergstrom, Bringmann, Eriksson & Gustafsson II '04

Kakizaki, Matsumoto, Sato & Senami '05 A 2nd look at the relic density calculation

+ superWimp KK graviton papers

Detailed relic density calculation

Direct and indirect detection

Direct detection

Direct detection

Prospects for neutrino telescopes

Indirect detection

Positron excess

Indirect detection

KK dark
matter in
UED

Agashe & Servant '04

Agashe & Servant II '04

Hooper & Servant, upcoming

Model building, relic density,
direct detection, collider signatures ...

Indirect detection

Warped
KK dark
matter

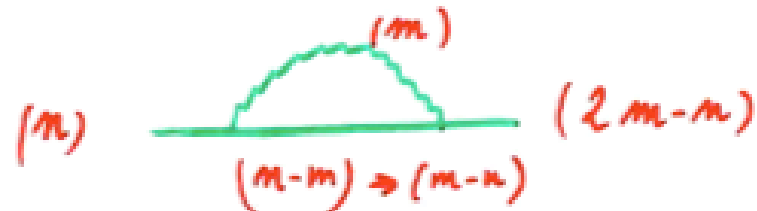
LKP dark matter in Universal Extra Dimensions

UED : ALL SM particles propagate into flat dimensions Appelquist, Cheng & Dobrescu '01

Conservation of momentum along extra dimension translates in 4D into conservation of KK number

by the orbifold we impose to recover a chiral theory

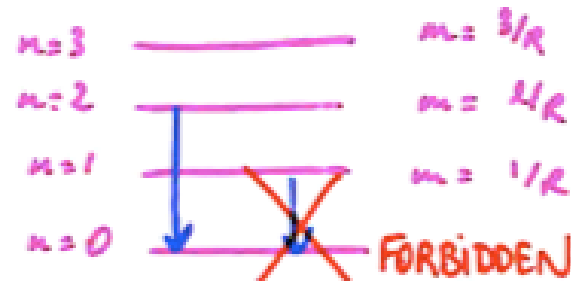
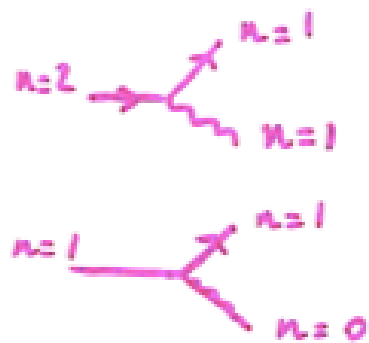
KK number is conserved at tree level but broken at loop level



n can change by an EVEN number only

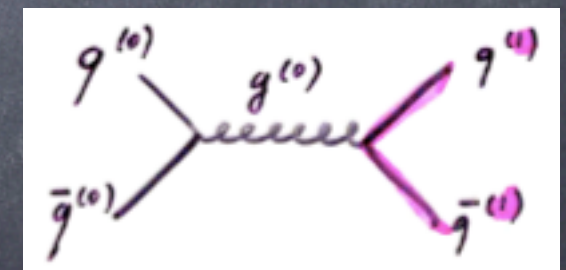
↳ KK PARITY : $(-1)^n$: preserved at all orders.

↳ Only interactions between an EVEN number of ODD KK modes



↳ the Lightest KK particle (LKP) is stable

Other consequence of KK parity:
Production of 1st KK modes
only by pairs:



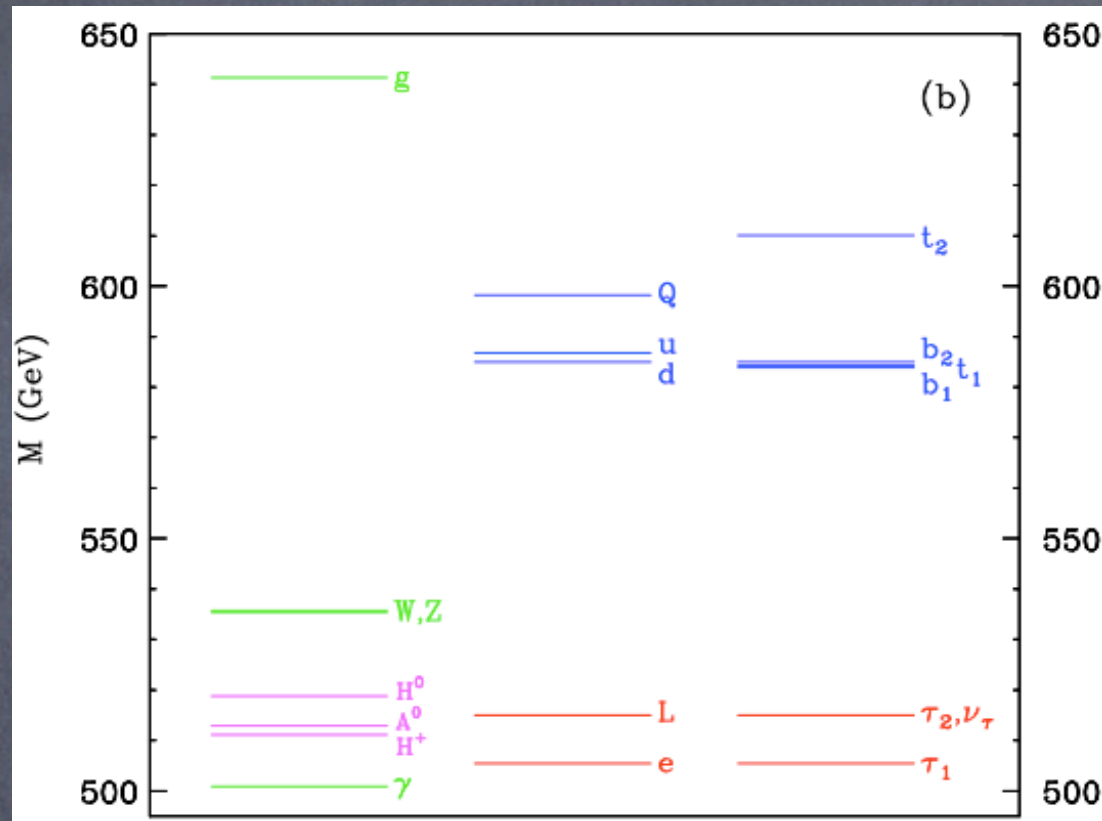
⇒ Weak bound on $1/R$

$$R^{-1} \gtrsim 300 \text{ GeV for } \delta=1$$

$$R^{-1} \gtrsim 500 \text{ GeV for } \delta=2$$

1-loop spectrum of 1st KK modes

Cheng, Matchev & Schmaltz'02



assuming: $1/R = 500 \text{ GeV}$, $\Lambda R = 20$, $m_h = 120 \text{ GeV}$
and vanishing boundary terms at the cutoff Λ

→ LKP: most likely a γ^1 (actually a B^1)

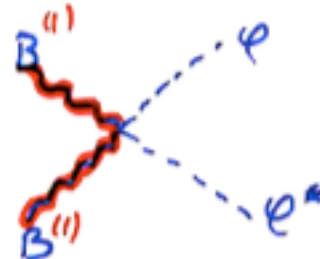
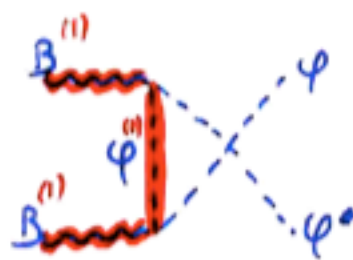
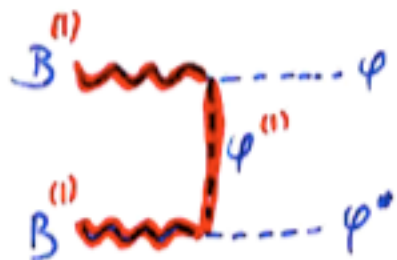
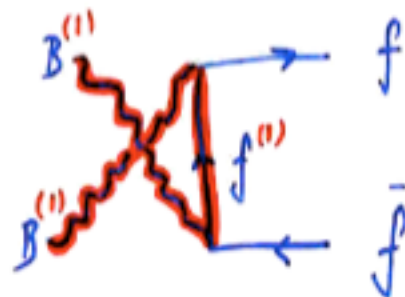
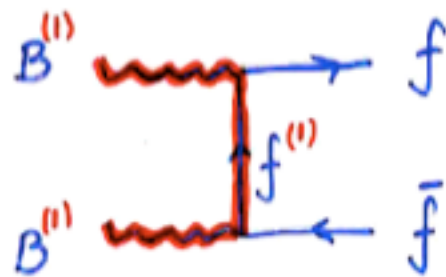
Another intriguing possibility: LKP=KK graviton (see S. Su's talk)

Relic density predictions

REMINDER : $\Omega h^2 \approx \frac{10^9}{m_{pl}} \frac{x_F}{\sqrt{g_*}} \frac{\text{GeV}^{-1}}{\langle \sigma_{eff} v \rangle}$

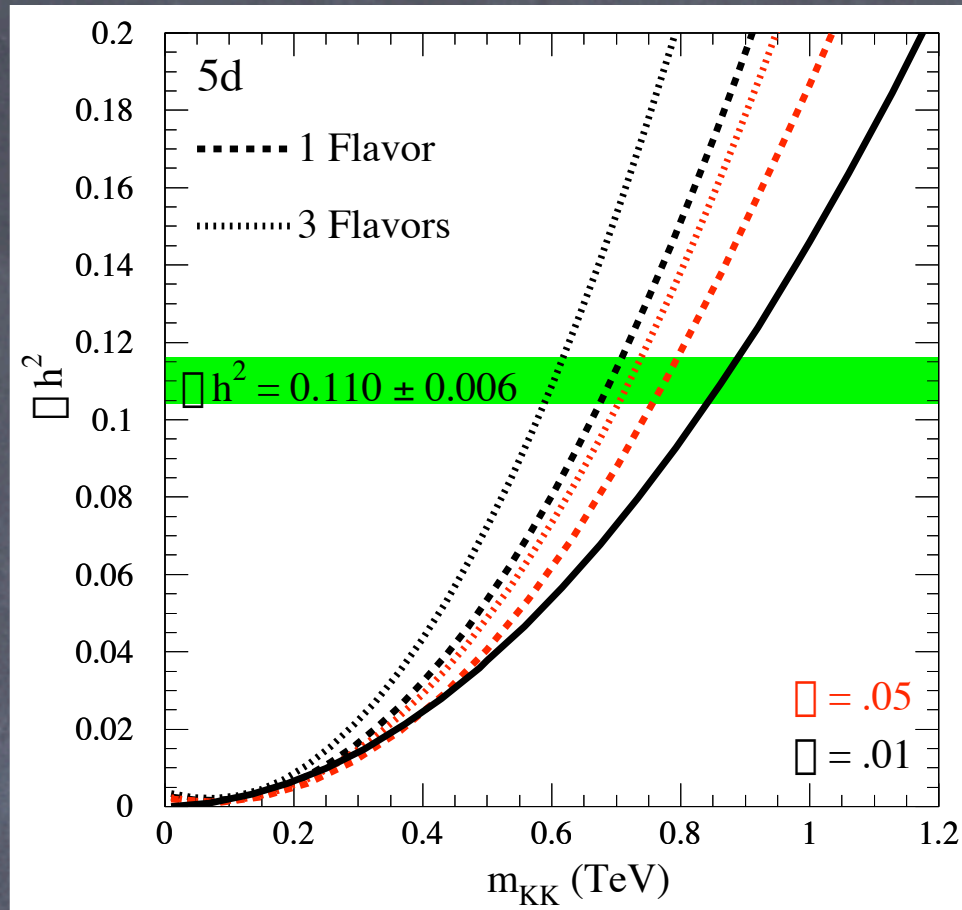
$\Omega h^2 \approx 0.11 \quad \leadsto \quad \langle \sigma_{eff} v \rangle \sim 1 \text{ pb} \quad (\alpha_F = 25-35)$

Annihilation cross sections

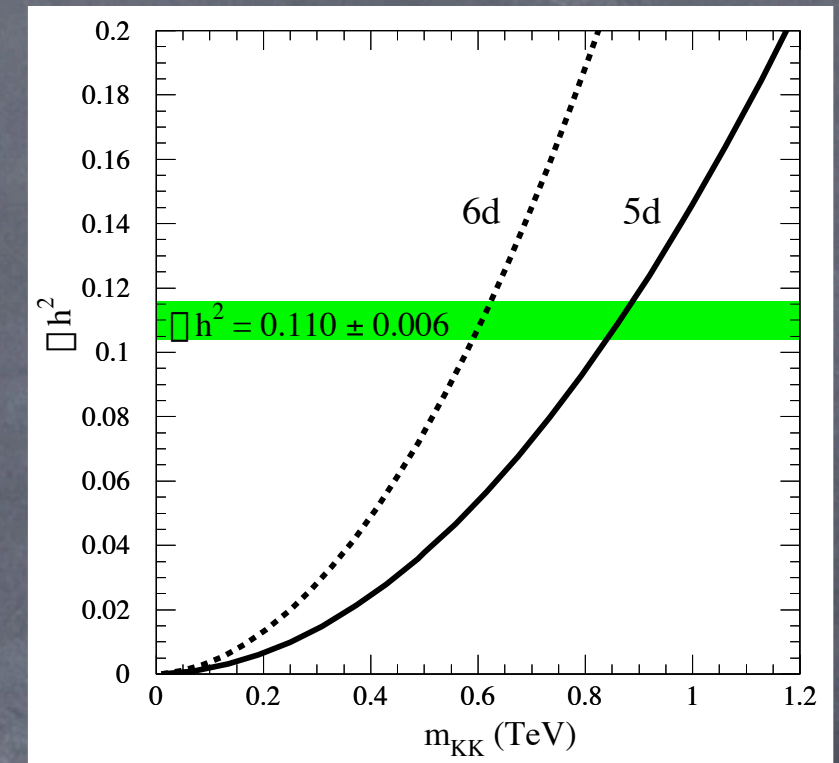


✓ Coannihilation effects

Servant-Tait



✓ Possible effect of additional dimension

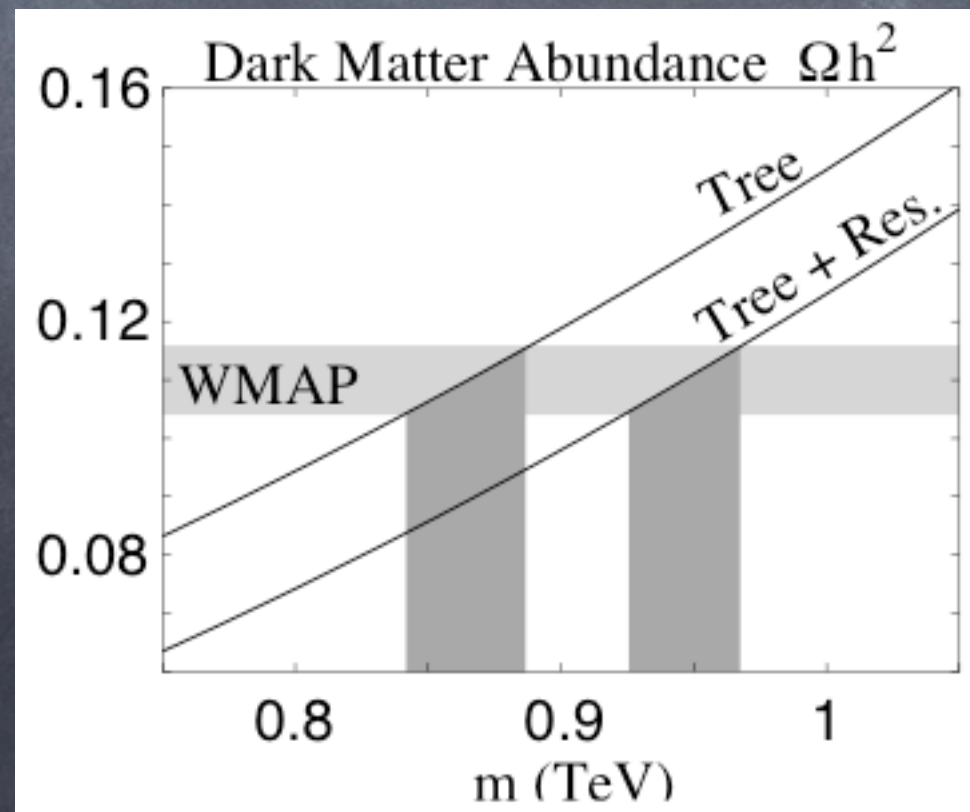
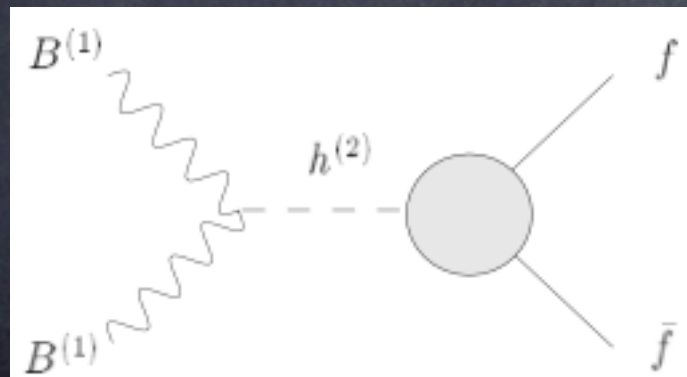


Servant-Tait

✓ Effect of 2nd KK modes

"natural KK resonance"

Kakizaki & al, hep-ph/0502059

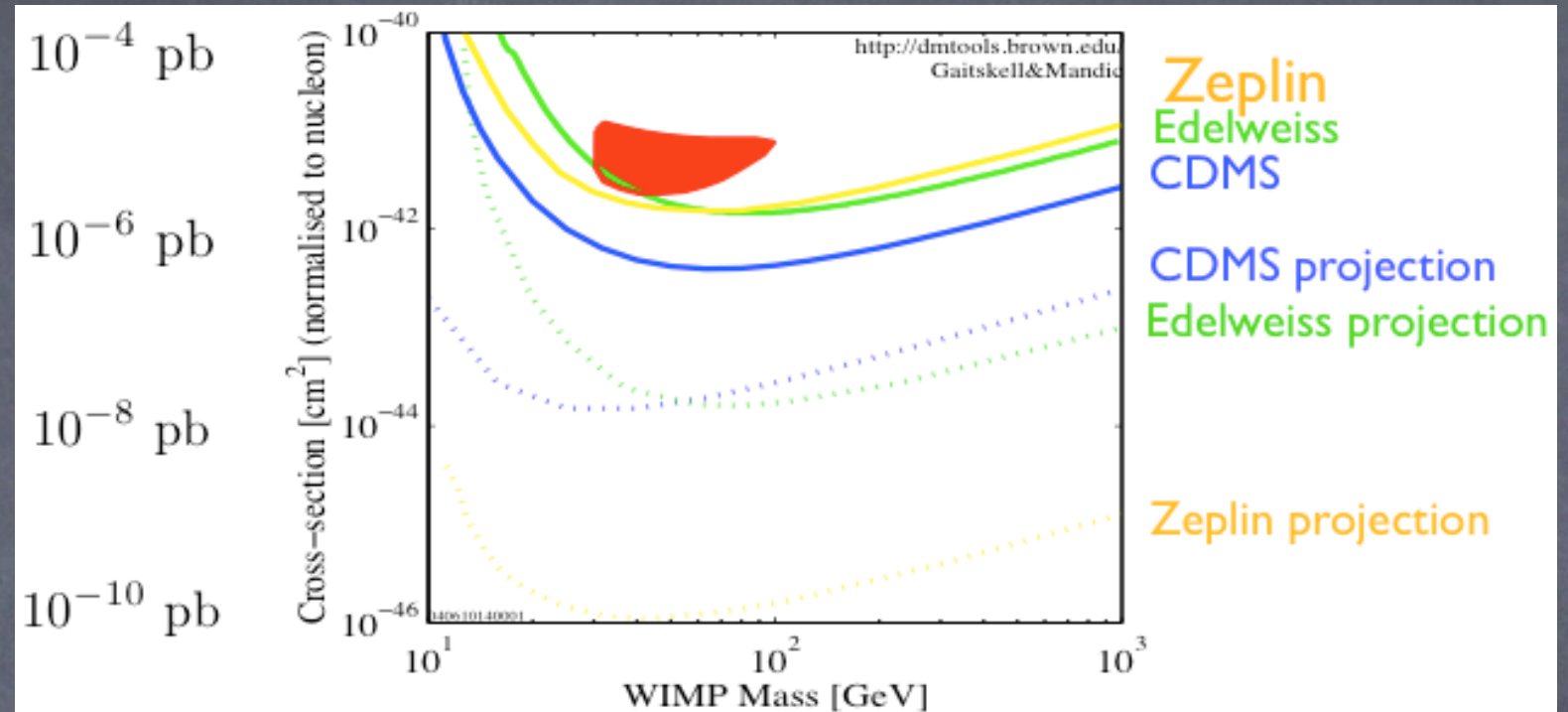


$$\delta \equiv (m_{h^{(2)}} - 2m)/2m$$

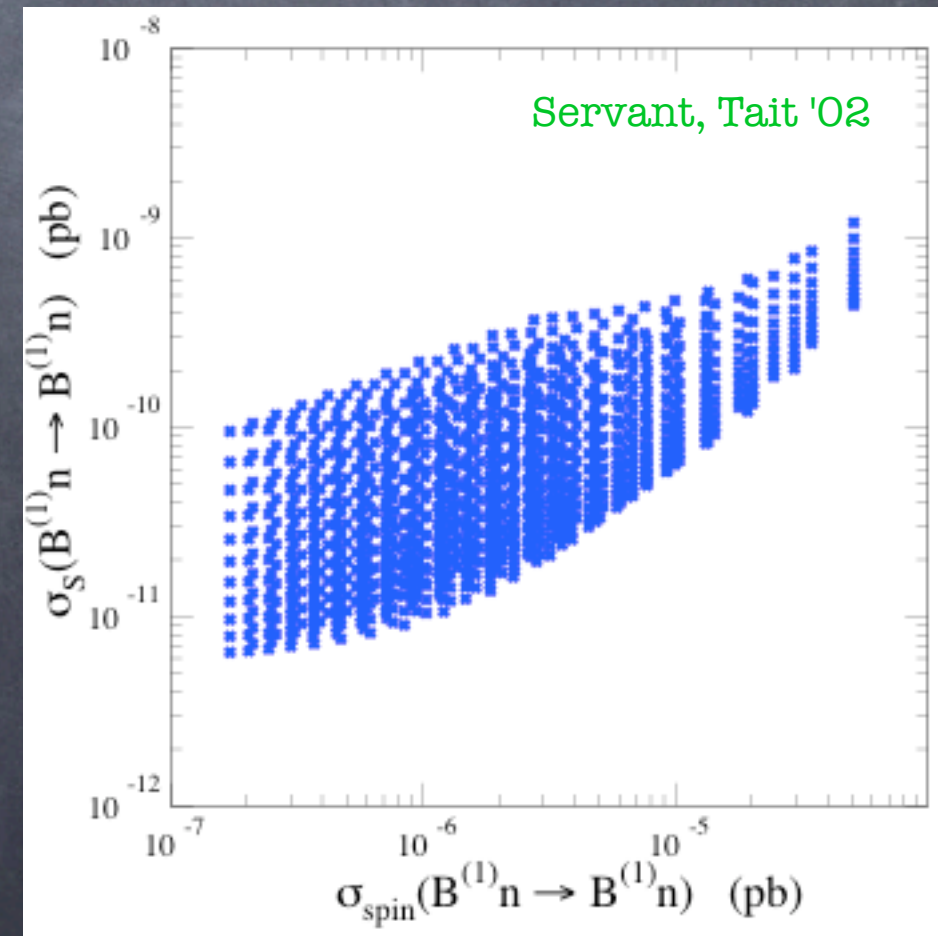
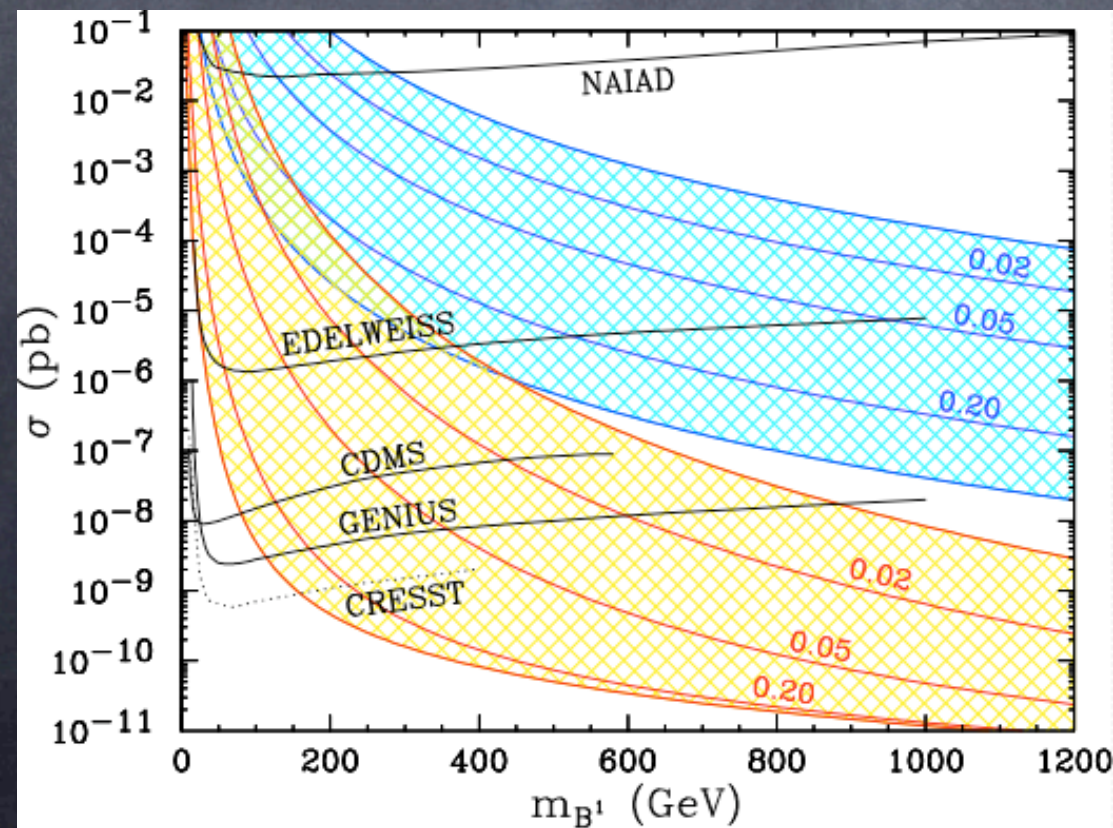
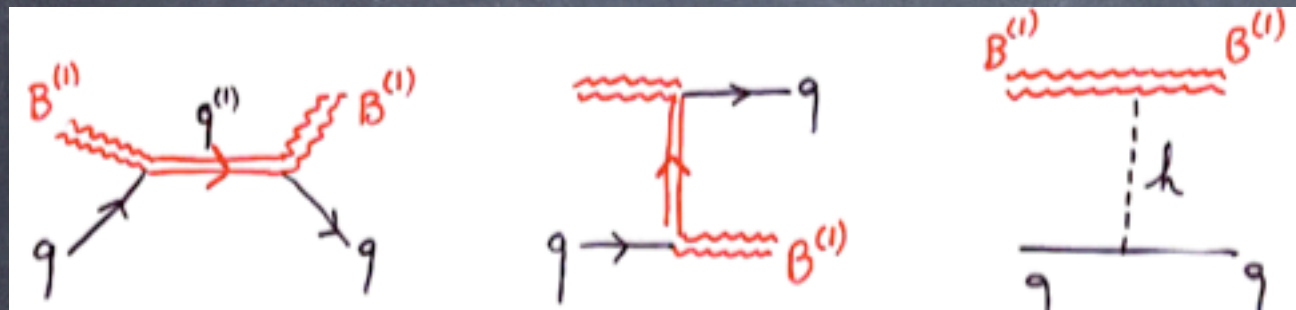
$$\delta \sim 0.01$$

Direct detection

Experimental limits :



LKP signal :

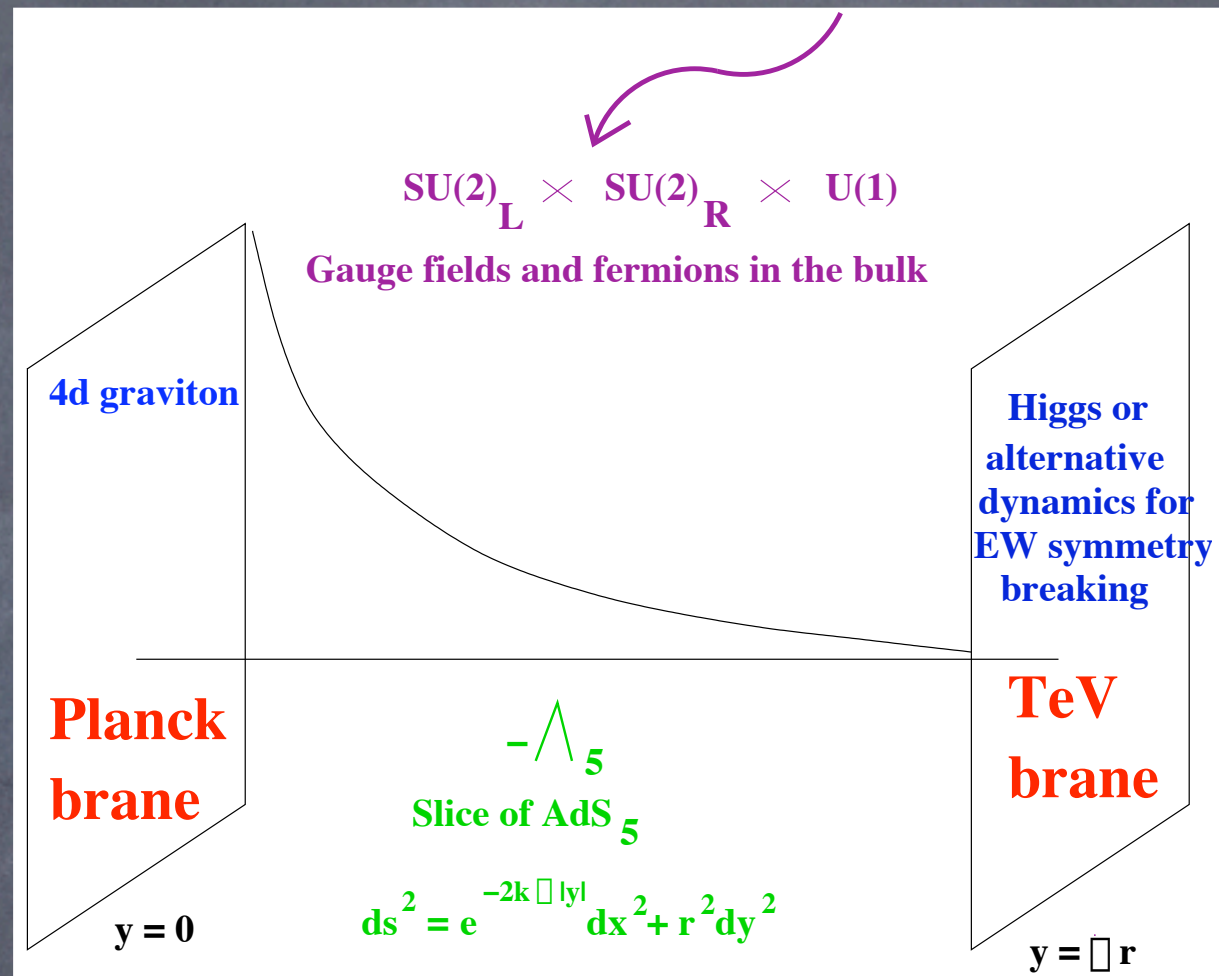


Cheng, Feng,
Matchev '02

Particle physics model building in warped space

2005 FAVOURITE SET-UP:

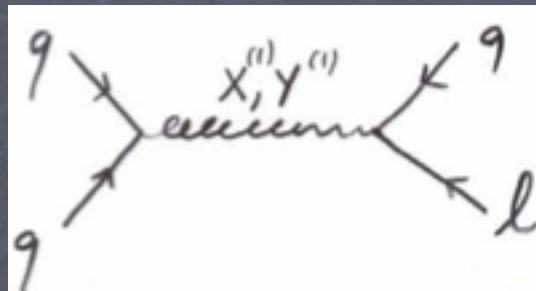
- ✓ hierarchy pb
- ✓ fermion masses
- ✓ High scale unification
- ✓ FRW cosmology



Now embed this into a GUT + solve proton stability

-
- ✓ Dark matter

In GUTs \Rightarrow



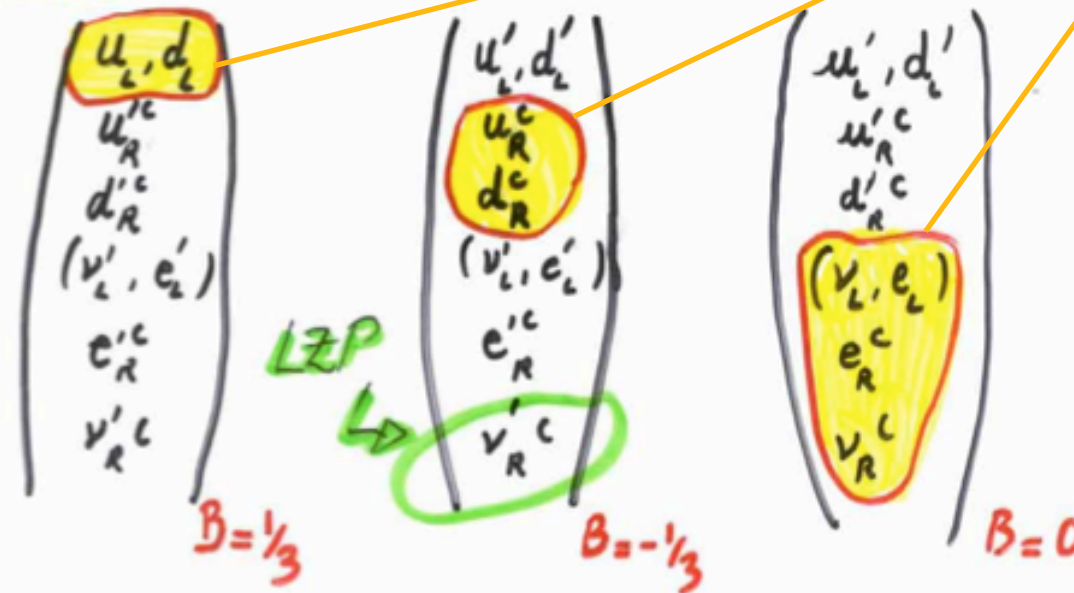
where $M_{X,Y} \sim \text{few TeV}$

\Rightarrow very fast proton decay

Solution: Break GUT by boundary conditions which split GUT multiplets

zero modes = SM fermions

$SO(10)$ example:



$$\hookrightarrow Z_3: \quad \Phi \rightarrow e^{2i\pi(B - (\frac{n_c - \bar{n}_c}{3}))} \Phi$$

B : baryon number ; $n_c (\bar{n}_c)$ = number of color (anticolor) indices of the particle.

SM particles are not charged under Z_3

\hookrightarrow LZP is stable

Mass spectrum of KK fermions

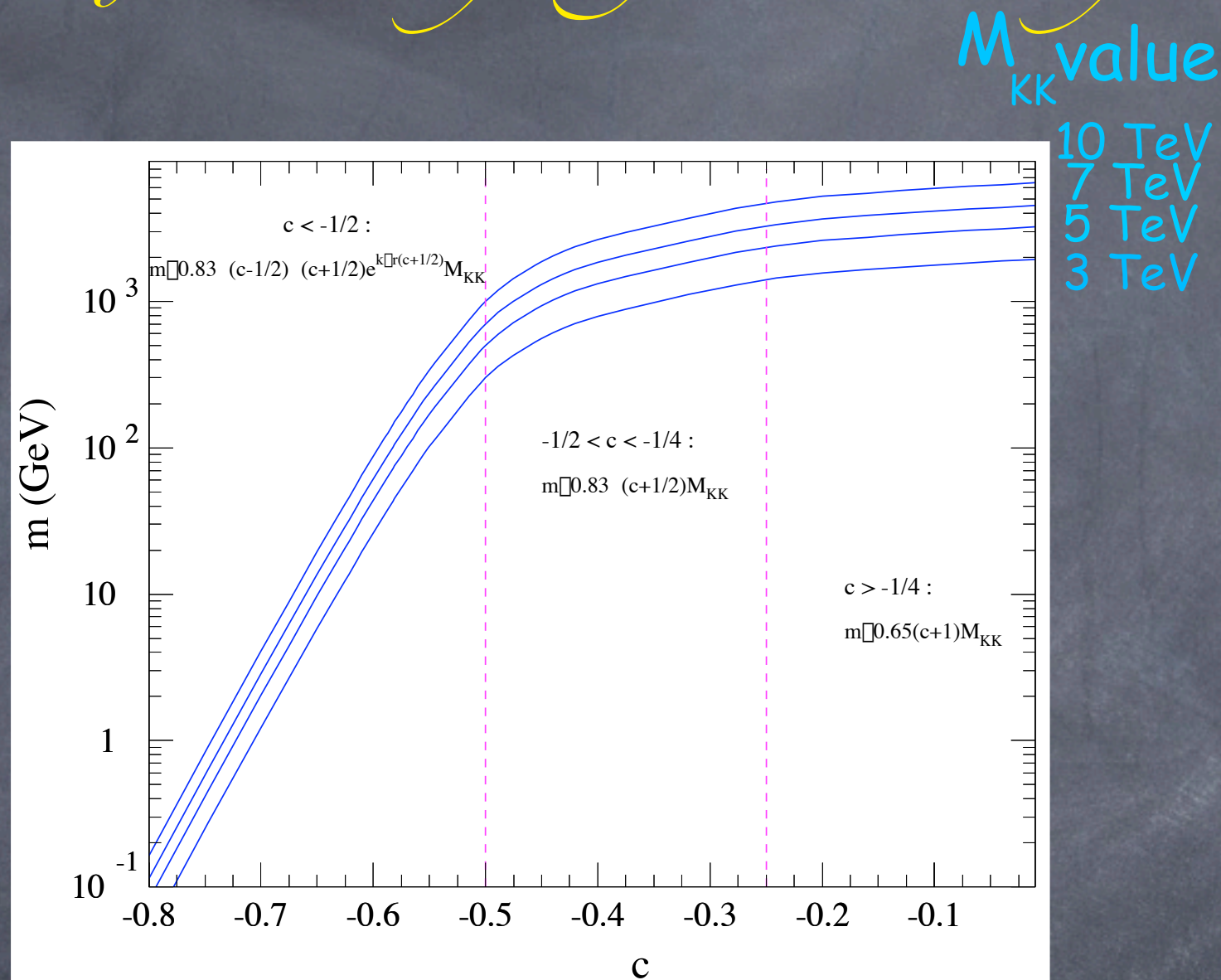
Depends on:

- ✓ type of boundary conditions on TeV and Planck branes
- ✓ c-parameter (=5D bulk mass)
(=localization of zero-mode wave function)

For certain type of boundary conditions on fermions,
there can be a hierarchy between the mass of KK
fermion and the mass of KK gauge bosons

⇒ Not a single KK scale

Mass spectrum of lightest KK fermion

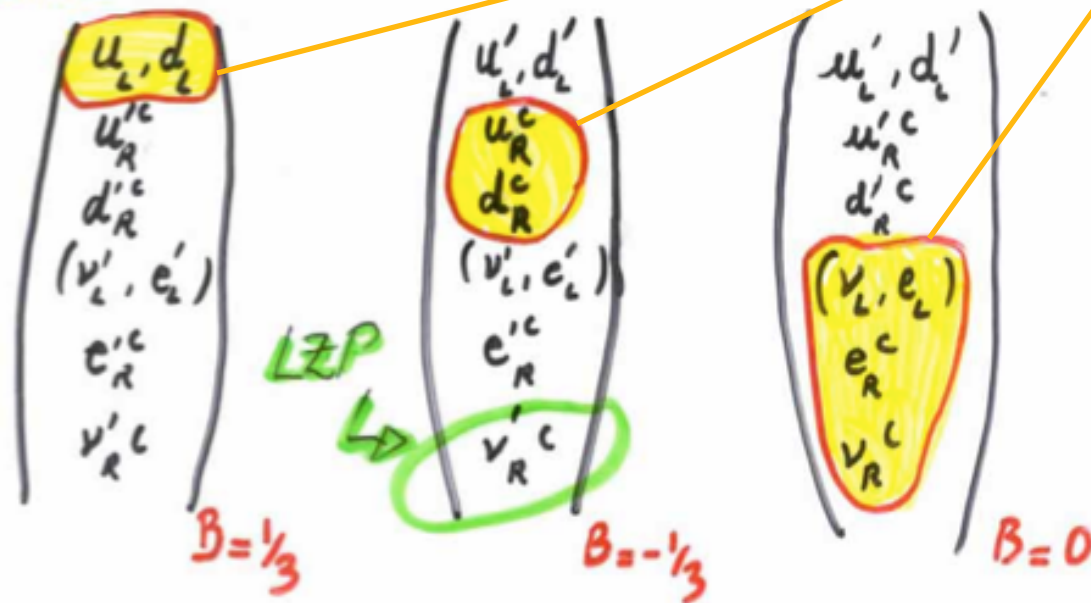


and smallest c : c of the top quark

⇒ LZF belongs to the multiplet containing SM top quark

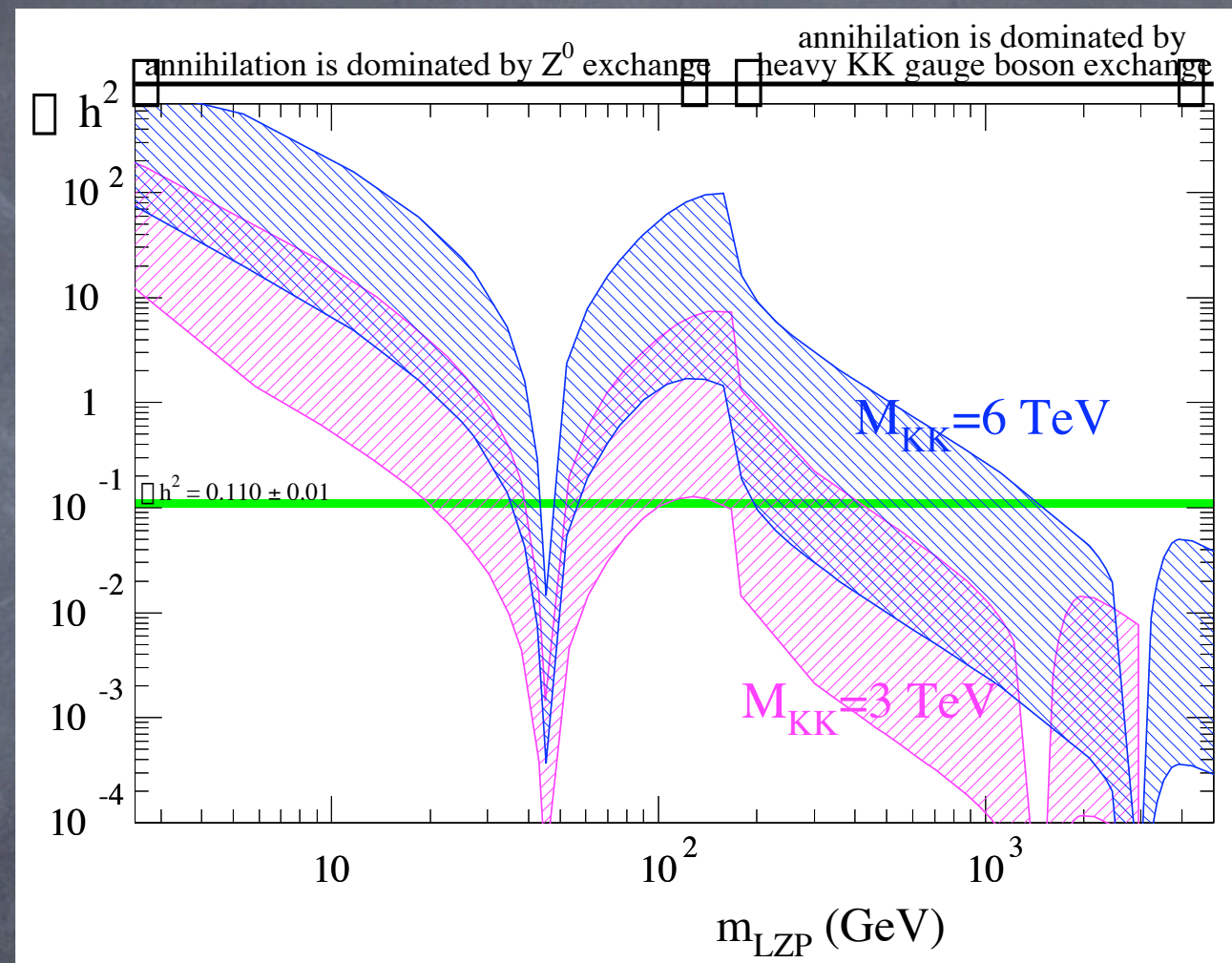
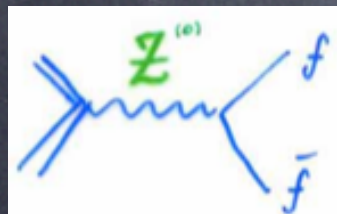
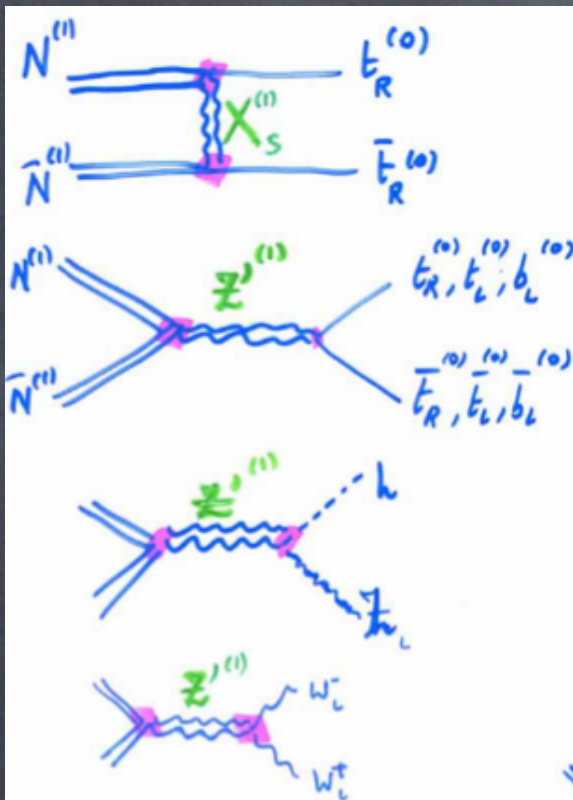
There exists a very light KK fermion as a consequence of the heaviness of the top quark

$SO(10)$ example:



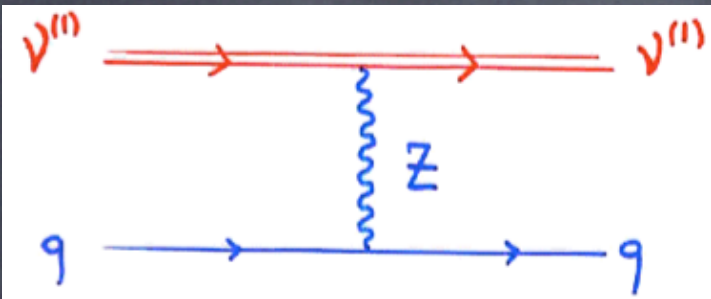
zero modes =
SM fermions

Relic density predictions



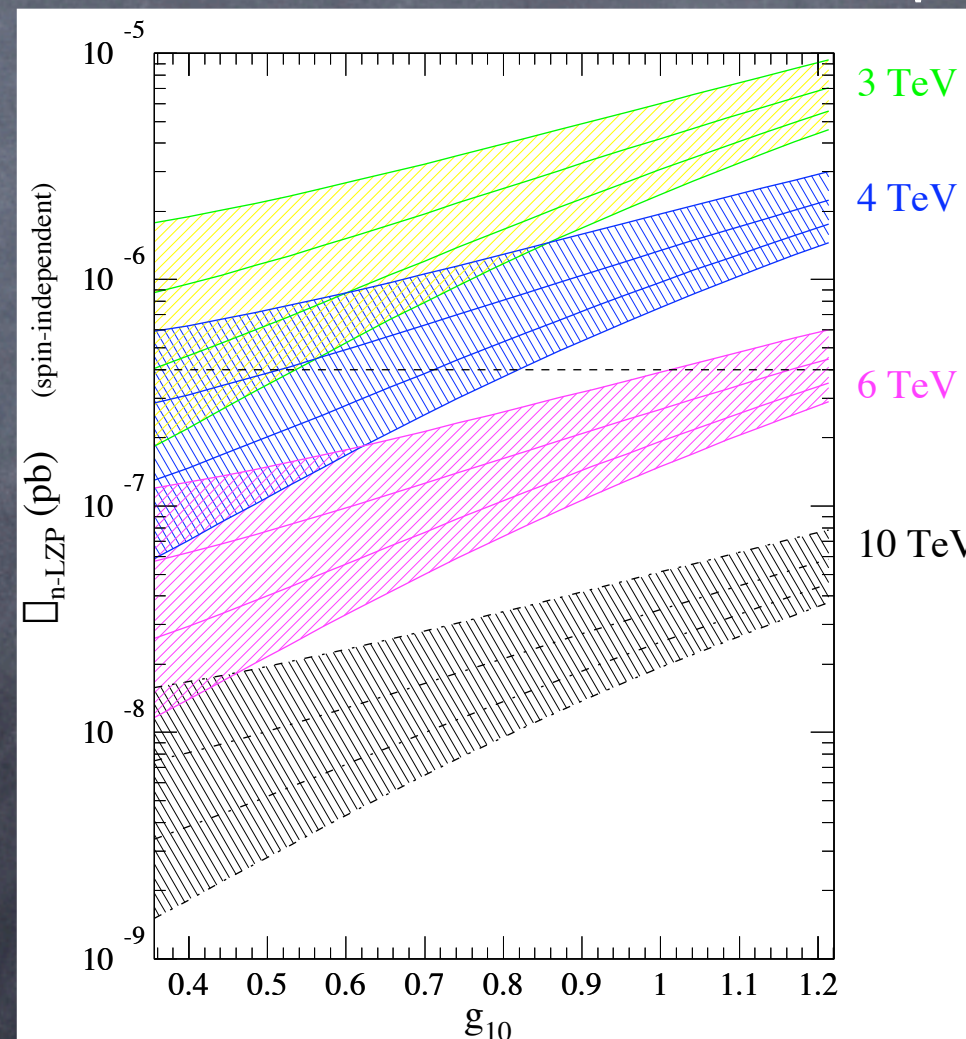
Agashe-Servant '04

Direct detection

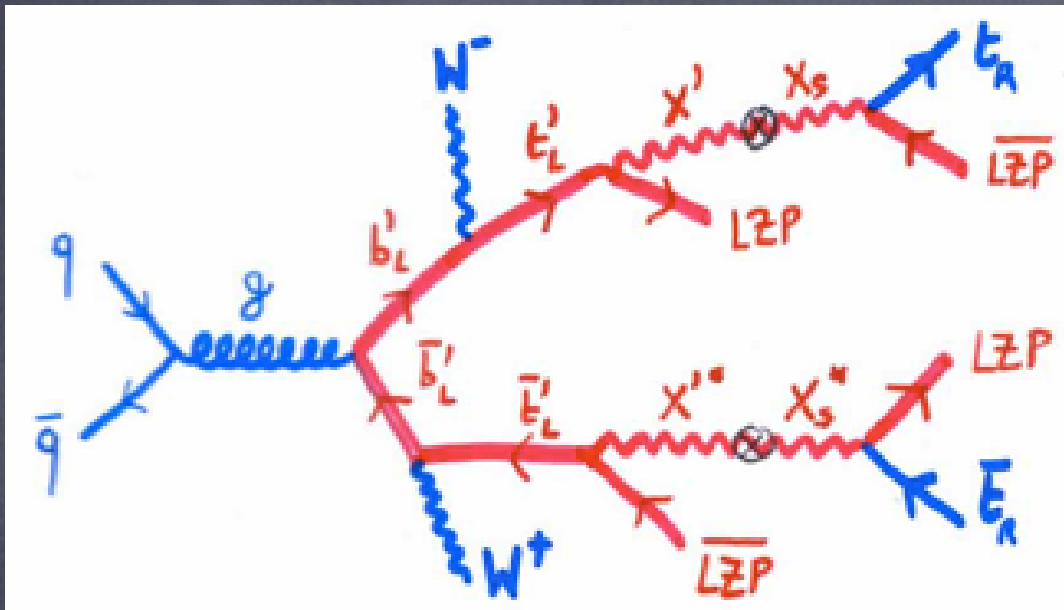


Wimp-nucleon elastic
scattering cross section
(spin-independent)

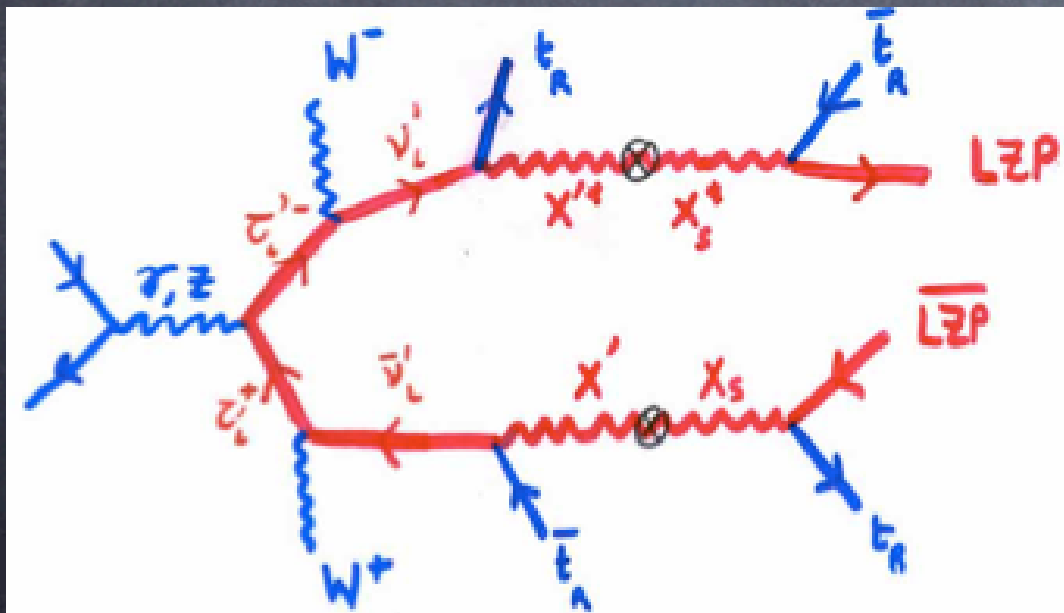
masses of KK
gauge bosons



Collider Signatures: examples



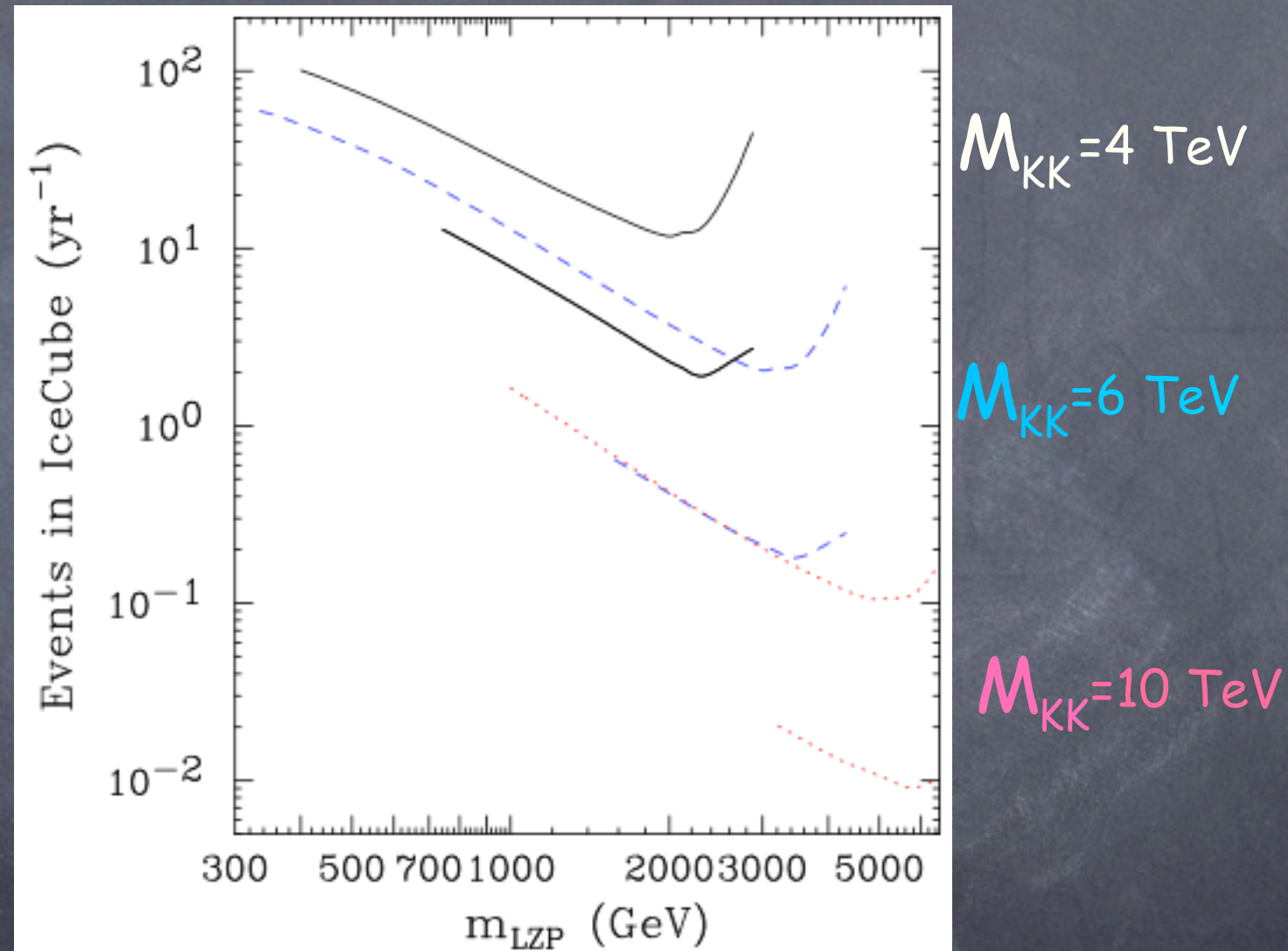
$$\Rightarrow 4 W + 2 b + \cancel{E}_T$$



$$\Rightarrow 6 W + 4 b + \cancel{E}_T$$

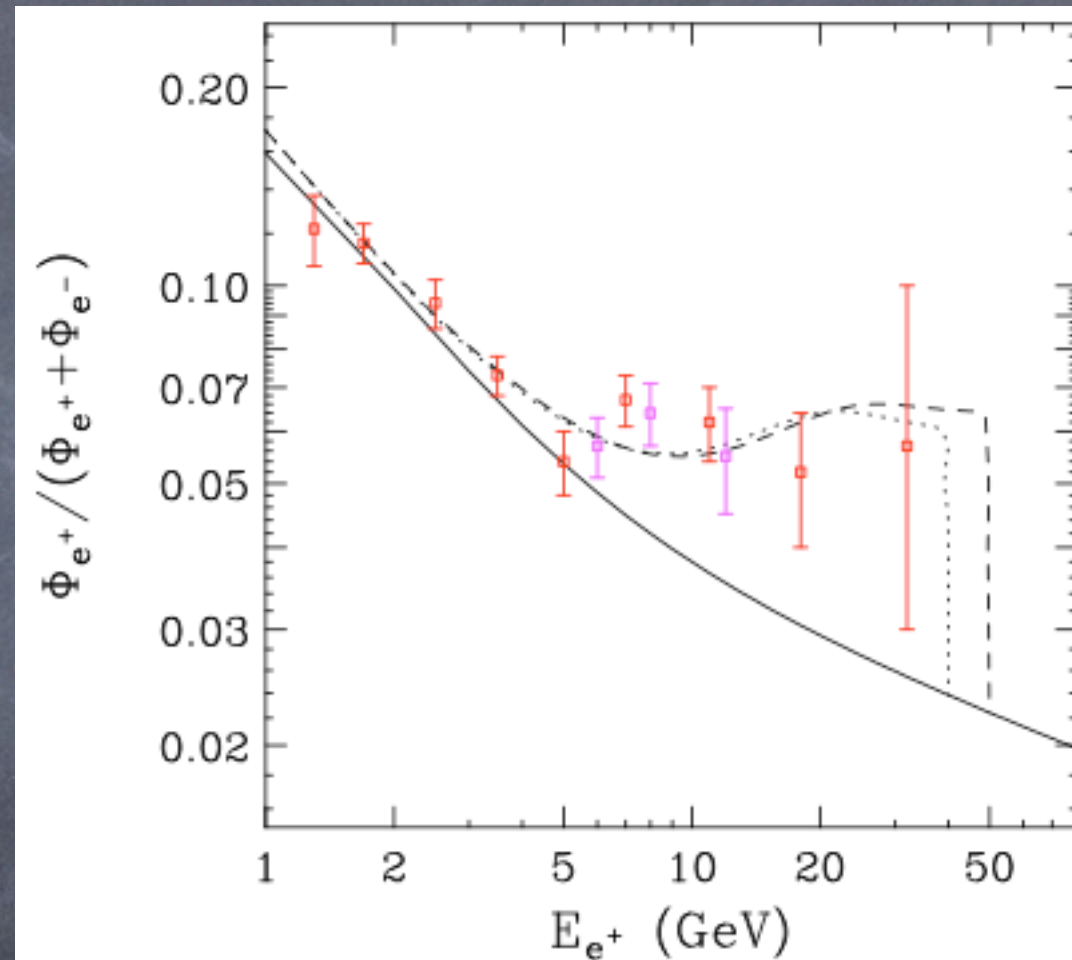
Indirect detection in neutrino telescopes

Large elastic scattering cross section: large capture rate in the Sun
Efficient production of neutrinos in annihilations



Hooper & Servant, in prep.

Cosmic positrons from LZP annihilations



Fit of the HEAT data
from LZP annihilations

..... 40 GeV LZP
----- 50 GeV LZP

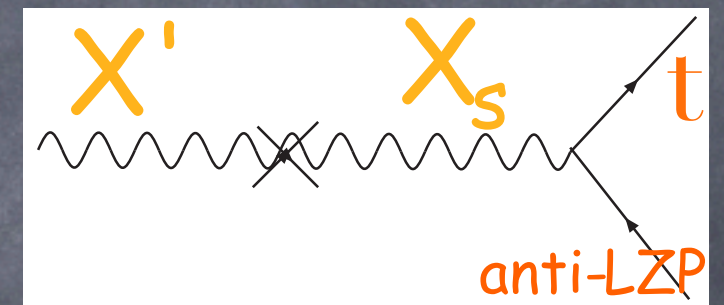
Hooper & Servant, in prep.

A natural framework to relate ^{baryons} and ^{dark matter}

Our dark matter candidate carries baryon number! ($B=1/3$)

$$B_{\text{UNIVERSE}} = 0 = \underbrace{B}_{\text{carried by baryons}} + \underbrace{(-B)}_{\text{carried by anti-LZP}}$$

Assume an asymmetry between t and \bar{t} is created via the out-of-equilibrium and CP-violating decay :




Baryon number conservation leads to:

$$3 (n_{\overline{\text{LZP}}} - n_{\text{LZP}}) = n_b - n_{\bar{b}}$$

Assuming efficient annihilation between LZP and $\overline{\text{LZP}}$, and b and \bar{b} :

$$\rho_{\text{DM}} = m_{\text{LZP}} n_{\text{LZP}} \approx 6 \rho_b \quad \longrightarrow \quad m_{\text{LZP}} \approx 18 \text{ GeV}$$

	LKP	LZP	LSP
<i>nature</i>	gauge boson	Dirac fermion	Majorana fermion
<i>symmetry</i>	KK parity $(-1)^n$	Z_3 $B - \frac{(n_c - \bar{n}_c)}{3}$	R-parity $(-1)^{3(B-L)+2S}$
		 related to proton stability	
<i>mass range</i>	~600-1000 GeV	20 GeV-few TeV	~50 GeV-1 TeV
<i>annihilation cross section</i>	s-wave	s-wave	helicity suppressed (p-wave)
<i>favourite detection</i>	<ul style="list-style-type: none"> ✓ LHC ✓ Indirect detection 	<ul style="list-style-type: none"> ✓ Direct detection! ✓ LHC! ✓ Indirect detection! entire parameter space is testable	<ul style="list-style-type: none"> ✓ LHC

To conclude

Abundance of experimental activity related to dark matter detection:

- Colliders
- Direct detection: CDMS, Edelweiss, Dama, Cresst, Zeplin, Xenon, Naiad ...
- Indirect detection:
 - Gamma-ray telescopes: Hess, Veritas, Glast, Magic
 - Neutrino telescopes: Amanda, IceCube, Antares
 - Cosmic positron experiments: HEAT, Pamela, AMS-2

⇒ It is timely to study the distinctive signatures expected in different dark matter scenarios.

LKPs, LZPs: viable alternatives to LSPs