

Distinguishing an MSSM Higgs from a SM Higgs at a Linear Collider

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Linear Collider Workshop 2000
Fermilab, October 24 - 28, 2000

- Distinguish MSSM Higgs from SM Higgs by measuring its BRs
- Radiative corrections to the MSSM Higgs sector affect the BRs:
→ dependence on SUSY parameters as well as M_A and $\tan \beta$
- How well can a LC do in various scenarios

Talk available at <http://www-theory.fnal.gov/people/logan/lcws.ps.gz>

Two sources of contributions to h^0 BRs:

- Renormalization of Higgs mass matrix \rightarrow mixing angle α :

Radiative corrections well known up to two loops
Okada, Yamaguchi, Yanagida; Ellis, Ridolfi, Zwirner; Li, Sher; Barbieri, Frigeni; Drees, Nojiri; Casas, Espinosa, Quiros, Riotto; Brignole, Ellis, Ridolfi, Zwirner; Zhang; Espinosa, Zhang; Gunion, Turski; Haber, Hempfling; Brignole; Diaz, Haber; Berger; Chankowski, Pokorski, Rosiek; Yamada; Dabelstein; Pierce, Bagger, Matchev, Zhang; Hempfling, Hoang; Heinemeyer, Hollik, Weiglein; Carena, Heinemeyer, Wagner, Weiglein; Barbieri, Frigeni, Caravaglios; Pierce, Papadopoulos, Johnson; Sasaki, Carena, Wagner; Hempfling; Kodaira, Yasui, Sasaki; Espinosa, Quiros; Carena, Espinosa, Quiros, Wagner; Carena, Quiros, Wagner; Carena, Haber, Heinemeyer, Hollik, Wagner, Weiglein

- Vertex corrections to fermion Yukawa couplings:
Corrections to the relation between the fermion masses and their Yukawa couplings. $\tan\beta$ enhanced for down-type fermions \rightarrow can be very significant.
Coarasa, Jimenez, Sola; Jimenez, Sola; Bartl, Eberl, Hikasa, Kon, Majerotto, Yamada; Hall, Rattazzi, Sarid; Hempfling; Carena, Olechowski, Pokorski, Wagner; Pierce, Bagger, Matchev, Zhang; Dabelstein; Haber, Herrero, H.L., Penaranda, Rigolin, Temes; Heinemeyer, Hollik, Weiglein

We use a version of HDECAY with these corrections included, put together by Steve Mrenna.

Radiatively corrected Higgs mass matrix and α

The tree-level Higgs mass matrix

$$\mathcal{M}_0^2 = \begin{pmatrix} m_A^2 s_\beta^2 + m_Z^2 c_\beta^2 & -(m_A^2 + m_Z^2) s_\beta c_\beta \\ -(m_A^2 + m_Z^2) s_\beta c_\beta & m_A^2 c_\beta^2 + m_Z^2 s_\beta^2 \end{pmatrix} = \begin{pmatrix} \mathcal{M}_{11}^2 & \mathcal{M}_{12}^2 \\ \mathcal{M}_{12}^2 & \mathcal{M}_{22}^2 \end{pmatrix}$$

is diagonalized by the mixing angle α :

$$s_\alpha c_\alpha = \frac{\mathcal{M}_{12}^2}{\sqrt{(\text{Tr} \mathcal{M}^2)^2 - 4 \det \mathcal{M}^2}}, \quad c_\alpha^2 - s_\alpha^2 = \frac{\mathcal{M}_{11}^2 - \mathcal{M}_{22}^2}{\sqrt{(\text{Tr} \mathcal{M}^2)^2 - 4 \det \mathcal{M}^2}}$$

If $\mathcal{M}_{12}^2 \rightarrow 0$, then either $\sin \alpha \rightarrow 0$ or $\cos \alpha \rightarrow 0$.

Radiatively corrected \mathcal{M}_{12}^2 : dominant one-loop top squark effects plus two-loop, leading log effects [Carena, Mrenna, Wagner] (approx.):

$$\mathcal{M}_{12}^2 \simeq -[m_A^2 + m_Z^2] s_\beta c_\beta - \left[\frac{3h_t^4 v^2}{48\pi^2 M_S} \mu x_t (6 - x_t a_t) s_\beta^2 - \frac{3h_t^2 m_Z^2}{32\pi^2} \mu x_t \right] \left[1 + \frac{9h_t^2 - 32g_s^2}{32\pi^2} \ln \left(\frac{M_S^2}{m_t^2} \right) \right]$$

($a_t = A_t/M_S$, $x_t = X_t/M_S$) $\rightarrow \alpha$ gets corrected. (Full result used in numerical calcs.)

m_t^4 -dependent rad. corrs. to \mathcal{M}_{12}^2 depend strongly on the sign of μX_t . (Note $A_t \simeq X_t$ for large $\tan \beta$ and μ not too big.)

$|a_t| \lesssim \sqrt{11/2} \rightarrow hbb$ coupling suppressed for $\mu A_t < 0$ and enhanced for $\mu A_t > 0$

$|a_t| \gtrsim \sqrt{11/2} \rightarrow$ vice versa

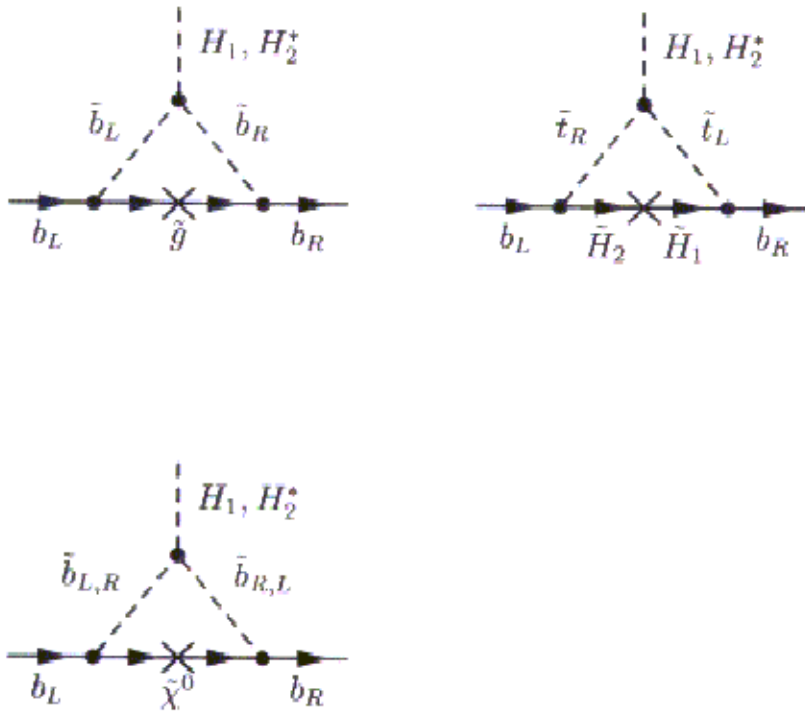
Loinaz, Wells; Kane, Kribs, Martin, Wells; Baer, Wells; Babu, Kolda

Correction to α affects BR(b) and BR(τ) in the same way: Ignoring the correction to the Yukawa couplings (next slide),

$h\bar{b}b$, $h\tau^+\tau^- \sim -(\sin \alpha / \cos \beta) \times$ SM coupling.

Δ_b and corrections to Yukawa couplings

In unbroken supersymmetry, b quarks couple only to Φ_d^0 . But SUSY is broken, and the b quarks receive a small coupling to Φ_u^0 from radiative corrections:



The effective Yukawa Lagrangian becomes (keeping only dimension ≤ 4 terms):

$$-\mathcal{L}_{\text{Yuk.}} \simeq h_b \Phi_d^0 \bar{b} b + (\Delta h_b) \Phi_u^0 \bar{b} b$$

Although Δh_b is loop suppressed compared to h_b , for sufficiently large $\tan \beta = v_u/v_d$ the contribution of both terms to the b mass can be comparable:

$$m_b = \frac{h_b v_d}{\sqrt{2}} + \frac{(\Delta h_b) v_u}{\sqrt{2}} = \frac{h_b v_d}{\sqrt{2}} (1 + \Delta_b)$$

Main contributions to Δ_b come from a bottom squark-gluino loop and a top squark-higgsino loop. Neglecting terms suppressed by powers of $1/M_{SUSY}$,

$$\Delta_b \simeq \frac{2\alpha_s}{3\pi} M_{\tilde{g}} \mu \tan \beta I(M_{\tilde{b}_1}, M_{\tilde{b}_2}, M_{\tilde{g}}) + \frac{Y_t}{4\pi} A_t \mu \tan \beta I(M_{\tilde{t}_1}, M_{\tilde{t}_2}, \mu)$$

$$(Y_t \equiv h_t^2/4\pi)$$

$$I(a, b, c) = \frac{a^2 b^2 \ln(a^2/b^2) + b^2 c^2 \ln(b^2/c^2) + c^2 a^2 \ln(c^2/a^2)}{(a^2 - b^2)(b^2 - c^2)(a^2 - c^2)} > 0$$

There are similar corrections to the relation between the τ mass and its Yukawa coupling:

$$m_\tau = \frac{h_\tau v_d}{\sqrt{2}} (1 + \Delta_\tau)$$

Δ_τ comes from a tau slepton-neutralino loop and a tau sneutrino-chargino loop; it is expected to be **much smaller** than Δ_b because it depends on the electroweak gauge couplings:

$$\Delta_\tau = \frac{\alpha_1}{4\pi} M_1 \mu \tan \beta I(M_{\tilde{\tau}_1}, M_{\tilde{\tau}_2}, M_1) + \frac{\alpha_2}{4\pi} M_2 \mu \tan \beta I(M_{\tilde{\nu}_\tau}, M_2, \mu)$$

→ we do not include it in our analysis.

The couplings to h^0 are modified by Δ_b, Δ_τ :

$$g_{hbb} = \frac{gm_b \sin \alpha}{2m_W \cos \beta} \frac{1}{1 + \Delta_b} \left[1 - \frac{\Delta_b}{\tan \alpha \tan \beta} \right]$$

$$g_{h\tau\tau} = \frac{gm_\tau \sin \alpha}{2m_W \cos \beta} \frac{1}{1 + \Delta_\tau} \left[1 - \frac{\Delta_\tau}{\tan \alpha \tan \beta} \right] \simeq \frac{gm_\tau \sin \alpha}{2m_W \cos \beta}$$

Partial widths are given by:

$$\Gamma(b) \propto g_{hbb}^2 \quad (\text{modulo kinematics})$$

$$\Gamma(\tau) \propto g_{h\tau\tau}^2$$

Branching ratios:

$$BR = \frac{\Gamma}{\Gamma_{total}}$$

Even if Δ_τ is negligible, $BR(\tau)$ is still affected:
 Δ_b changes $\Gamma(b)$ and thus affects all the BRs by changing the h^0 total width.

Some generalities:

- Correction to α depends on sign of μA_t , while Δ_b depends on sign of $\mu M_{\bar{g}}$.
→ Flip sign of both μ and A_t → α remains the same while Δ_b flips sign.
- Correction to α affects $BR(b)$ and $BR(\tau)$ in the same way.
In contrast, Δ_b affects $\Gamma(b)$ but not $\Gamma(\tau)$; however it affects all the BRs because it changes the total width.

Three “benchmark” scenarios:

1. Minimal stop mixing, $X_t = 0$
Parameters: $M_S = 1.5$ TeV, $M_{\tilde{g}} = 1$ TeV, $\mu = -200$ GeV, $M_2 = 200$ GeV
2. Maximal mixing, $X_t = \sqrt{6}M_S$ – Maximises h^0 mass
Parameters: $M_S = 1$ TeV, $M_{\tilde{g}} = 1$ TeV, $\mu = -200$ GeV, $M_2 = 200$ GeV
3. Large μ , $A_t \rightarrow$ large contributions to both α and Δ_b
Parameters: $M_S = 1$ TeV, $M_{\tilde{g}} = 500$ GeV, $M_2 = 200$ GeV, $\mu = -A_t = \pm 1.2$ TeV

We plot contours of:

$$\frac{|BR(MSSM) - BR(SM)|}{BR(SM)}$$

over the $M_A - \tan\beta$ plane.

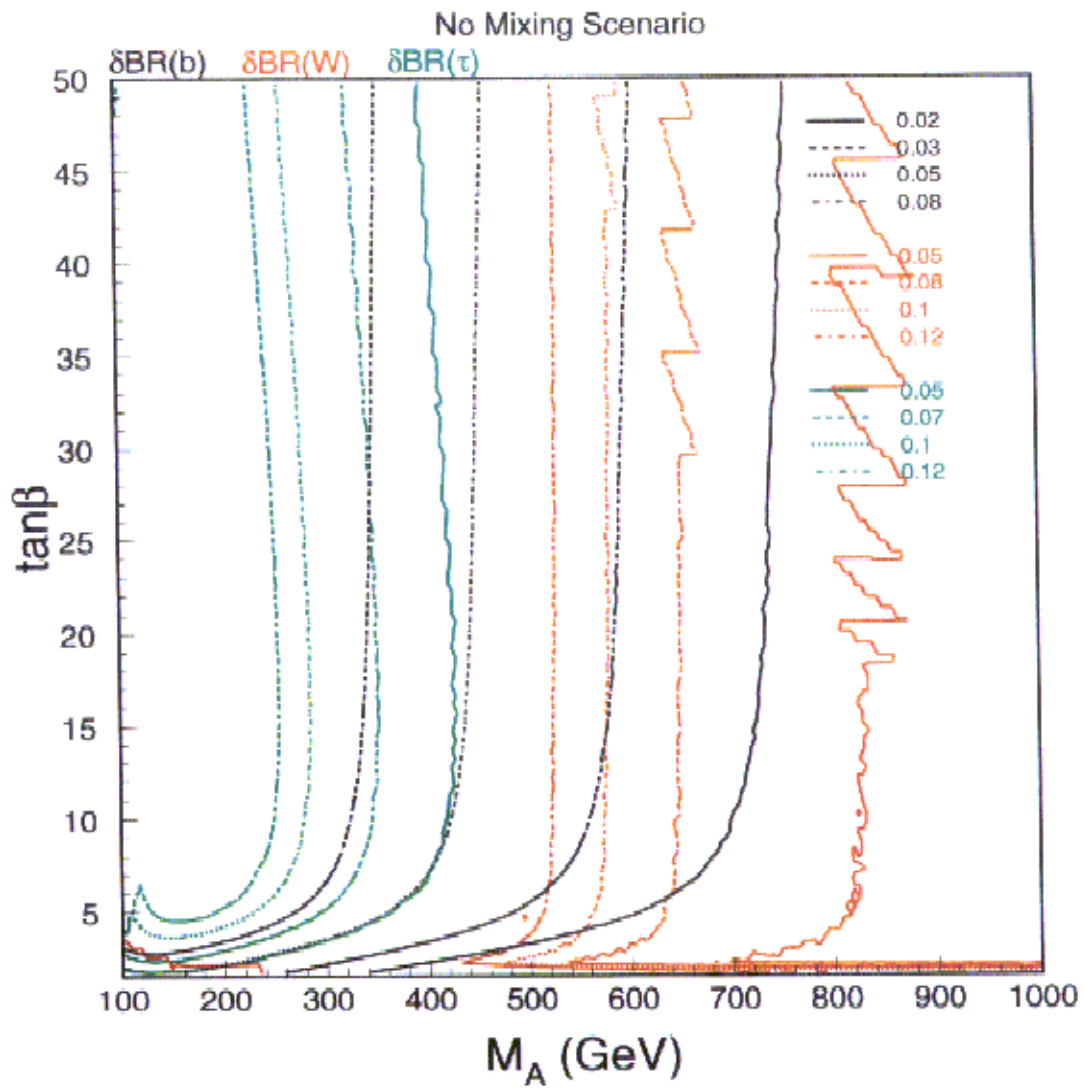
We use HDECAY with Δ_b corrections inside.

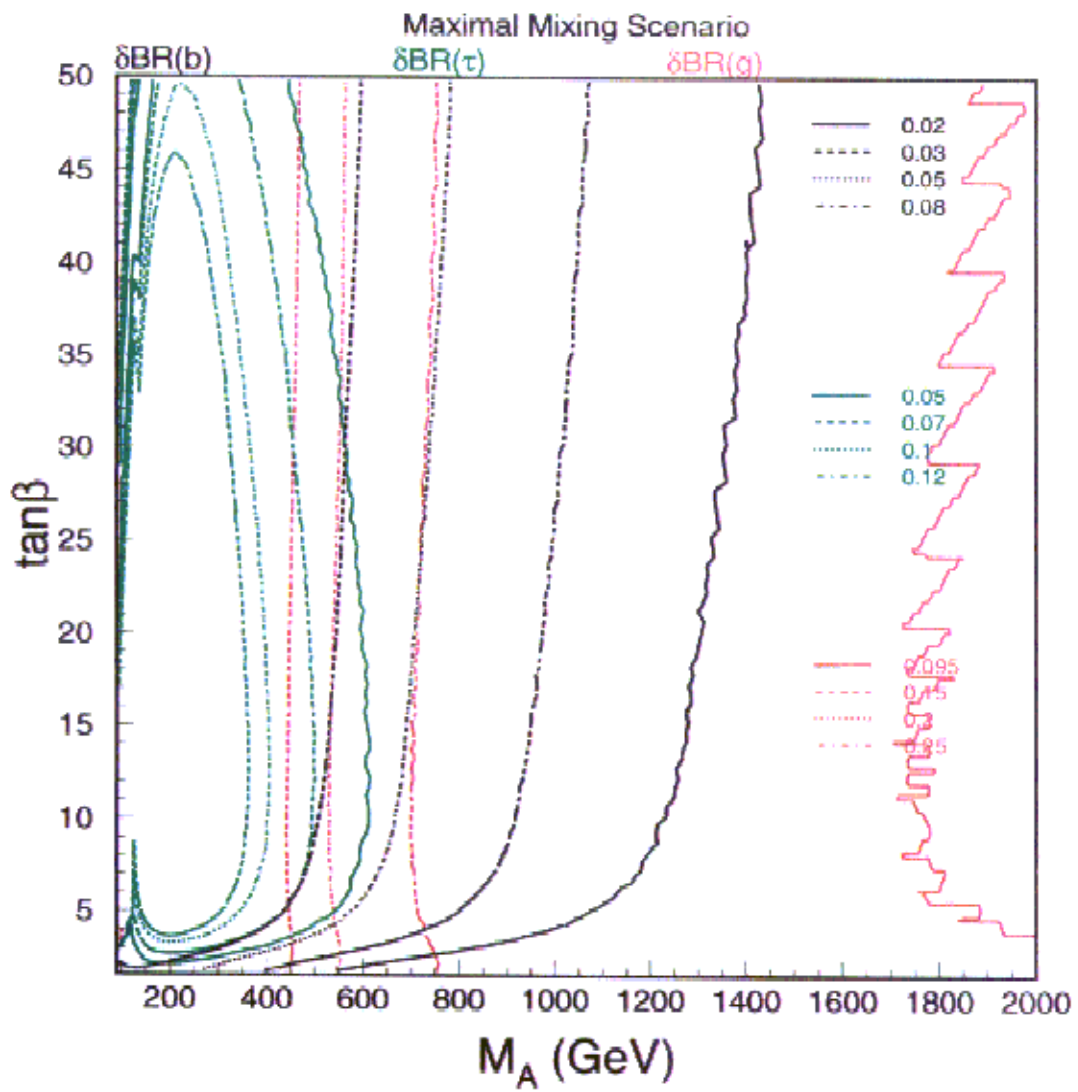
Plots made by Steve Mrenna.

Expected LC precision for BRs:

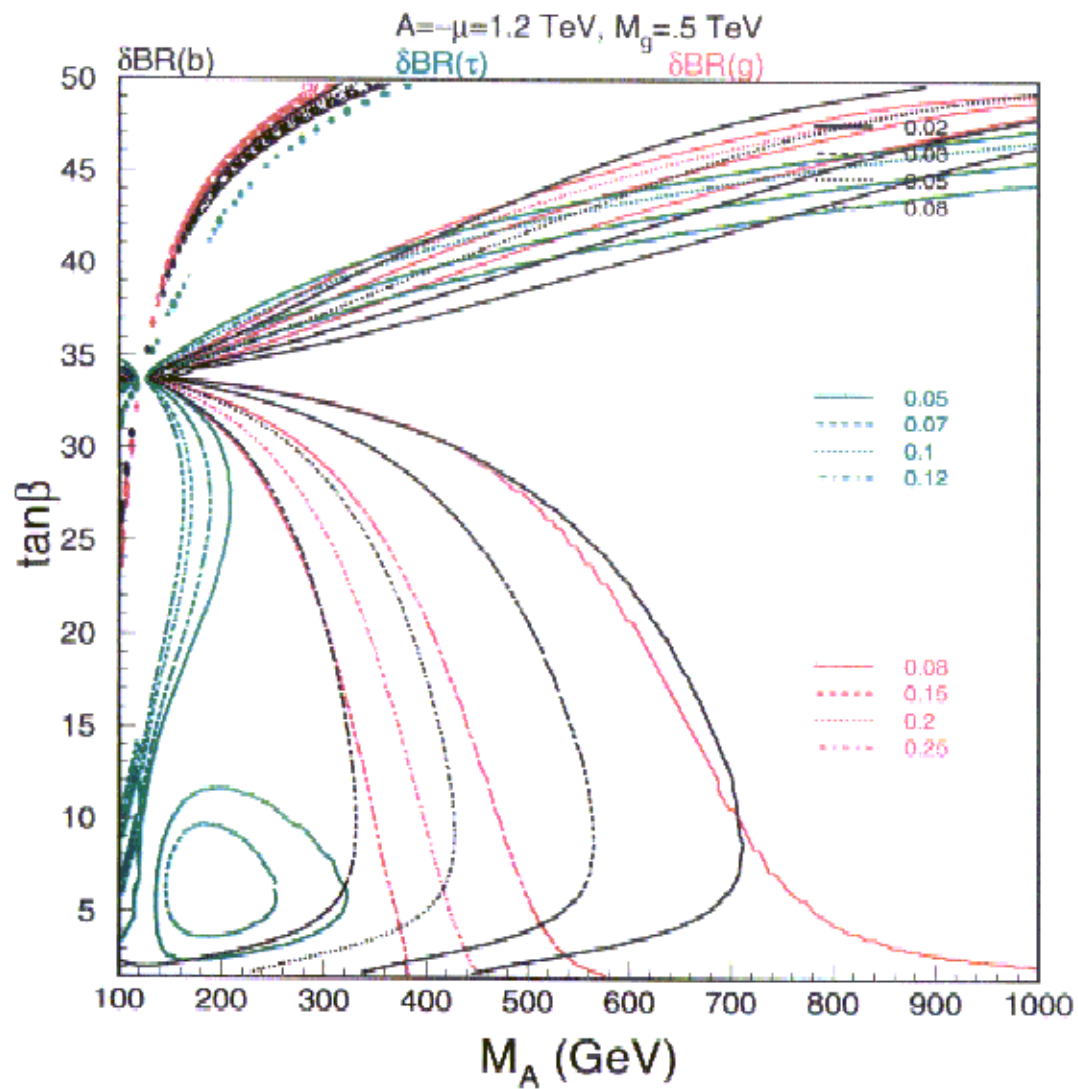
BR	(a)	(b)
bb	3%	2.4%
WW*	8%	5.4%
$\tau\tau$	7%	8.3 - 13.5%
cc	15%	—
gg	8%	5.5%
$\gamma\gamma$	22%	—

- (a) Rick Van Kooten at March LC meeting at LBL for SM H^0 at 120 GeV with $\sqrt{s} = 500$ GeV and 200 fb^{-1}
(b) Marco Battaglia, LCWS99, Sitges: for SM H^0 at 120 GeV with $\sqrt{s} = 350$ or 500 GeV and 500 fb^{-1}

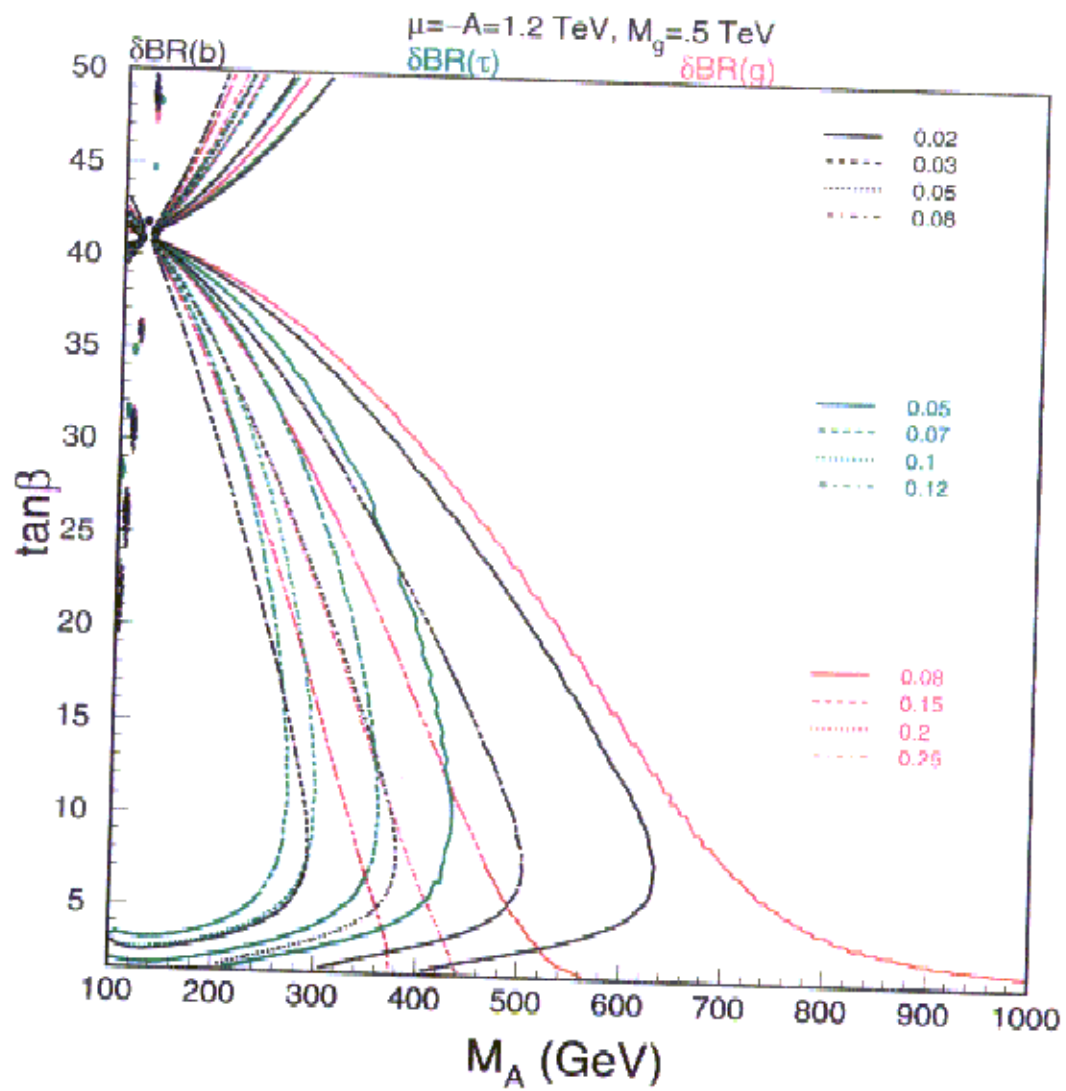




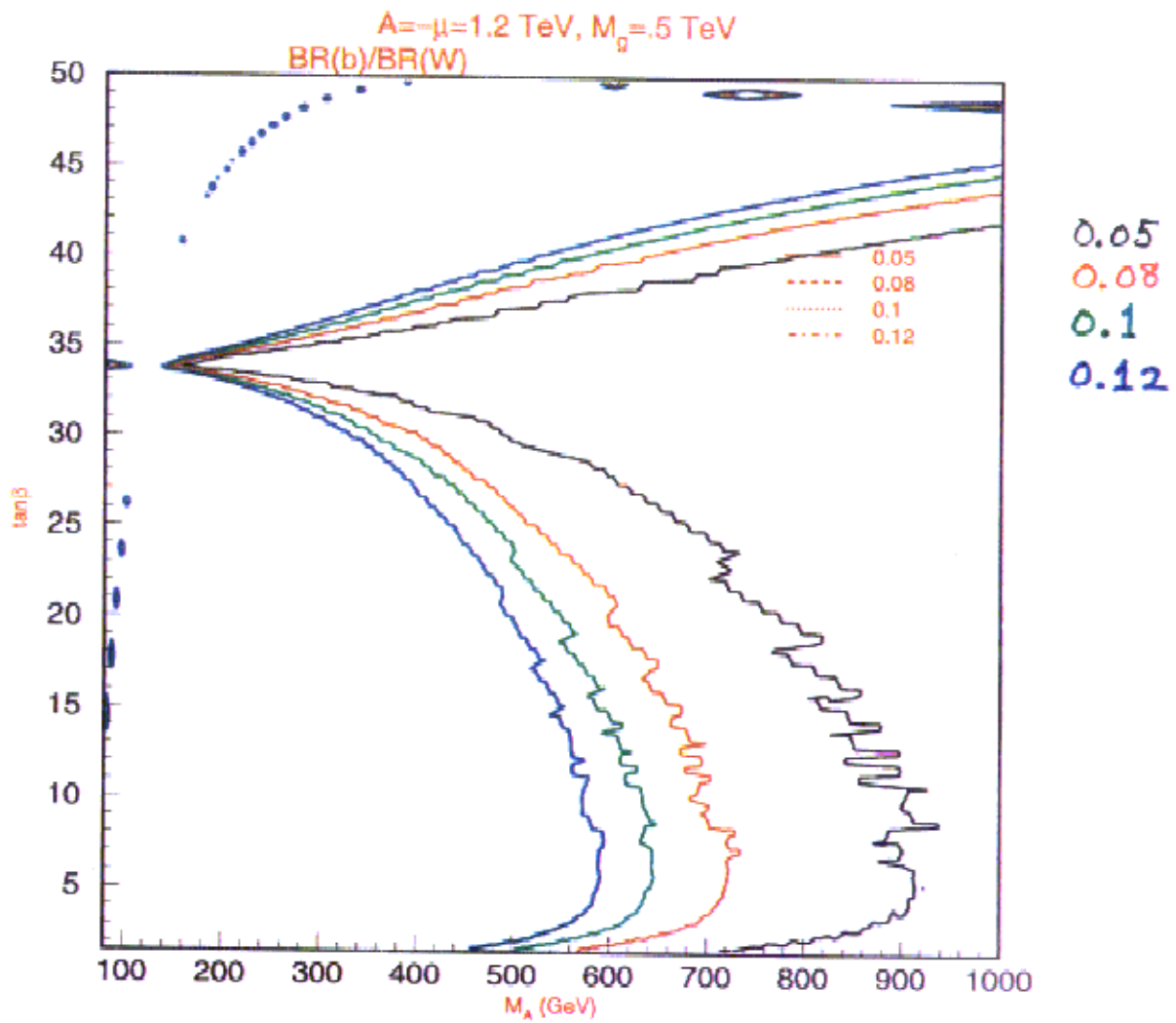
Large μ and A_t scenario with $\mu < 0$ (y_b increased)



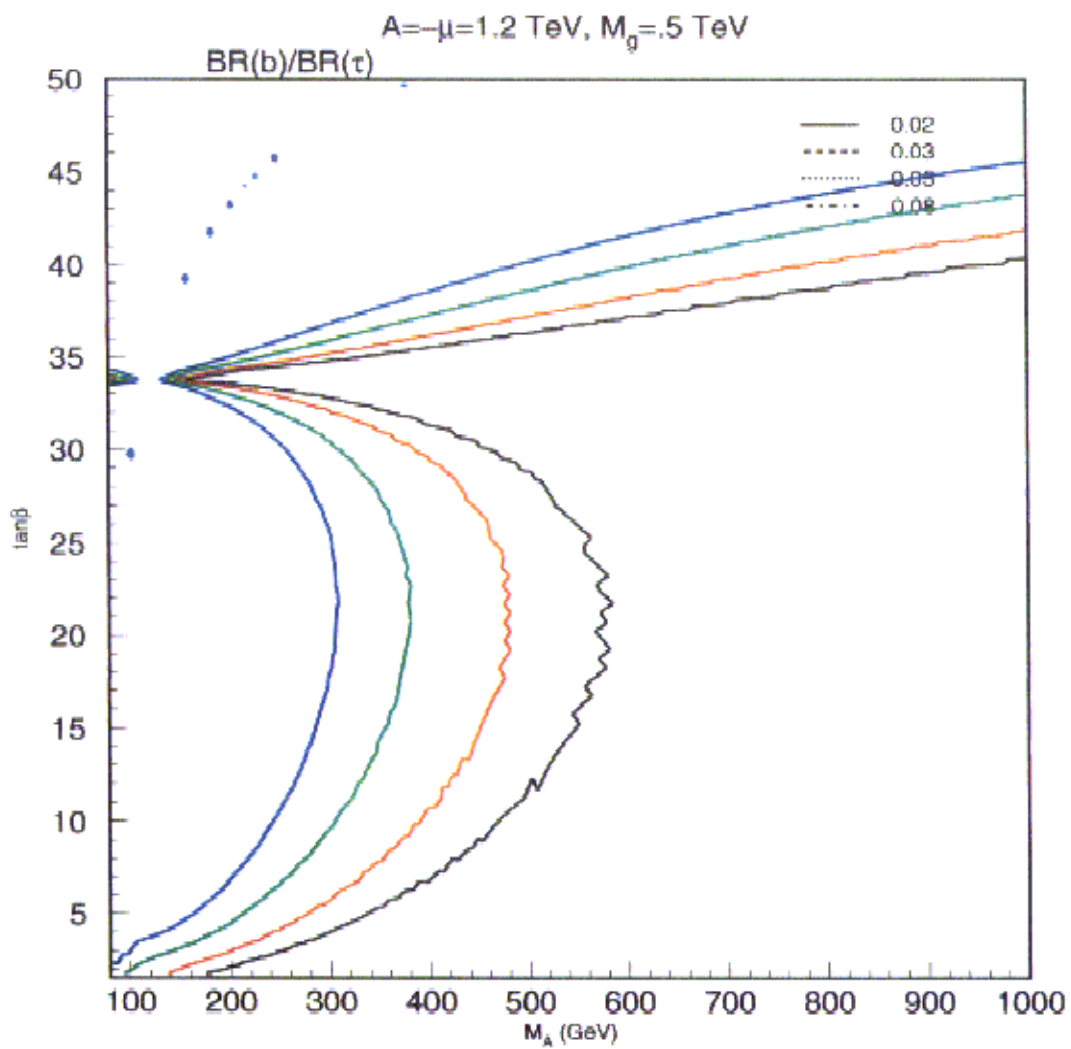
Large μ and A_t scenario with $\mu > 0$ (y_b decreased)



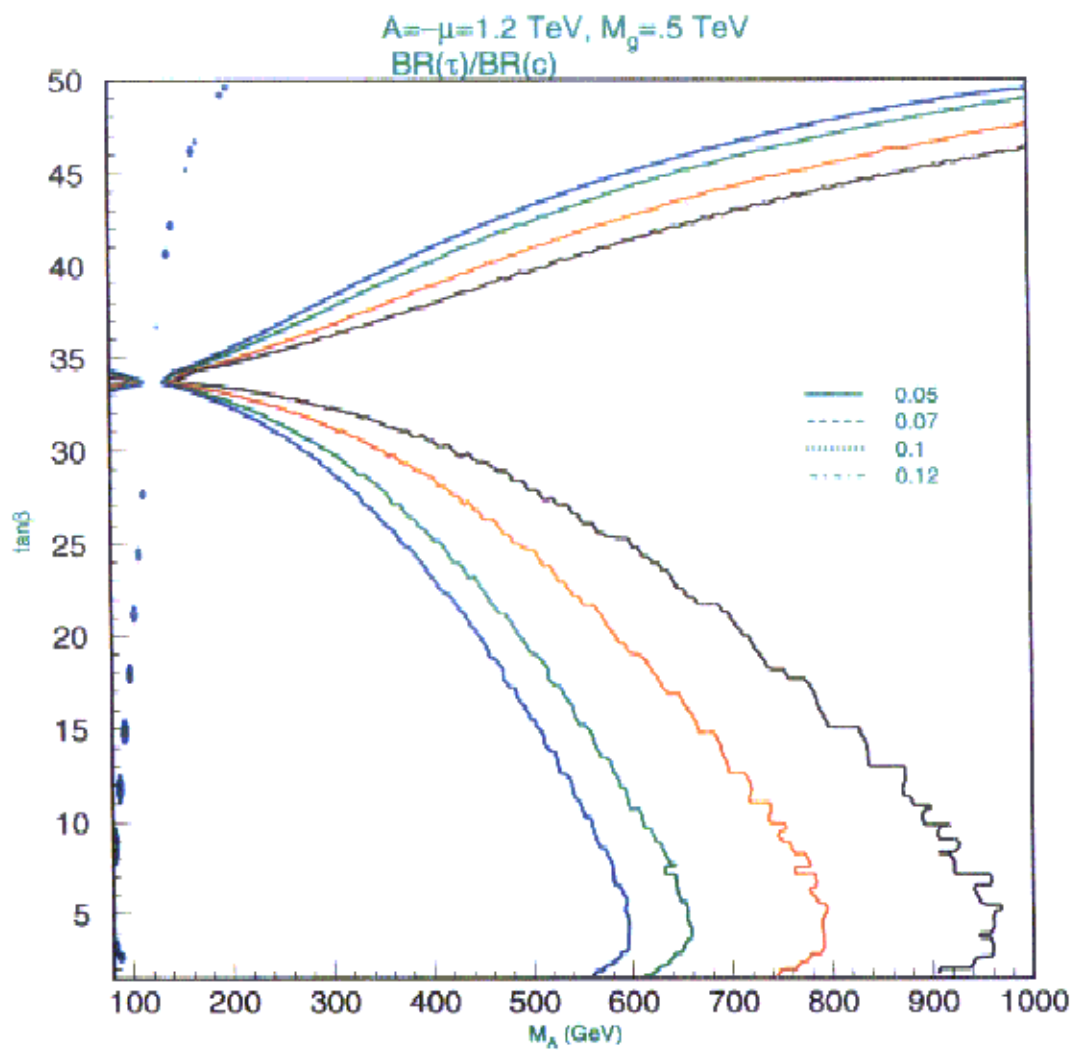
Measuring ratios of BRs can increase the sensitivity:



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Can extract Δ_b in some regions of parameter space:

$$\frac{g_{hbb}}{g_{hbb}^{SM}} = -\frac{\sin \alpha}{\cos \beta} \frac{1}{1 + \Delta_b} \left[1 - \frac{\Delta_b}{\tan \alpha \tan \beta} \right]$$

$$\frac{g_{h\tau\tau}}{g_{h\tau\tau}^{SM}} = -\frac{\sin \alpha}{\cos \beta}$$

so,

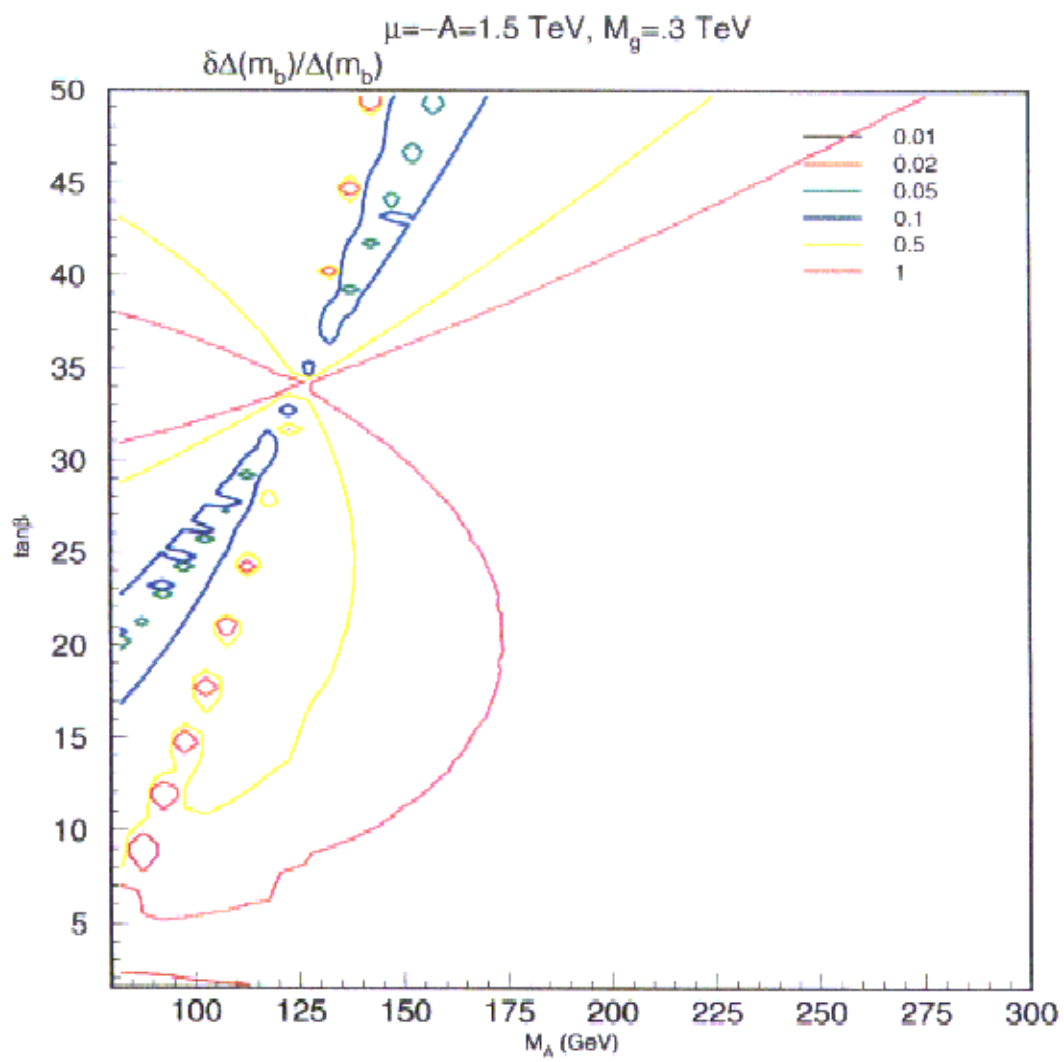
$$\frac{g_{hbb}/g_{hbb}^{SM}}{g_{h\tau\tau}/g_{h\tau\tau}^{SM}} = \frac{1}{1 + \Delta_b} \left[1 - \frac{\Delta_b}{\tan \alpha \tan \beta} \right]$$

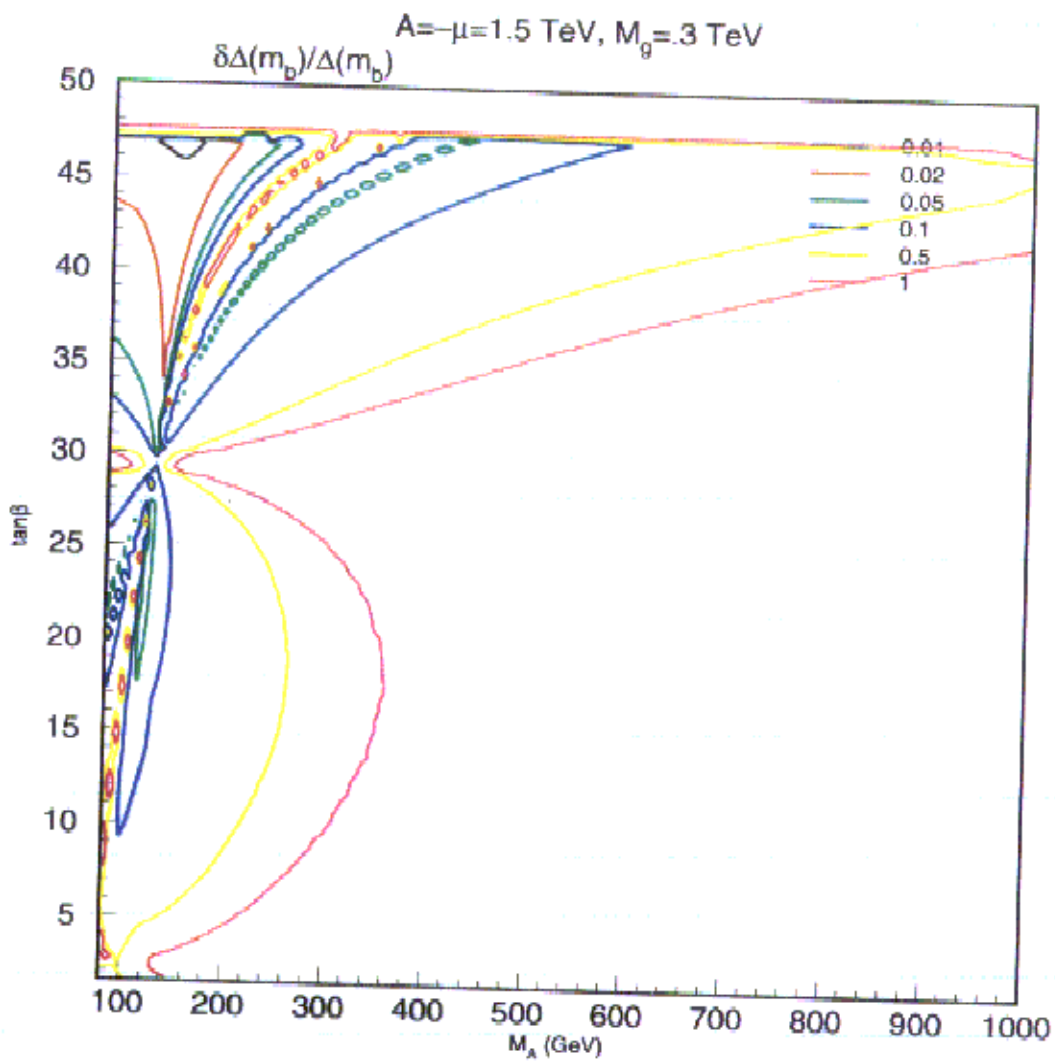
Get this from ratios of branching ratios:

$$\begin{aligned} \frac{BR(b)/BR(b)_{SM}}{BR(\tau)/BR(\tau)_{SM}} &= \left(\frac{g_{hbb}/g_{hbb}^{SM}}{g_{h\tau\tau}/g_{h\tau\tau}^{SM}} \right)^2 \\ &= \left(\frac{1}{1 + \Delta_b} \left[1 - \frac{\Delta_b}{\tan \alpha \tan \beta} \right] \right)^2 \end{aligned}$$

Still need $\tan \alpha \tan \beta$: extract it from the ratio of the $\tau\tau$ and cc branching ratios:

$$\begin{aligned} \frac{BR(\tau)/BR(\tau)_{SM}}{BR(c)/BR(c)_{SM}} &= \left(\frac{g_{h\tau\tau}/g_{h\tau\tau}^{SM}}{g_{hcc}/g_{hcc}^{SM}} \right)^2 = \left(\frac{\sin \alpha / \cos \beta}{\cos \alpha / \sin \beta} \right)^2 \\ &= (\tan \alpha \tan \beta)^2 \end{aligned}$$





Conclusions

- Different MSSM parameter space regions
→ different reach in distinguishing SM from MSSM Higgs.
 $\sqrt{s} = 500 \text{ GeV}$, 200 fb^{-1} (Van Kooten):
 - No mixing scenario is the usual benchmark: Can distinguish MSSM Higgs from SM Higgs up to $M_A \sim 600 - 700 \text{ GeV}$ using sensitivity to deviations from SM BR(W) (500 - 600 GeV with BR(b))
 - But you can do much better in some regions of parameter space: Can distinguish MSSM Higgs from SM Higgs up to $M_A \sim 900 - 1000 \text{ GeV}$ in the maximal mixing scenario using BR(b)
 - ... and in others you can do much worse: In the large μ and A_t scenario it can be very hard to distinguish MSSM Higgs from SM Higgs at moderate $\tan\beta$, even for M_A as light as 200 GeV.

- In some regions of parameter space, can extract Δ_b from the $BR(b)/BR(\tau)$ and $BR(\tau)/BR(c)$ measurements.
More work being done.
- Important to measure all the BRs that we can as accurately as possible:
 - Different BRs more sensitive in different regions of parameter space
 - Ratios of BRs sometimes more sensitive than individual BRs
- Ultimately use Higgs sector (and SUSY sector) to do precision measurements of SUSY parameters.